

$\mu(l)$ rheology for granular flows with *Gerris*
applications for
avalanches and column collapse



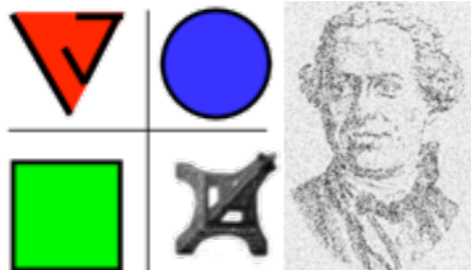
Pierre-Yves Lagrée*,

Lydie Staron*, Stéphane Popinet *o, (Daniel Lhuillier *, Christophe Josserand *)

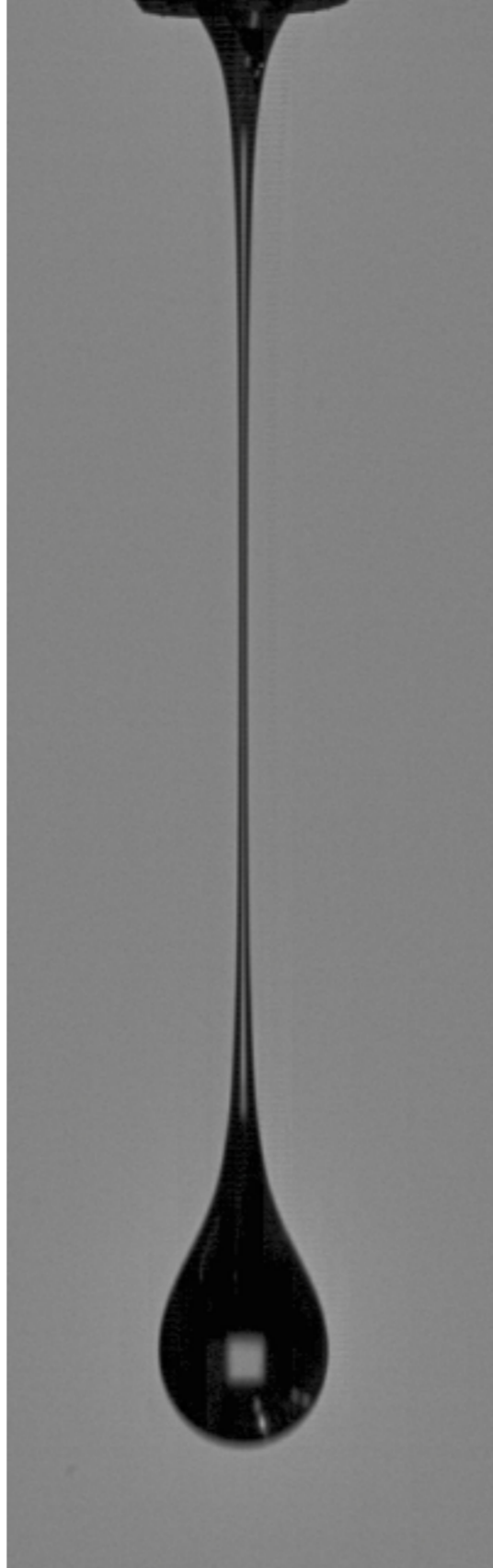
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o National Institute of Water and Atmospheric Research, PO Box 14-901 Kilbirnie, Wellington, New Zealand

Saint Gobain 05/04/12



Institut Jean Le Rond d'Alembert





outline

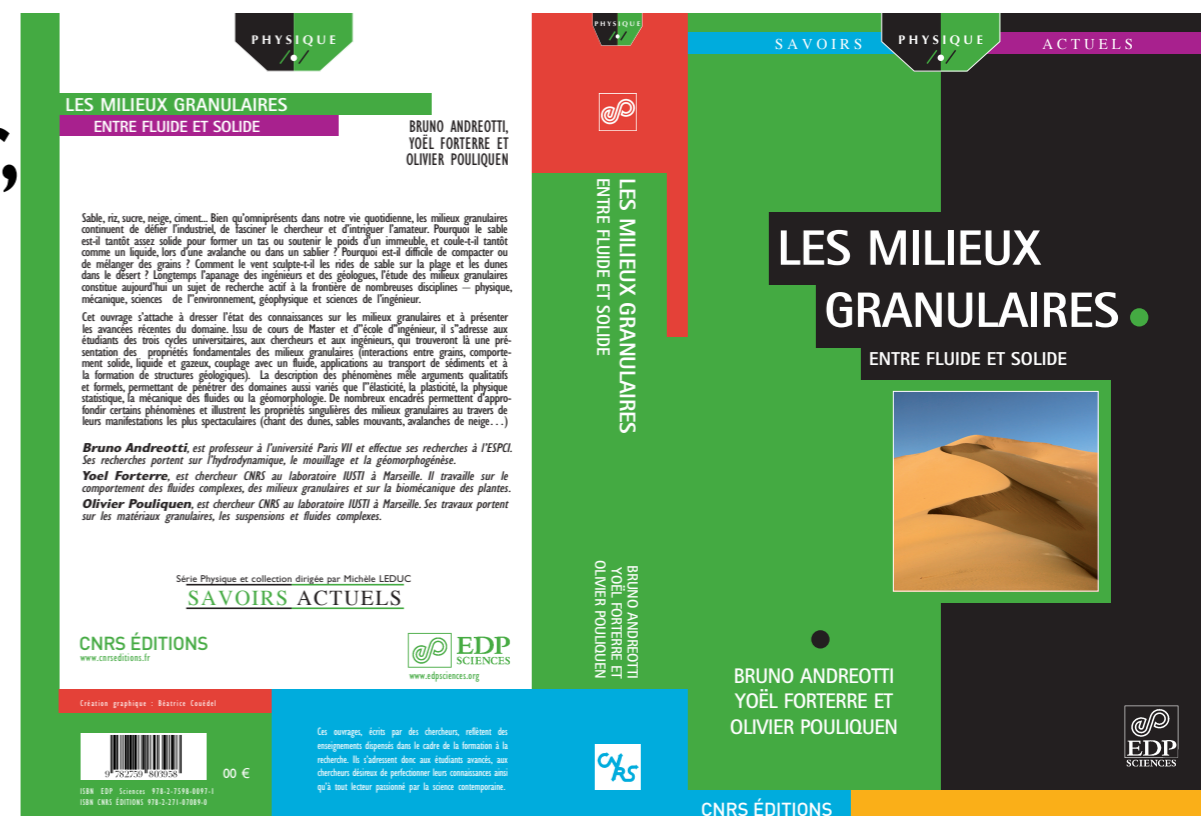
- what is a granular fluid? some images
- the $\mu(I)$ friction law obtained from experiments and discrete simulation
- the viscosity associated to the $\mu(I)$ friction law
- the Saint Venant Savage Hutter Hyperbolic model
- implementing the $\mu(I)$ friction law in Navier Stokes
- Examples of flows: focusing on the granular column collapse and the Hour Glass



- What is a granular media?
- size $> 100\mu\text{m}$
- grains of sand, small rocks, glass beads, animal feed pellet, medicines, cereals, wheat, sugar, rice...
- 50 % of the traded products



FIG. 1.2 – Les milieux granulaires forment une famille extrêmement vaste.





spoil tip (boney pile, gob pile, bing or pit heap), «terril» in french

pyl

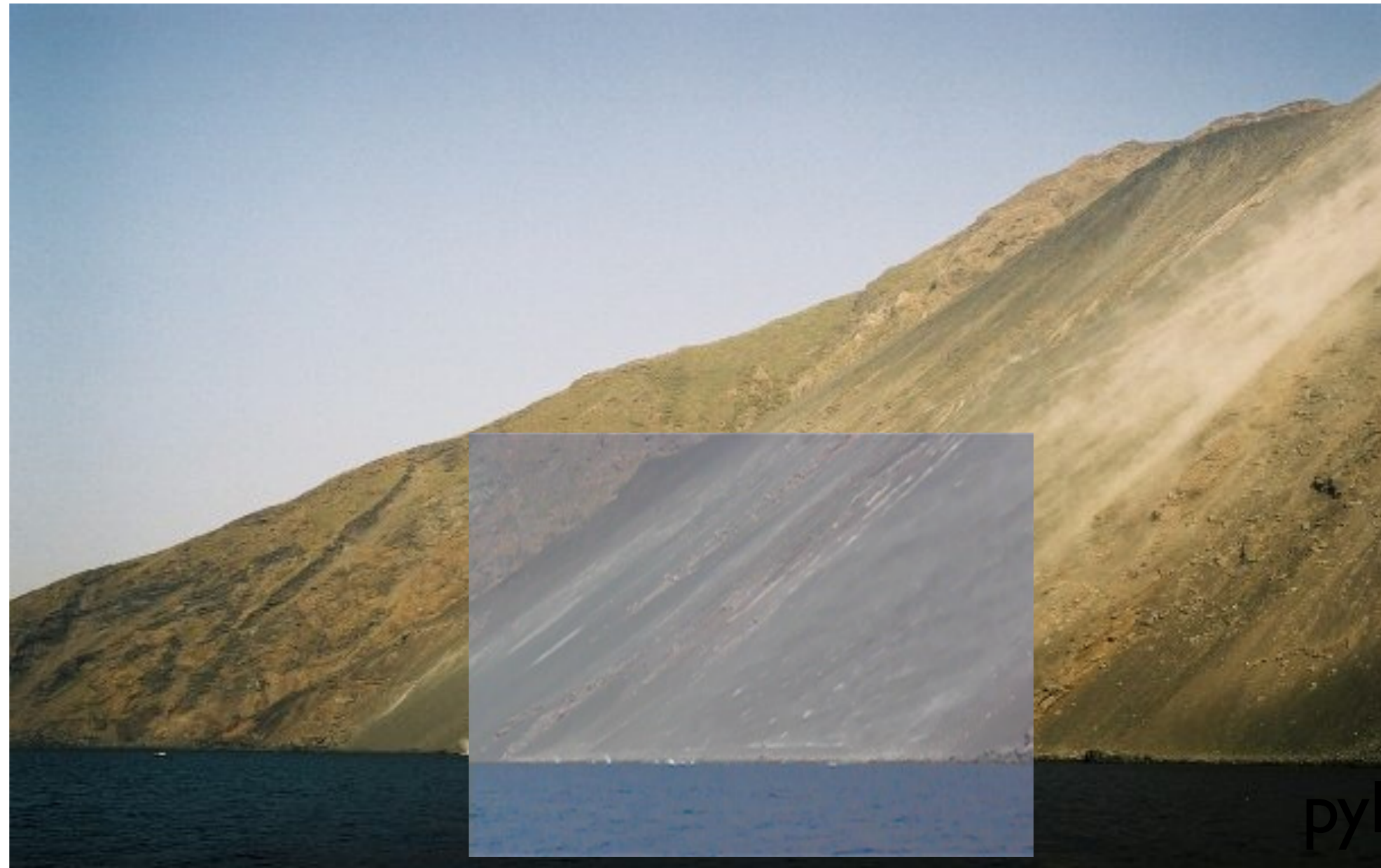
pyl



ру





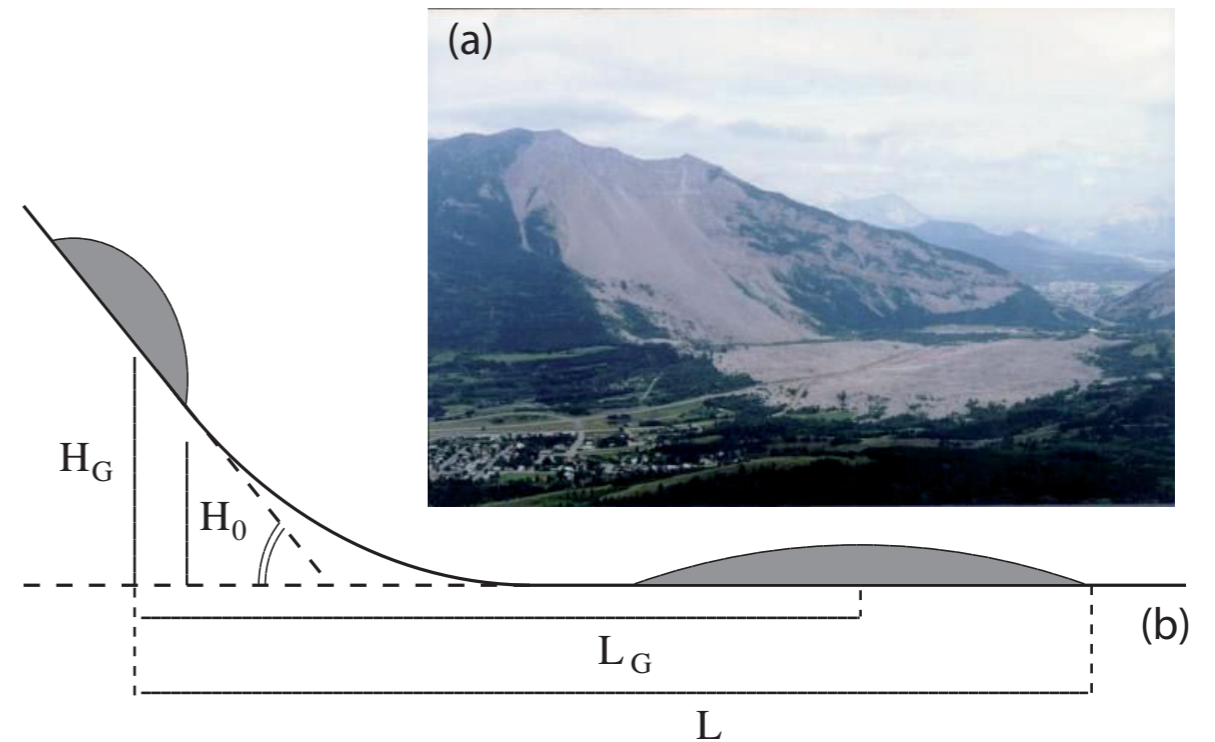




- the Saint Venant Savage Hutter Hyperbolic model



Fig. 20. Frank slide.



$$H \times L = 1000\text{m} \times 2500\text{m}$$

Staron

Environmental Modelling & Software xx (2006) 1e18

www.elsevier.com/locate/envsoft

The effect of the earth pressure coefficients on the runout of granular material

Marina Pirulli ^{a,*}, Marie-Odile Bristeau ^b, Anne Mangeney ^c, Claudio Scavia

1903, Alberta Canada

90 morts



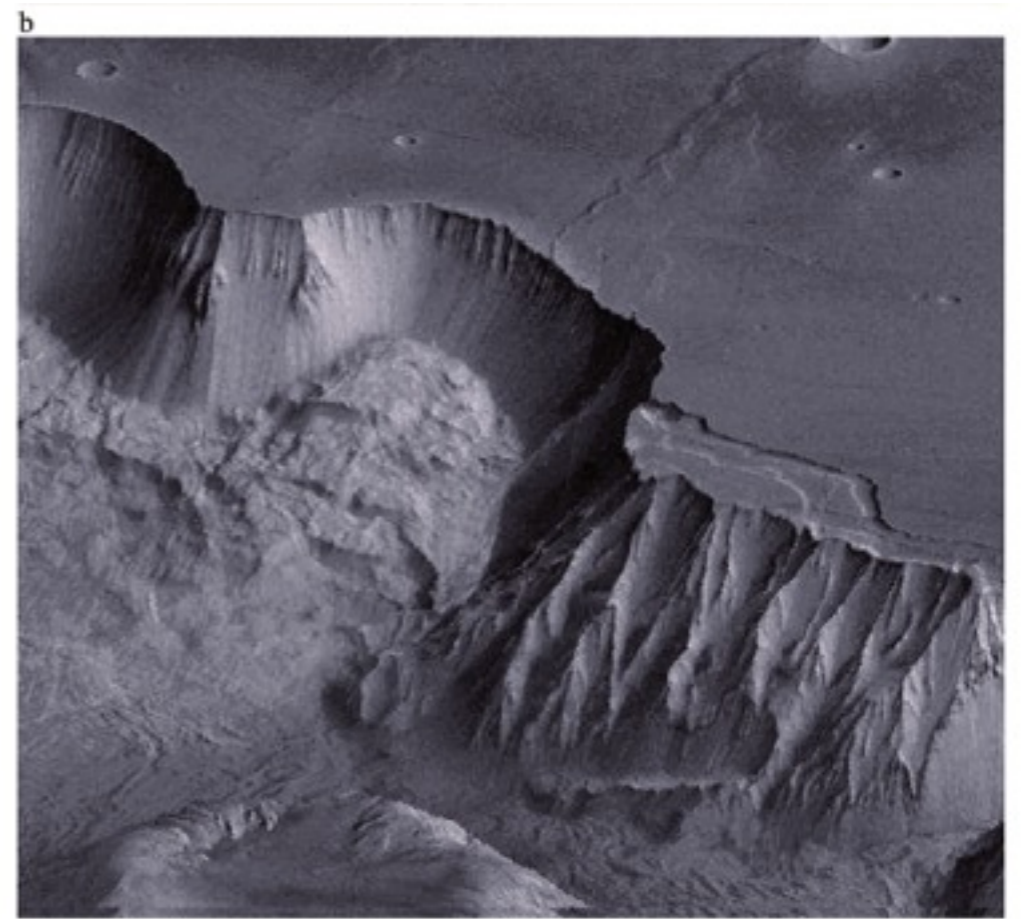
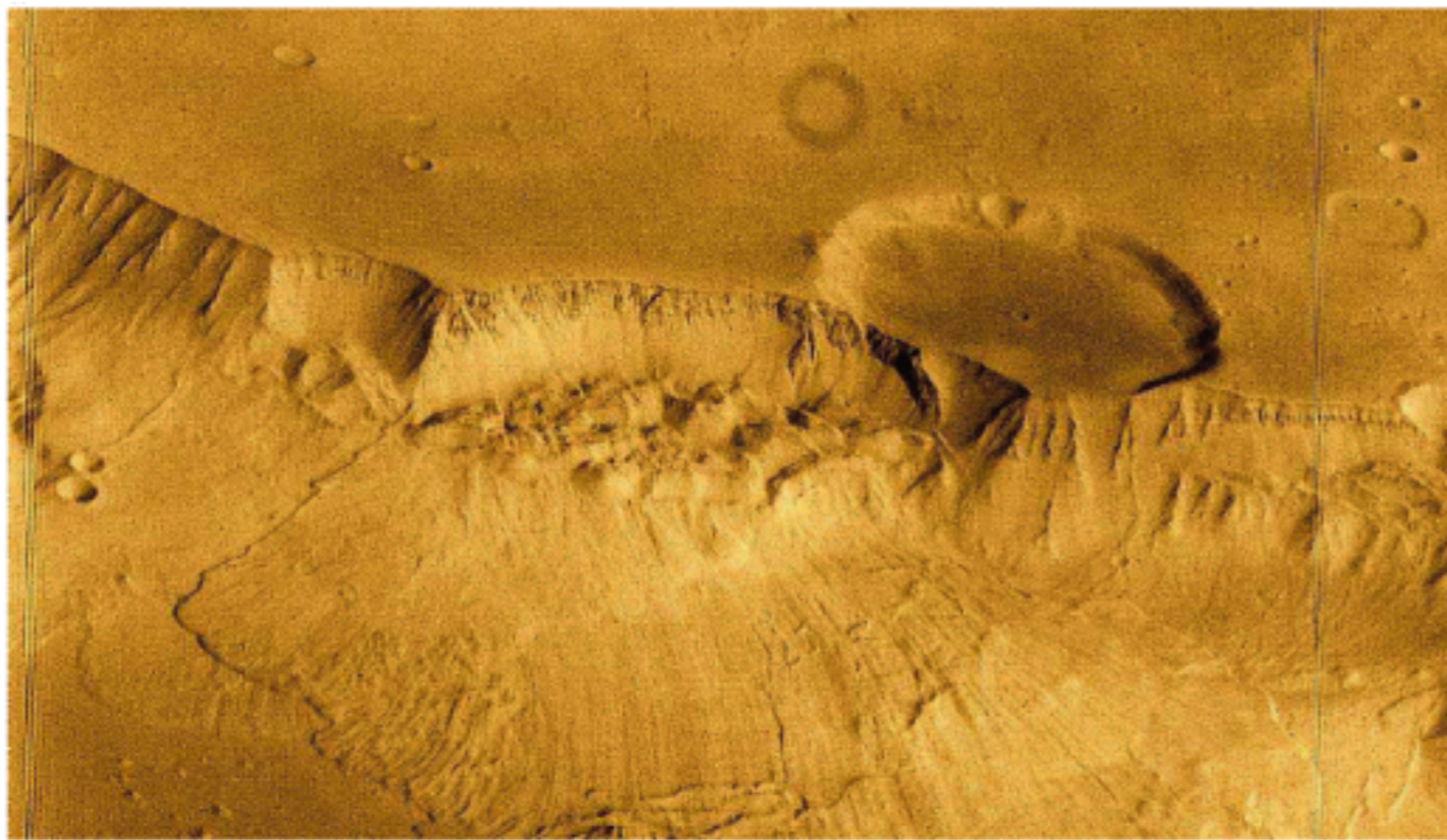
<http://www.pbase.com/image/63044602>

2006 Gary Hebert



Lofoten Norway





<http://books.google.fr/books?id=HY6Z5od4-E4C&pg=PA49&dq=granular>

[+flow&hl=fr&ei=lamtTaa_NYyVOoToldcL&sa=X&oi=book_result&ct=result&resnum=10&ved=0CFkQ6AEwCTgK#v=onepage&q&f=true](http://books.google.fr/books?id=HY6Z5od4-E4C&pg=PA49&dq=granular+flow&hl=fr&ei=lamtTaa_NYyVOoToldcL&sa=X&oi=book_result&ct=result&resnum=10&ved=0CFkQ6AEwCTgK#v=onepage&q&f=true)



http://www.cieletespace.fr/image-du-jour/5126_la-saison-des-avalanches-sur-mars

Granular Column Collapse

A model for avalanches

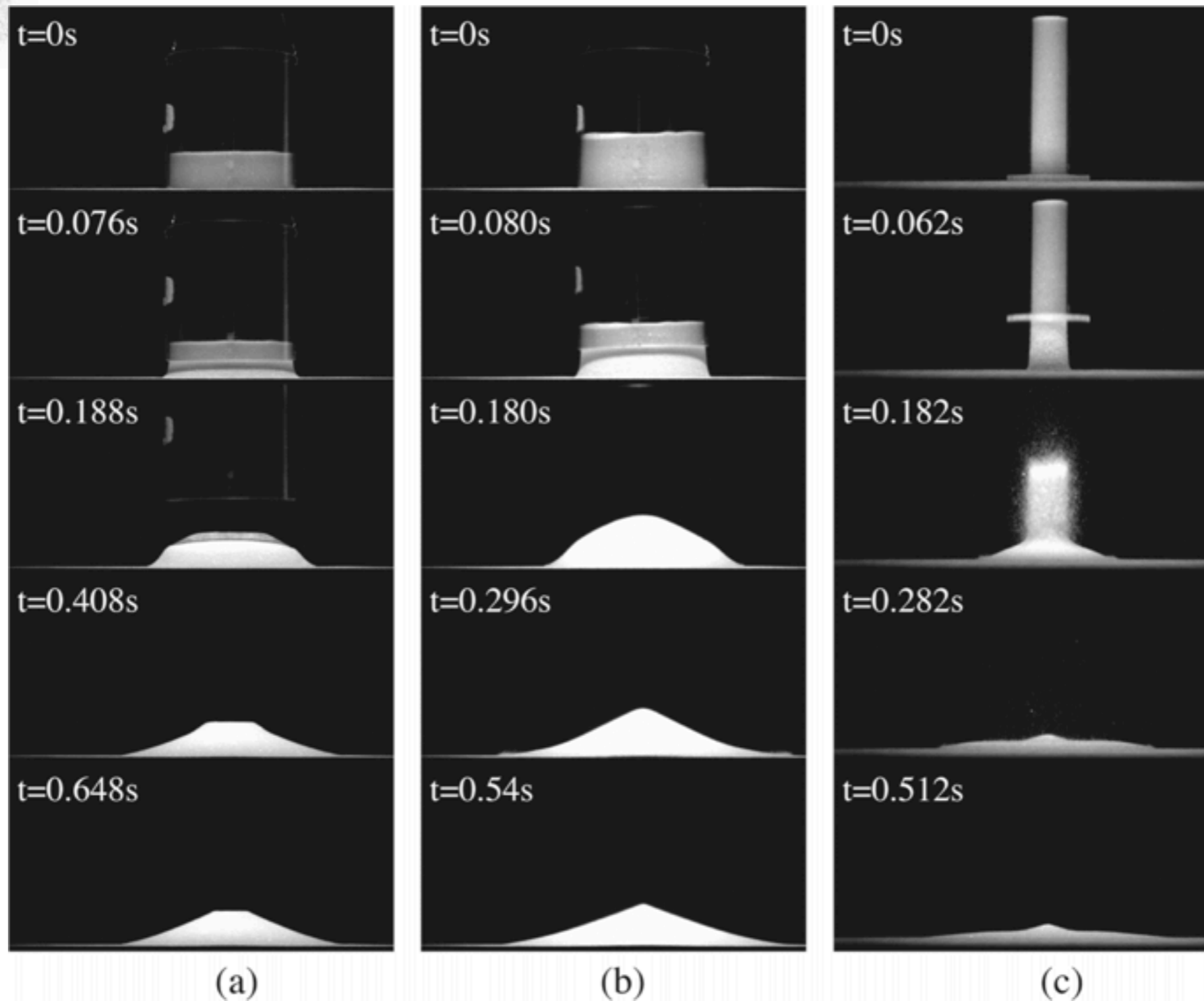


The sand pit problem: quickly remove the bucket of sand



Granular Column Collapse

E. Lajeunesse A. Mangeney-Castelnau and J. P. Vilotte PoF 2004



(a)

(b)

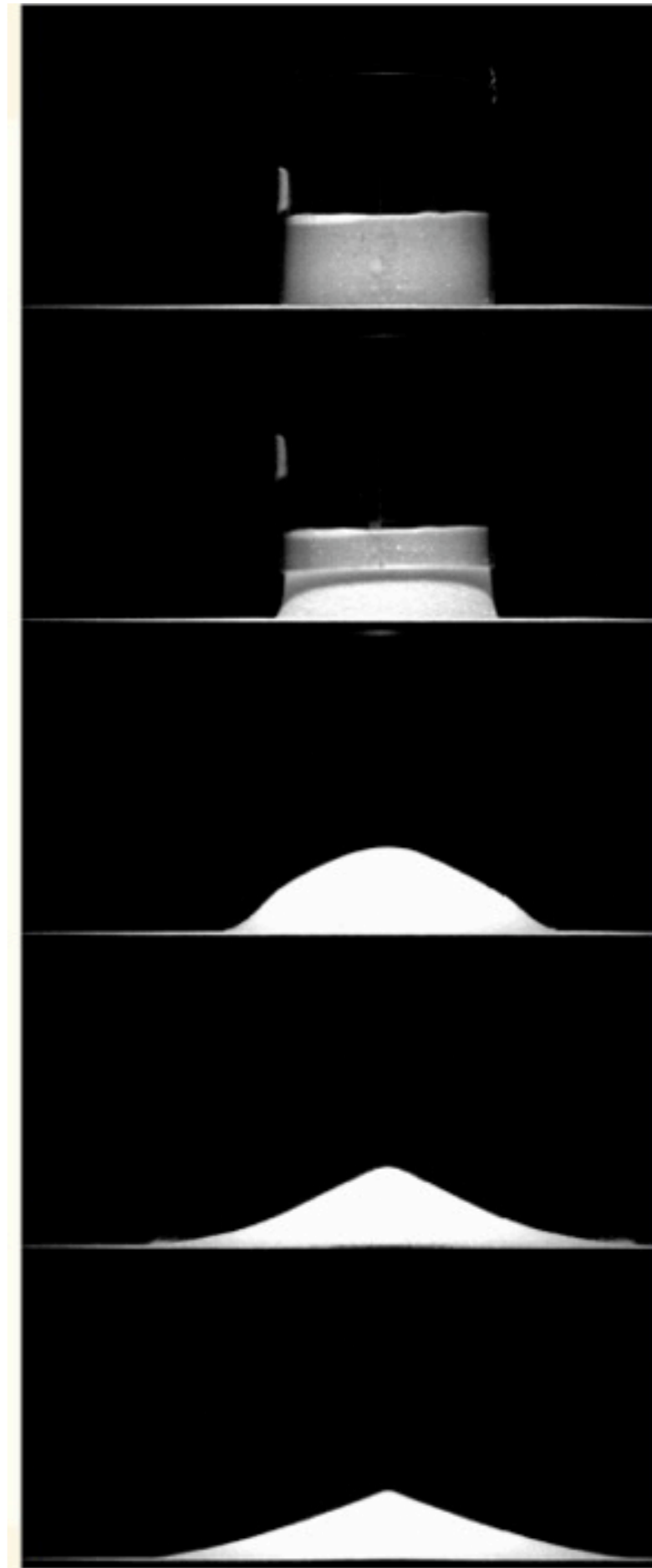
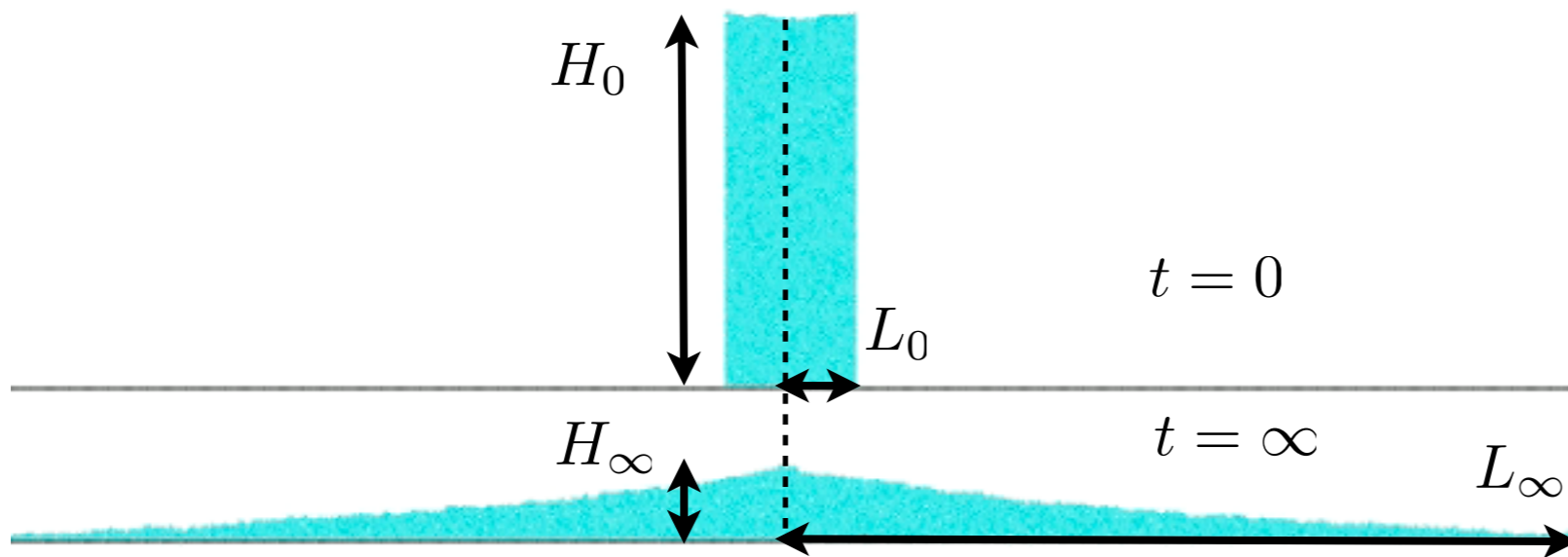
(c)

The sand pit problem: quickly remove the bucket of sand



Granular Column Collapse

aspect ratio $a = H_0/R_0 = H_0/L_0$



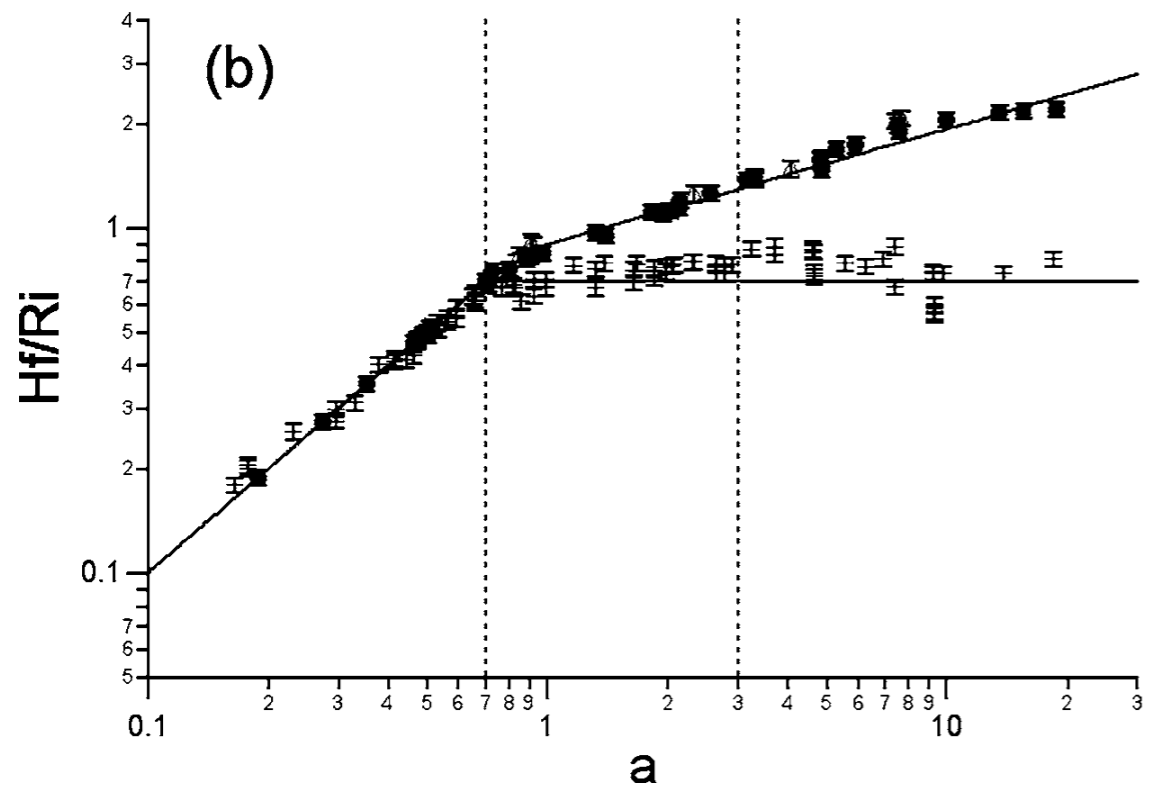
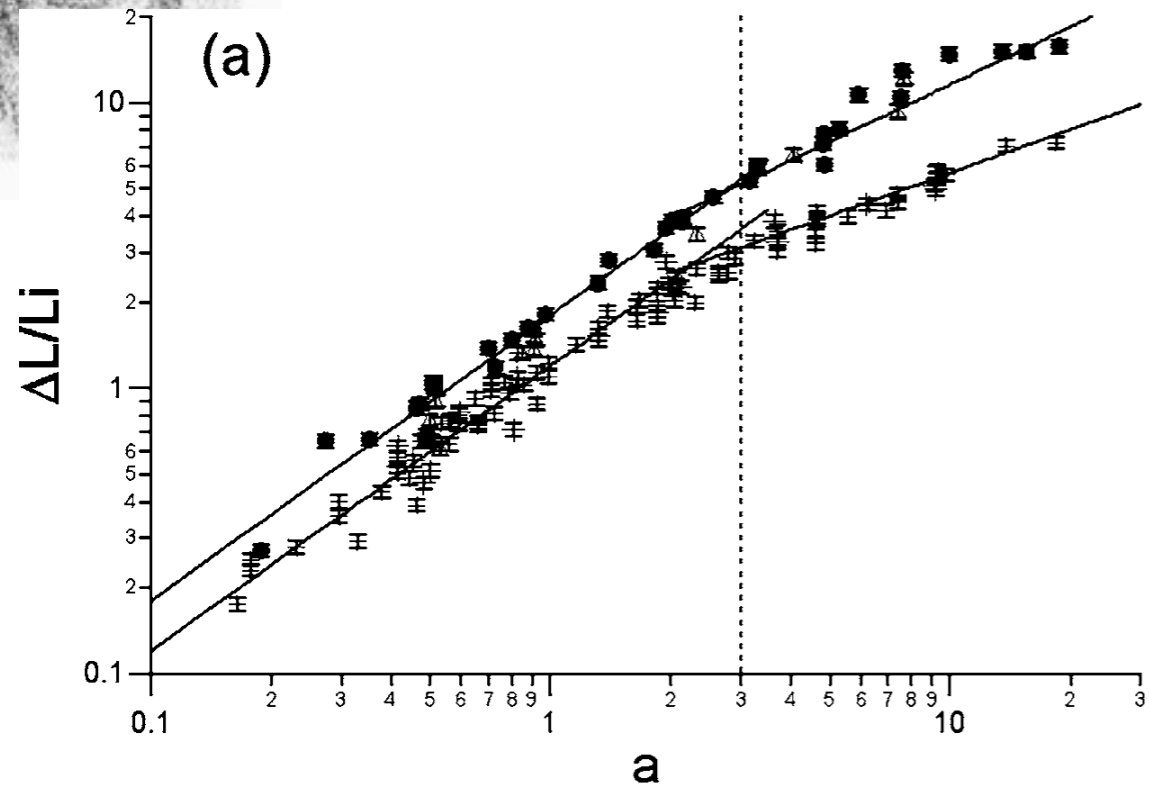
The sand pit problem: quickly remove the bucket of sand

Granular Column Collapse



A possible experimental set up is a container filled by sand (left), the aspect ratio (height/length) is a . At initial time, the gate is opened quickly. After the avalanche, the grains stop, the final configuration is at rest (right). We compare results from Discrete Contact Method Simulations (simulation of the displacement of each grain) to a continuum Navier Stokes simulation with the $\mu(I)$ rheology *Gerris*.

The sand pit problem: quickly remove the bucket of sand



• In the axisymmetric geometry

$$\frac{H_f}{L_i} = \begin{cases} a & a \lesssim 0.74, \\ 0.74 & a \gtrsim 0.74, \end{cases}$$

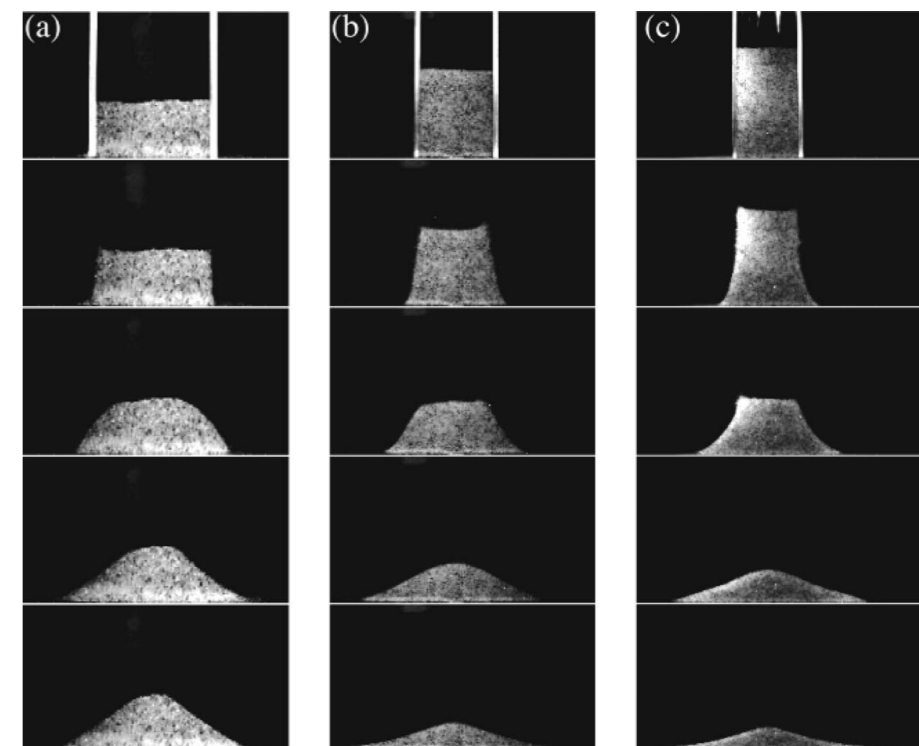
$$\frac{\Delta L}{L_i} \propto \begin{cases} a & a \lesssim 3, \\ a^{1/2} & a \gtrsim 3. \end{cases}$$

• In the rectangular channel:

$$\frac{H_f}{L_i} \propto \begin{cases} a & a \lesssim 0.7, \\ a^{1/3} & a \gtrsim 0.7, \end{cases}$$

$$\frac{\Delta L}{L_i} \propto \begin{cases} a & a \lesssim 3, \\ a^{2/3} & a \gtrsim 3. \end{cases}$$

FIG. 6. Scaled runout $\Delta L/L_i$ (a) and scaled deposit height H_f/L_i (b) as functions of a . Circles and triangles correspond to experiments performed in the 2D channel working respectively with glass beads of diameter $d = 1.15$ mm or $d = 3$ mm. Crosses correspond to the data set of axisymmetric collapses from Lajeunesse *et al.* (Ref. 10).



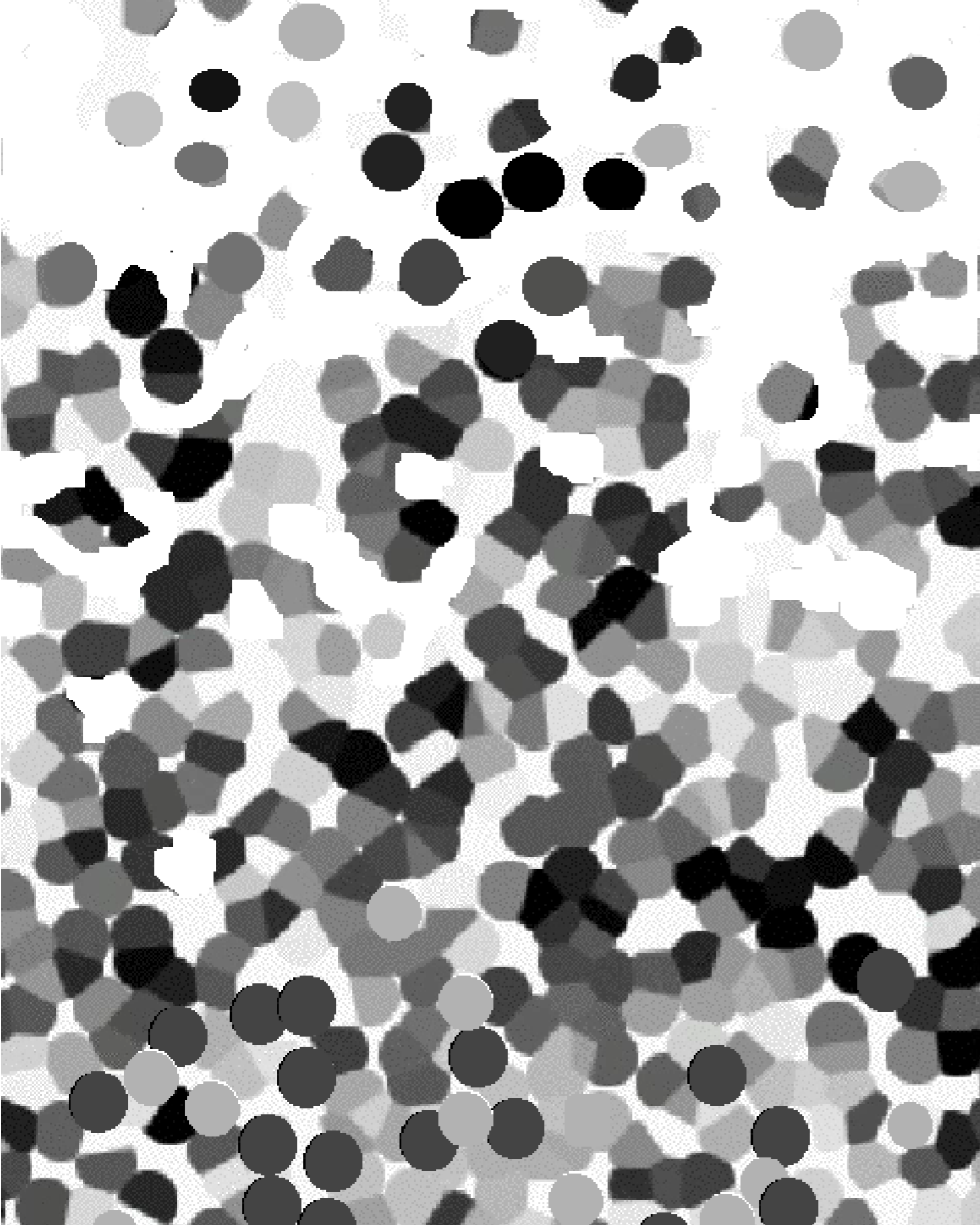
- Silo, hopper, hourglass





outline

- what is a granular fluid? some images
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grains

gaz

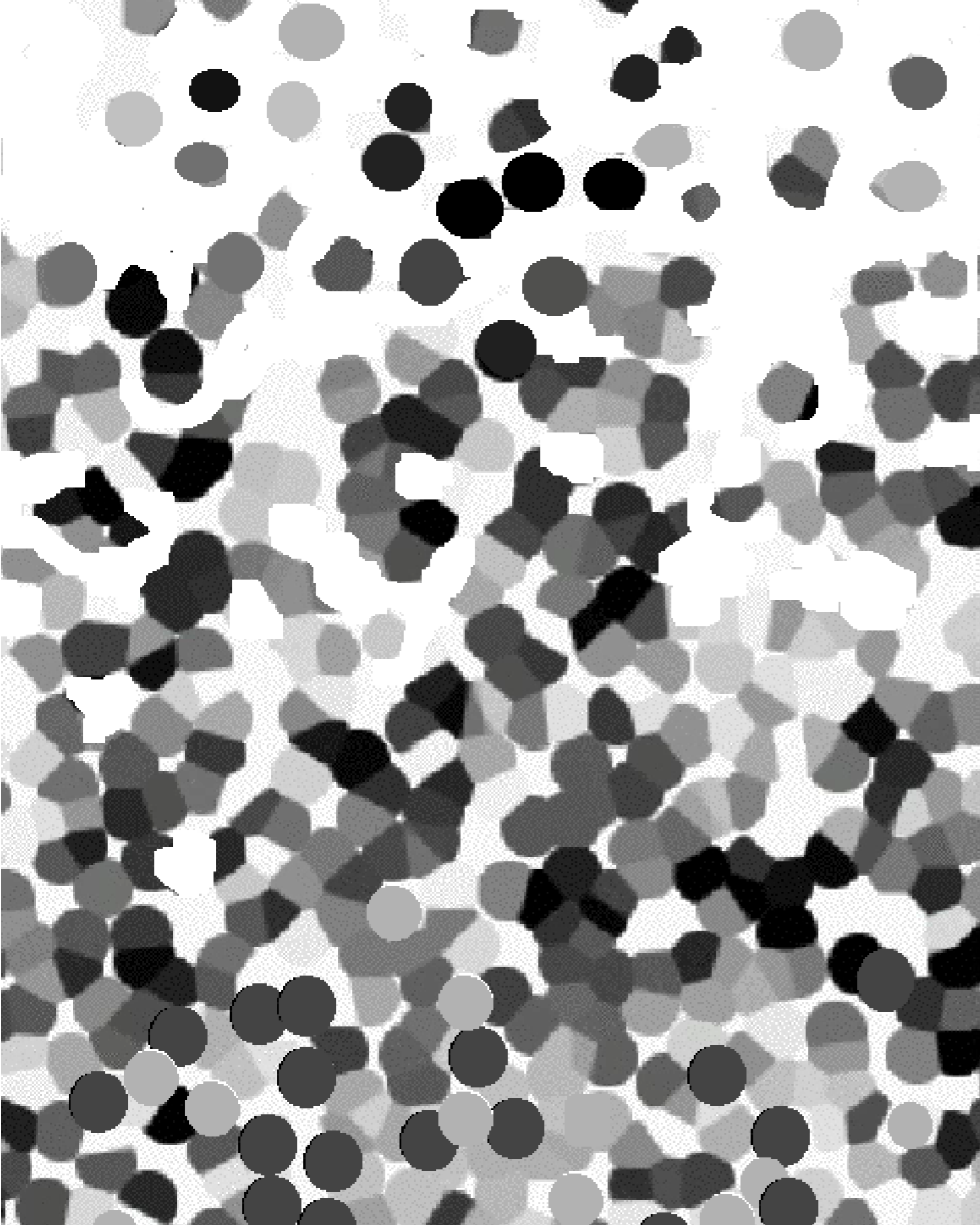
impacts:
suspension

granular media
contacts

liquid

like a solid:

from the grains to the fluid



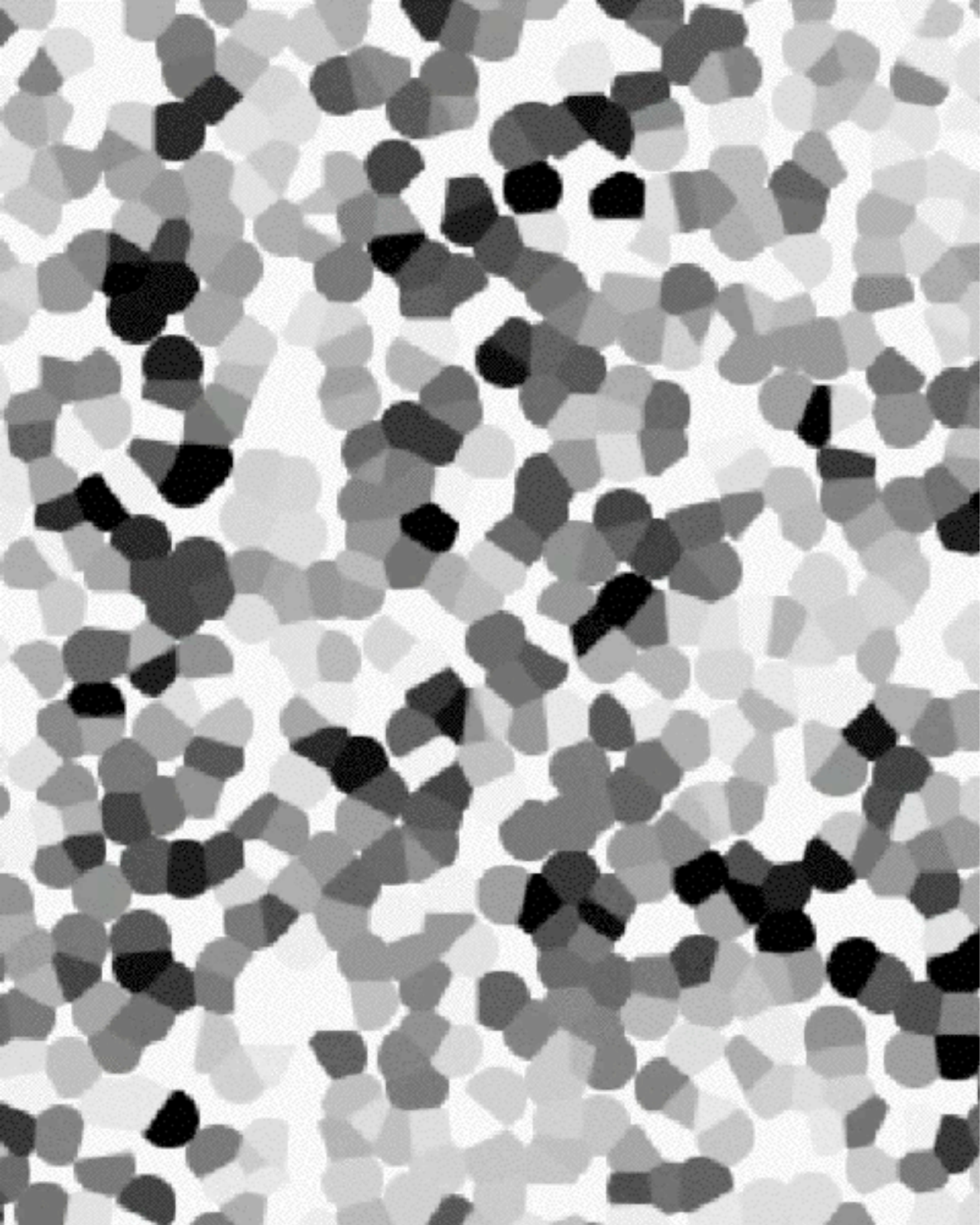
grains

ϕ_{min} 0.5 (2D) 0.55 (3D)

$\phi_{min} < \phi < \phi_{Max}$

ϕ_{max} 0.8 (2D) 0.65 (3D)

from the grains to the fluid



~~grains~~

continuum media
hypothesis

from the grains to the fluid



- Looking for a continuum description
- Lot of recent experiments in simple configurations: shear/ inclined plane, with model material (glass beads, sand...)
- Simulations with Contact Dynamics (disks, polygons, spheres)

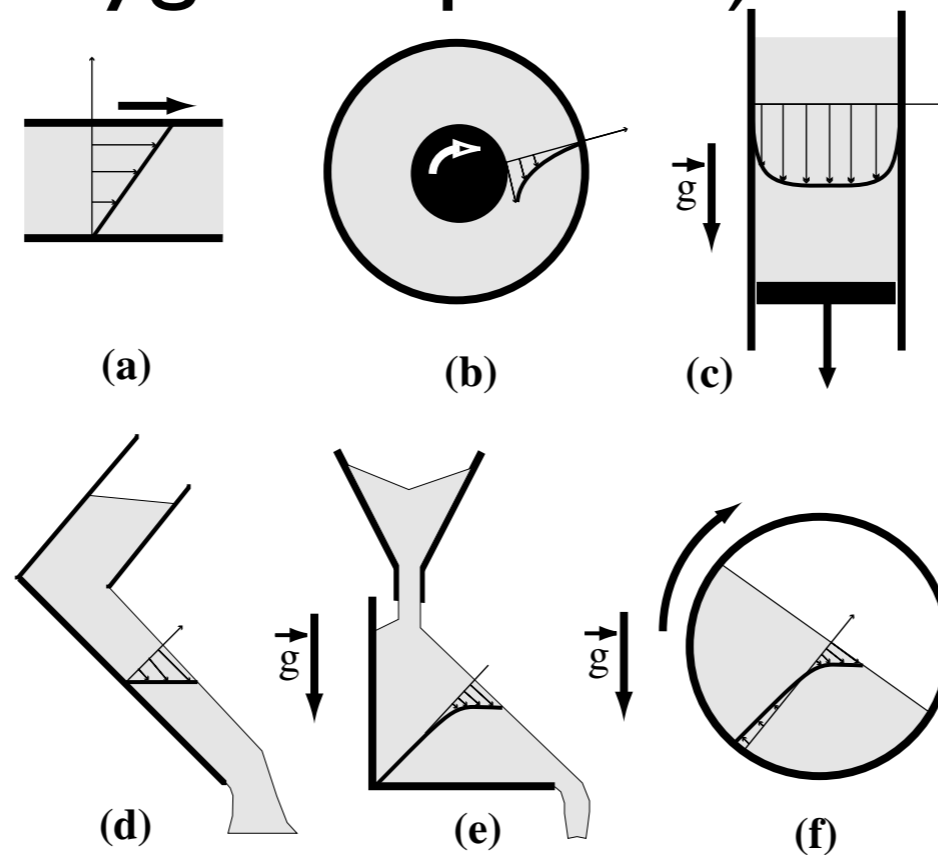


Fig. 1. The six configurations of granular flows: (a) plane shear, (b) annular shear, (c) vertical-chute flows, (d) inclined plane, (e) heap flow, (f) rotating drum.



- Looking for a continuum description
- Lot of recent experiments in simple configurations: shear/ inclined plane,
with model material (glass beads, sand...)
- Simulations with Contact Dynamics
(disks, polygons, spheres)
- Defining a «viscosity»
- Implement it in the Navier Stokes solver *Gerris*
- Test on exact «Bagnold» avalanche solution
- Test on granular collapse and hourglass



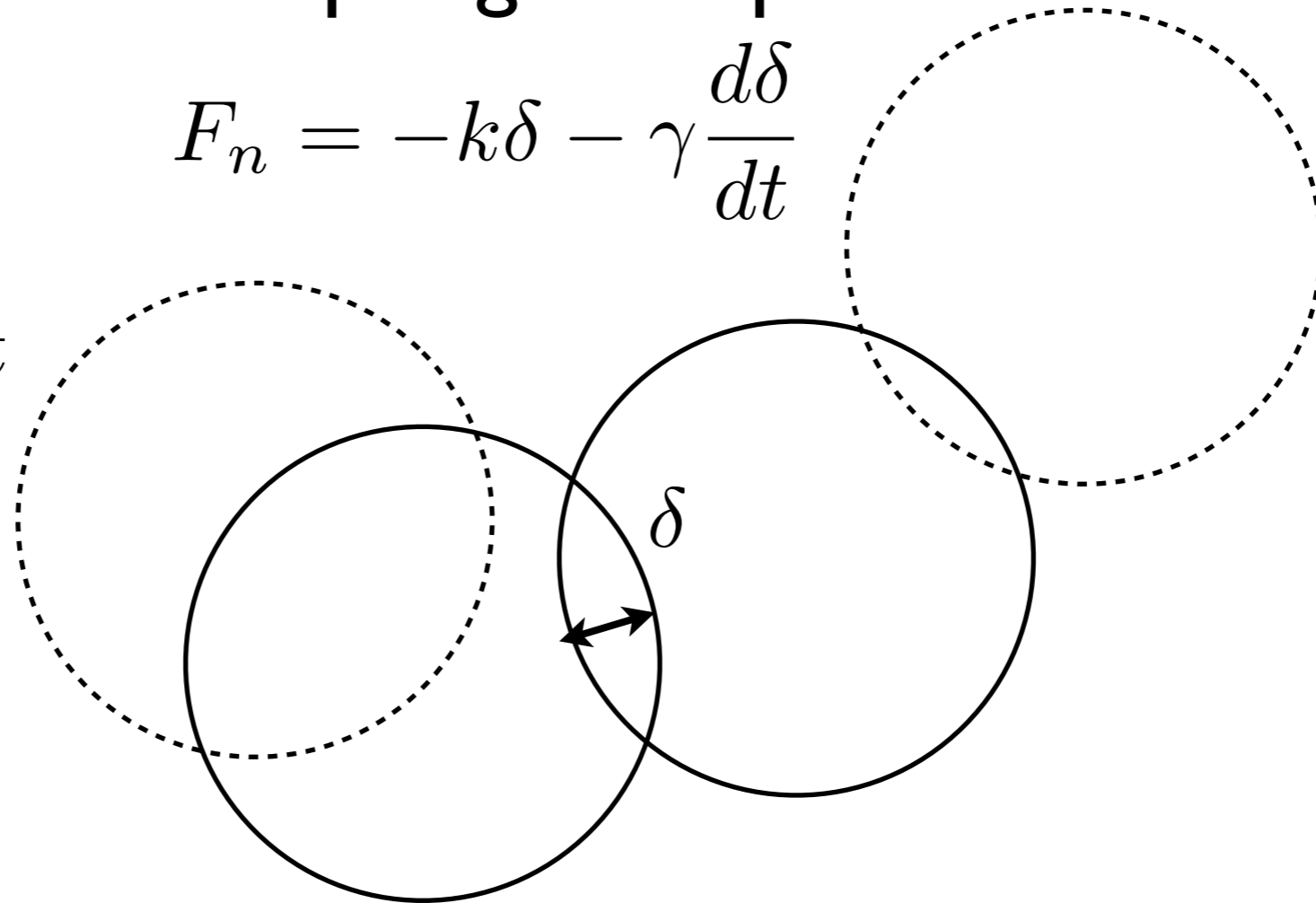
Molecular Dynamics:

$$m \frac{d}{dt} \vec{U} = \vec{F} + \vec{F}_n + \vec{F}_t \quad \text{Newton's equations}$$

branch a spring -dashpot

$$F_n = -k\delta - \gamma \frac{d\delta}{dt}$$

$t + \Delta t$



tangential Coulombic Friction $F_t < \mu F_n$



Contact Dynamics de (Moreau 1988)

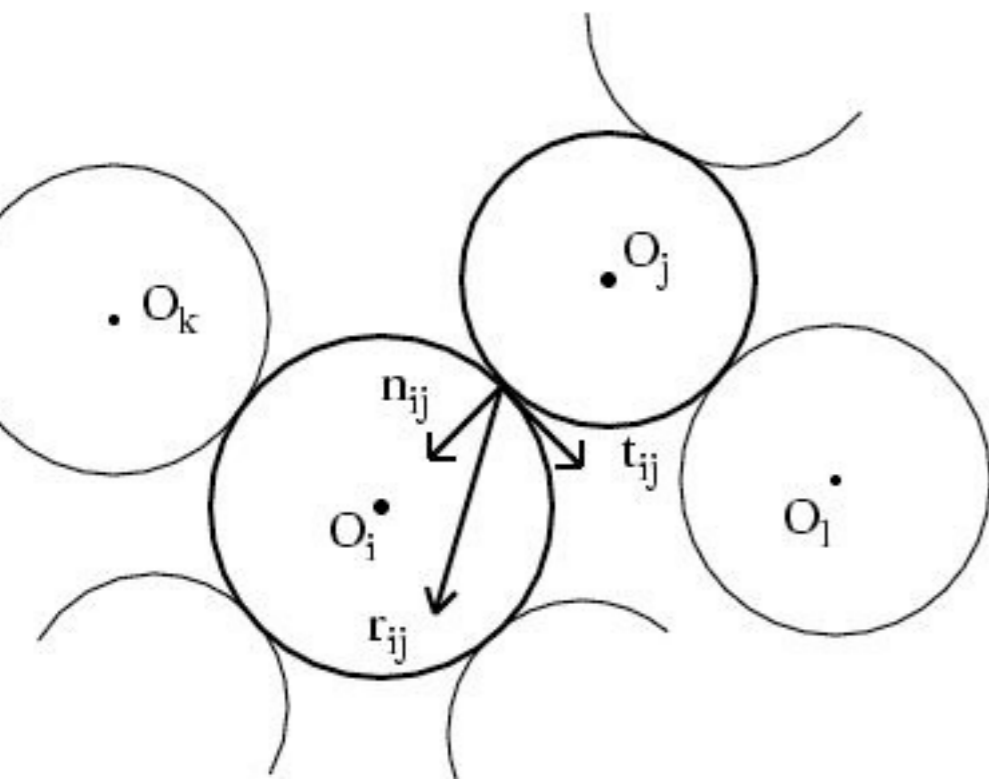
rigid grains

coefficient of friction μ

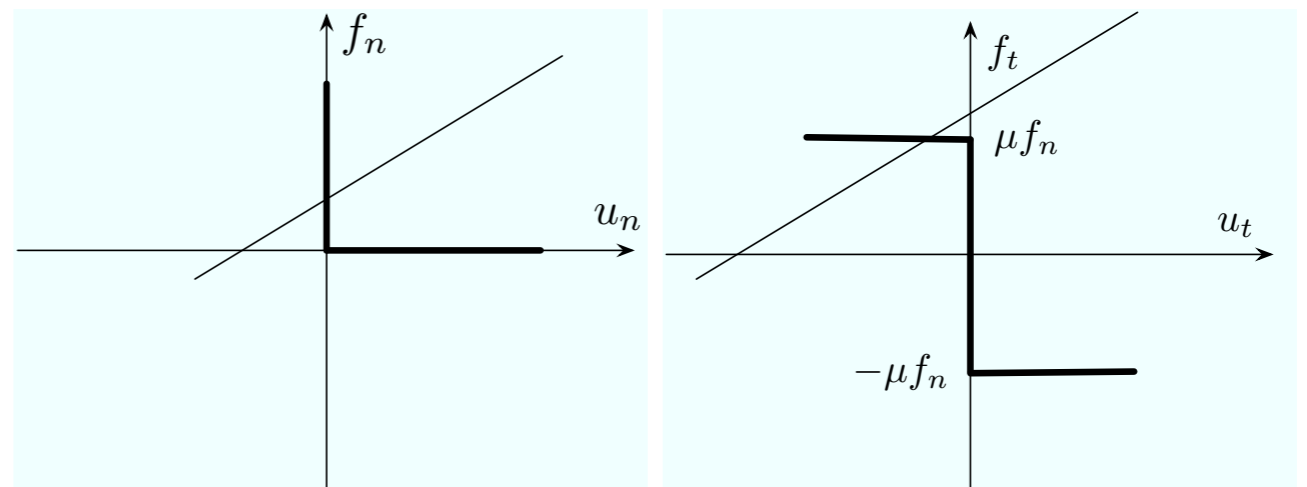
$$m(\vec{U}^+ - \vec{U}^-) = \vec{F} \delta t \quad \text{Newton's equations}$$

take the form of an equality between the change of momenta and the average impulse during δt .

written for each grain at the contact

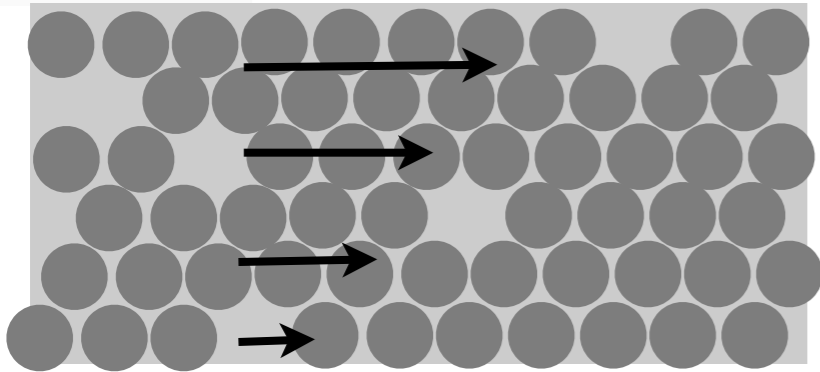


u_n, u_t

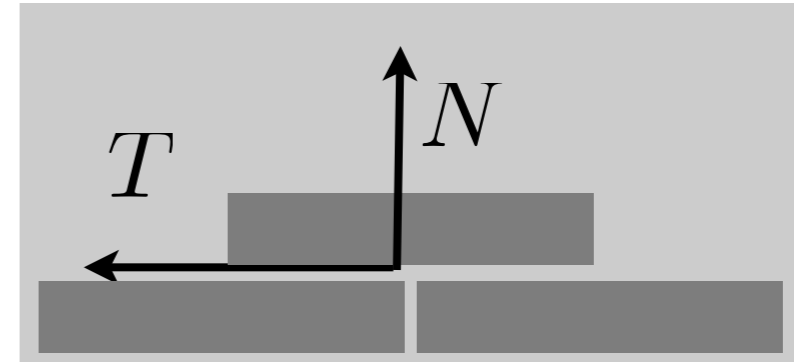


The $\mu(I)$ -rheology

$V(z)$



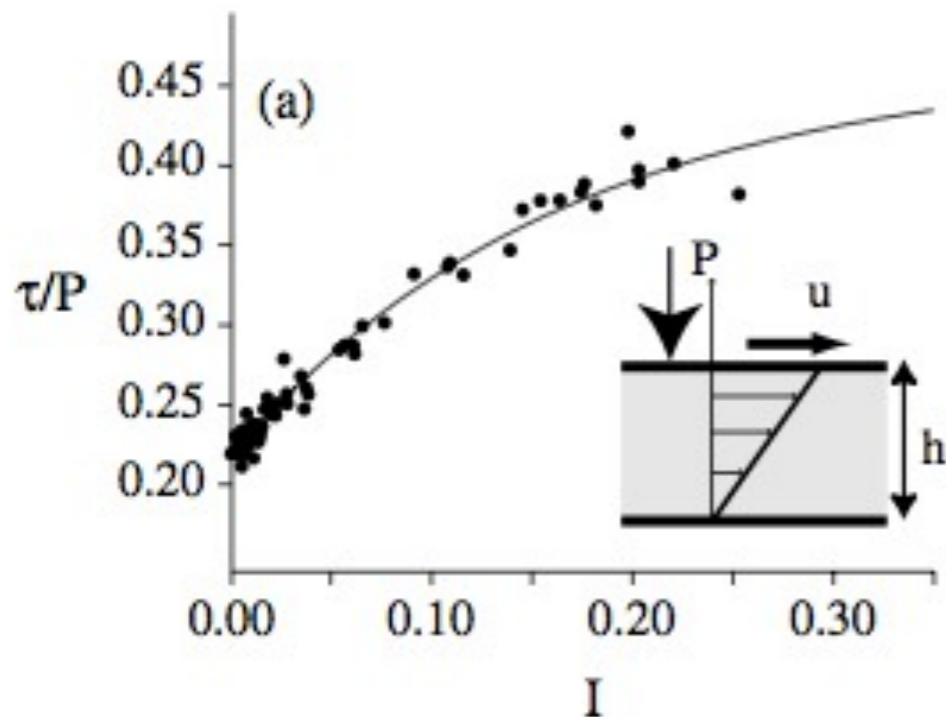
$V(z)$



$$T = \mu N$$

constitutive law?

Coulomb dry friction
Coulomb friction law

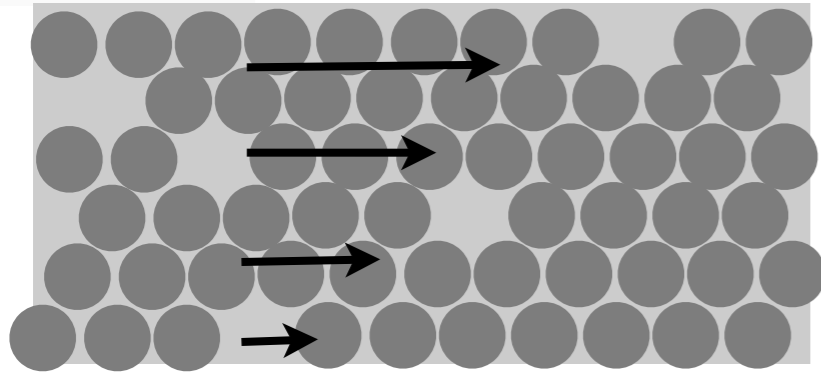


$$\tau = \mu P$$

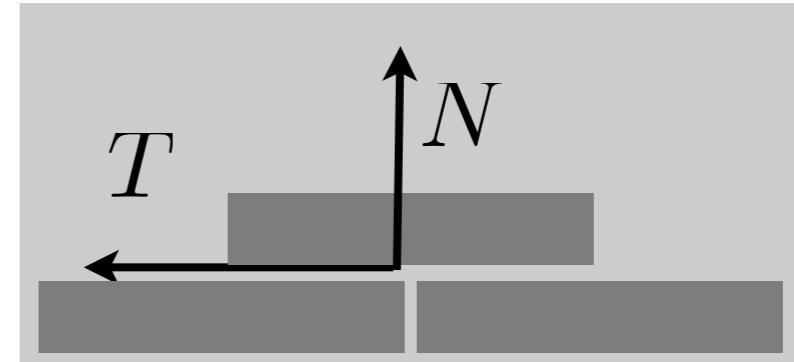


The $\mu(I)$ -rheology

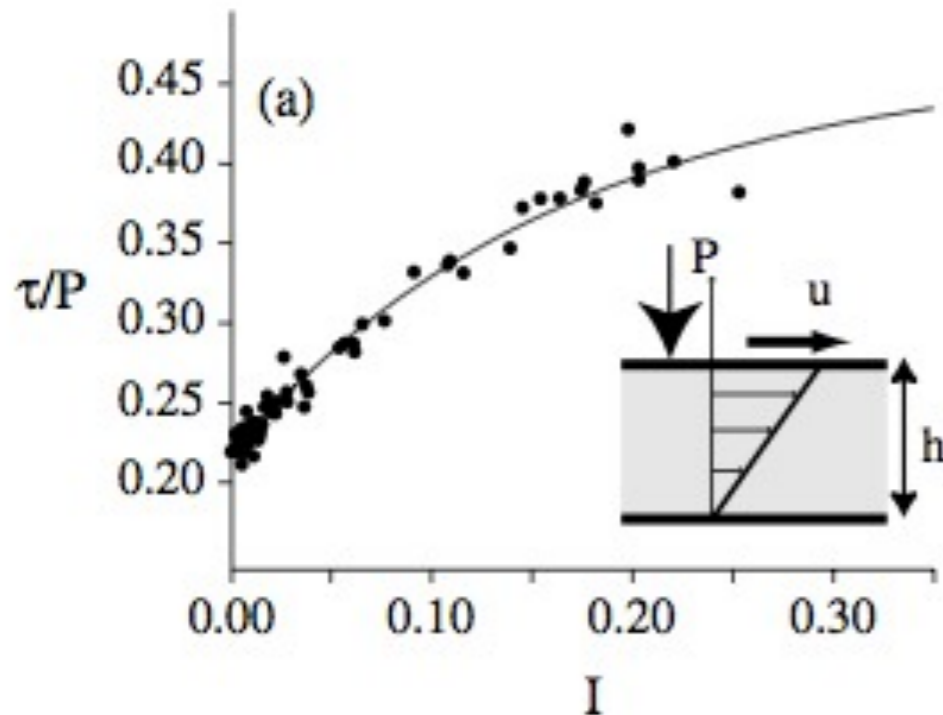
$V(z)$



$V(z)$



$$T = \mu N$$



Coulomb friction law

$$I = \frac{d \frac{\partial u}{\partial y}}{\sqrt{P/\rho}}$$

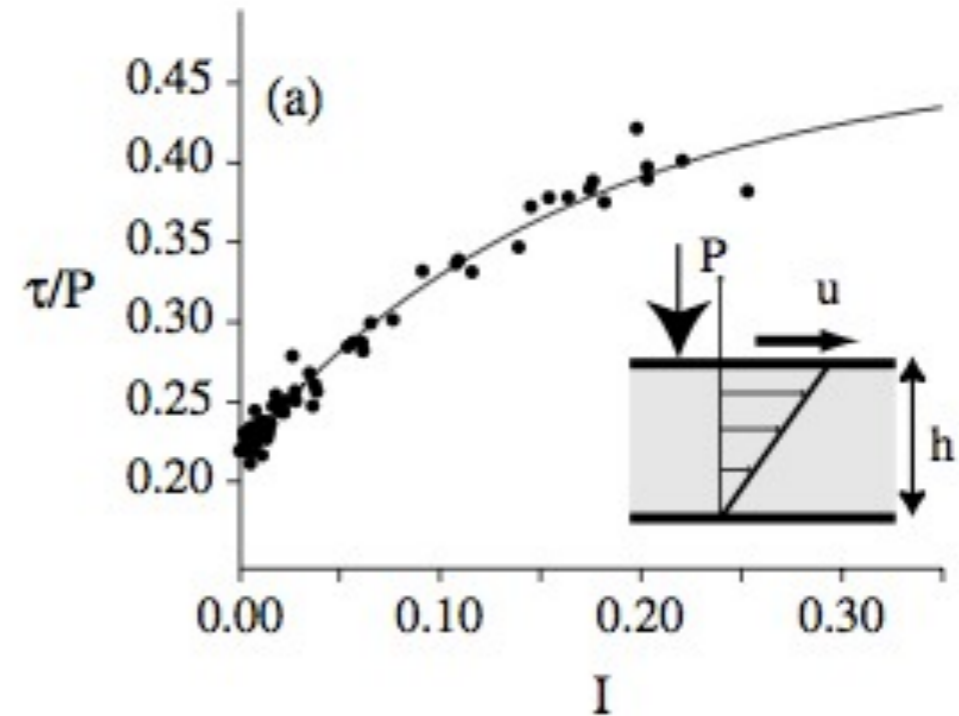
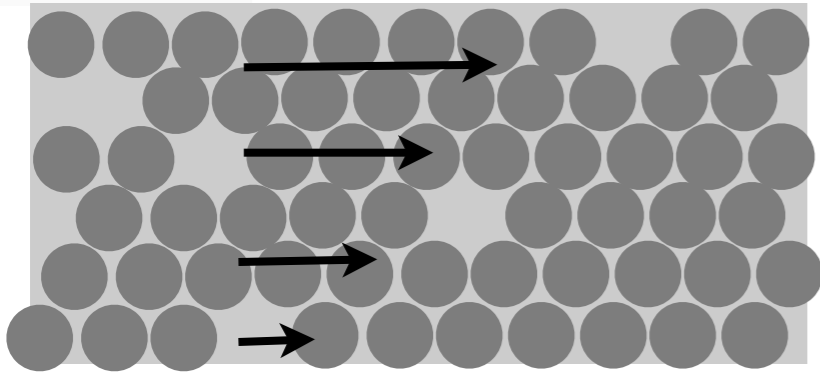
$$\tau = \mu(I)P$$

non dimensional number: «Froude»
 local «Inertial Number» (Da Cruz 04-05)

The $\mu(I)$ -rheology



$V(z)$

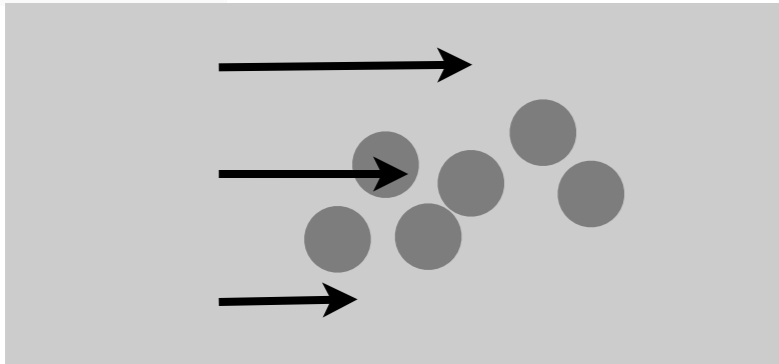


falling time
displacement time

$$I = \frac{d \frac{\partial u}{\partial y}}{\sqrt{P/\rho}}$$

non dimensional number: «Froude»
local «Inertial Number» (Da Cruz 04-05)

The $\mu(I)$ -rheology



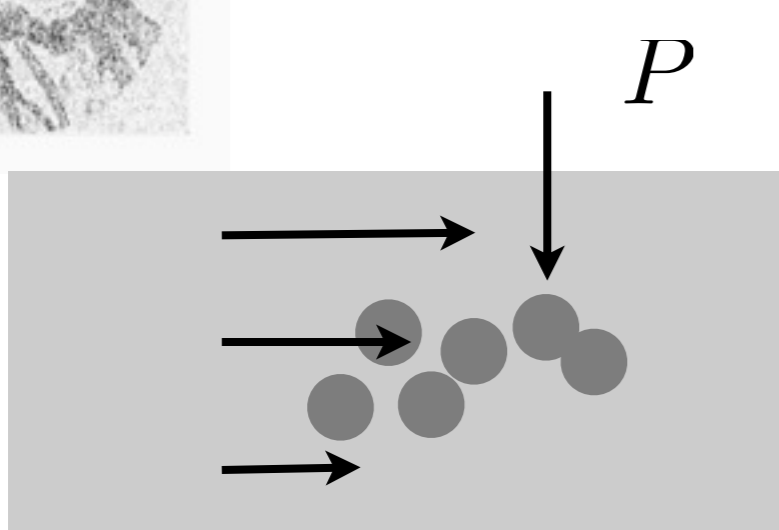
$$\frac{dx}{dt} = d \frac{\partial u}{\partial y} \quad t = 1 / \frac{\partial u}{\partial y}$$

falling time

displacement time

$$I = \frac{d \frac{\partial u}{\partial y}}{\sqrt{P/\rho}}$$

The $\mu(l)$ -rheology



$$m d^2 y / dt^2 = P d^2$$

$$t^2 = \rho d^2 / (P)$$

$$t = 1 / \frac{\partial u}{\partial y}$$

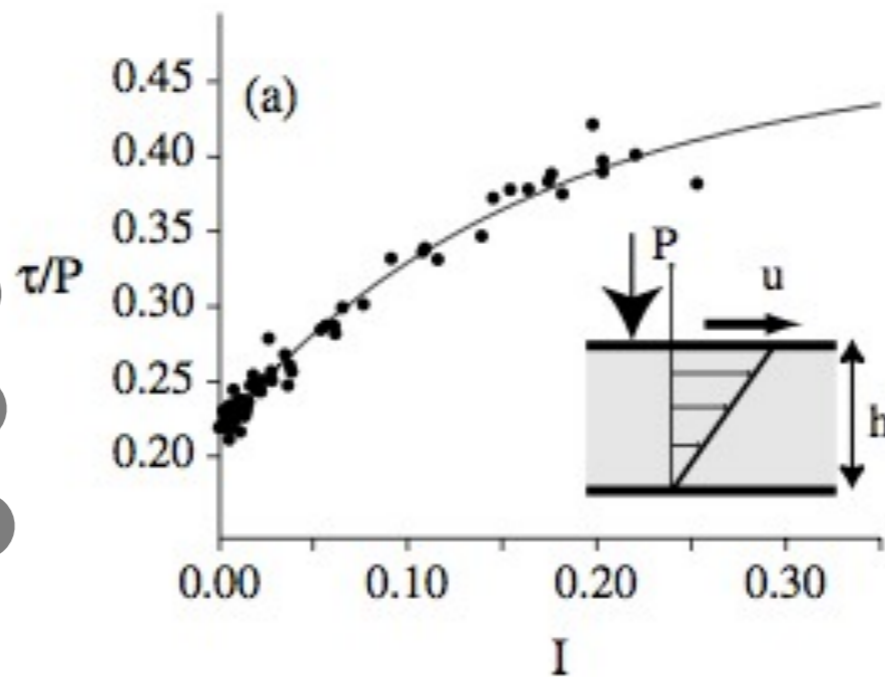
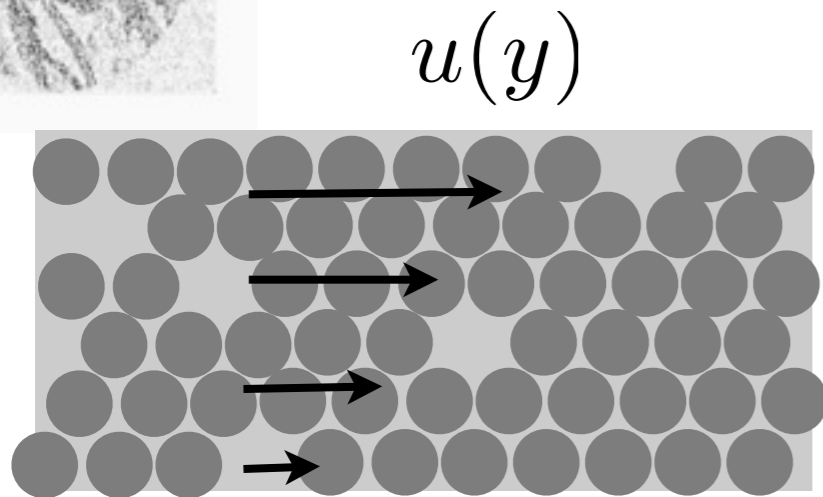
falling time

displacement time

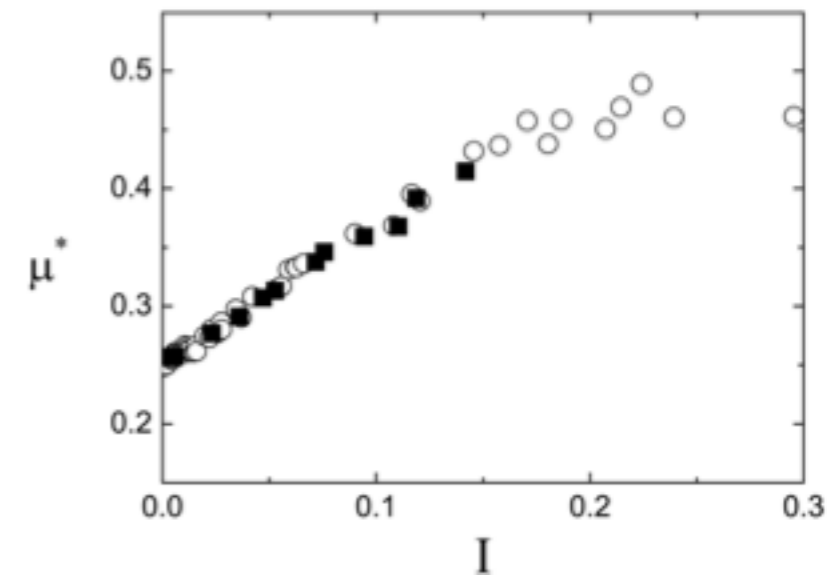
$$I = \frac{d \frac{\partial u}{\partial y}}{\sqrt{P / \rho}}$$



The $\mu(I)$ -rheology



by grain dynamics



Da Cruz PRE 05

Coulomb friction law

$$\tau = \mu(I)P$$

$\frac{\text{falling time}}{\text{displacement time}}$

$$I = \frac{d \frac{\partial u}{\partial y}}{\sqrt{P/\rho}}$$

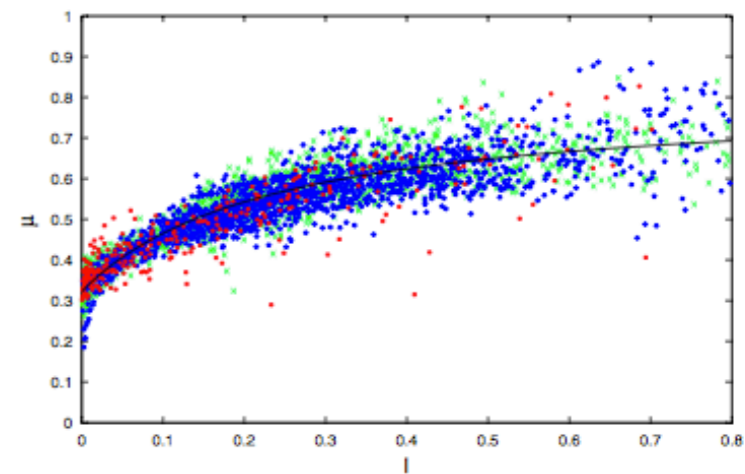
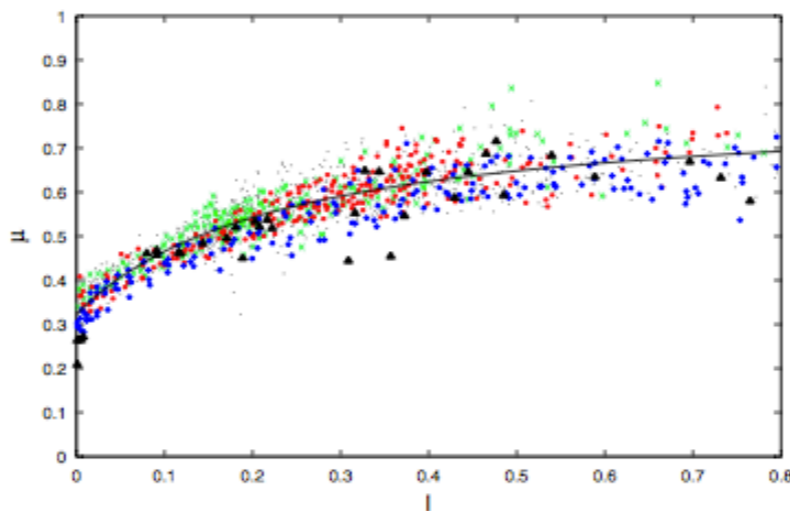
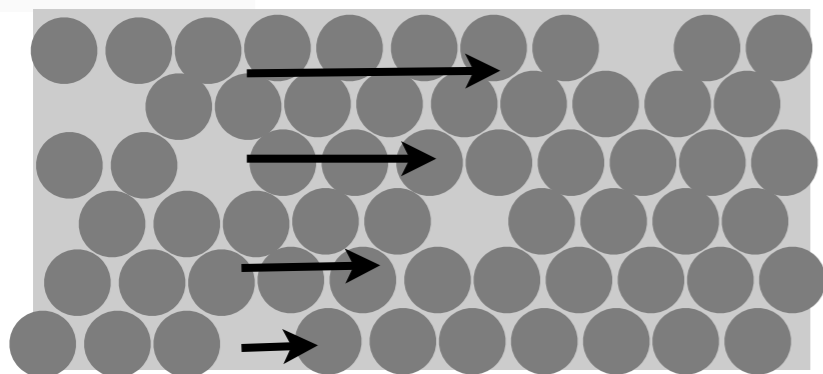
- Pouliquen 99
- Pouliquen Forterre JSM 06
- Da Cruz 04-05
- GDR Midi 04
- Josserand Lagrée Lhuillier 04

The $\mu(I)$ -rheology

by grain dynamics



$u(y)$



Lacaze Kerswell 09

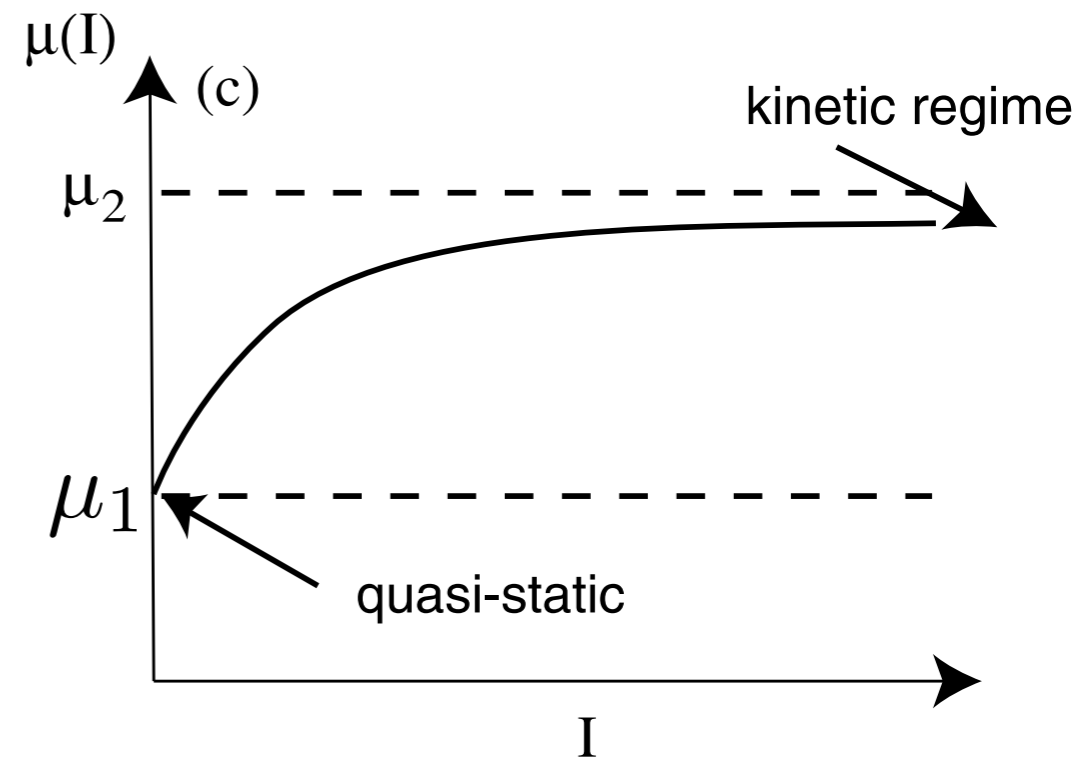
Coulomb friction law

$$\tau = \mu(I)P$$

$$I = \frac{d \frac{\partial u}{\partial y}}{\sqrt{P/\rho}}$$

$$\mu(I) = \mu_1 + \frac{\mu_2 - \mu_1}{I_0/I + 1}$$

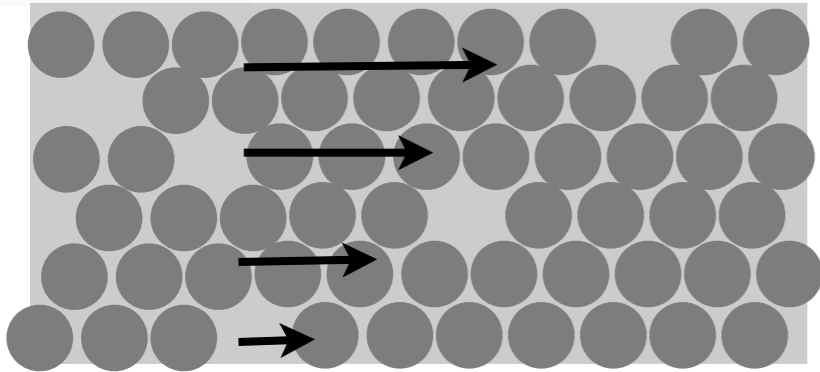
$$\mu_1 \simeq 0.32 \quad (\mu_2 - \mu_1) \simeq 0.23 \quad I_0 \simeq 0.3$$



The $\mu(I)$ -rheology

Jop Forterre Pouliquen 2005

$u(y)$



Coulomb friction law

$$\tau = \mu(I)P$$

$$I = \frac{d \frac{\partial u}{\partial y}}{\sqrt{P/\rho}}$$

«Drucker-Prager»
plastic flow

«equivalent» viscosity

$$\eta \frac{\partial u}{\partial y} = \mu(I)p \quad \rightarrow \quad \eta = \frac{\mu(I)p}{\frac{\partial u}{\partial y}}$$



The $\mu(I)$ -rheology implementation in Navier Stokes?

Jop Forterre Pouliquen 2005

$$\mu(I) = \mu_1 + \frac{\mu_2 - \mu_1}{I_0/I + 1}$$

$$D_2 = \sqrt{D_{ij}D_{ij}} \quad D_{ij} = \frac{u_{i,j} + u_{j,i}}{2}$$

construction of a viscosity based on the D_2 invariant and redefinition of I

$$I = d\sqrt{2}D_2 / \sqrt{(|p|/\rho)}.$$

$$\eta = \left(\frac{\mu(I)}{\sqrt{2}D_2} p \right) \quad \text{«Drucker-Prager»}$$

$$\nabla \cdot \mathbf{u} = 0, \quad \rho \left(\frac{\partial \mathbf{u}}{\partial t} + \mathbf{u} \cdot \nabla \mathbf{u} \right) = -\nabla p + \nabla \cdot (2\eta \mathbf{D}) + \rho g,$$

Boundary Conditions: no slip and $P=0$ at the interface



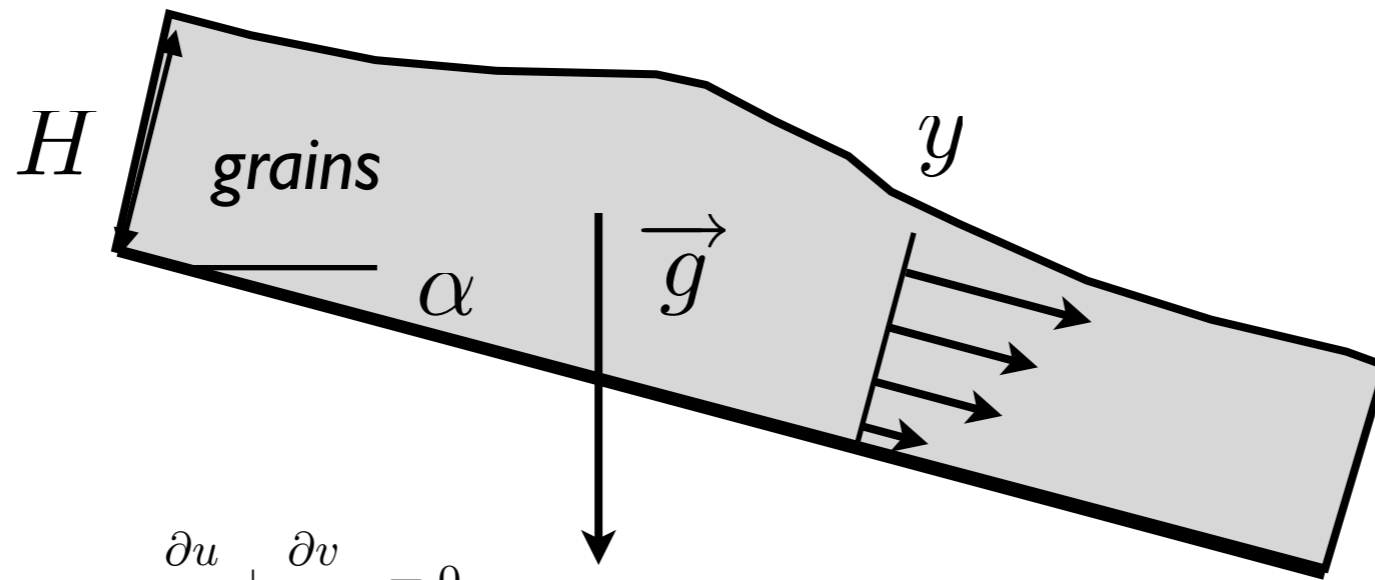
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S^T VENANT



● Couche Mince Saint Venant Shallow Water Savage Hutter



$$\varepsilon = \dot{H}/L$$

$$\left\{ \begin{array}{l} \frac{\partial u}{\partial t} + \frac{\partial u^2}{\partial x} + \frac{\partial uv}{\partial y} \\ \rho \left(\frac{\partial u}{\partial t} + \frac{\partial u^2}{\partial x} + \frac{\partial uv}{\partial y} \right) \\ \frac{\partial v}{\partial t} + \frac{\partial uv}{\partial x} + \frac{\partial v^2}{\partial y} \\ \rho \left(\frac{\partial v}{\partial t} + \frac{\partial uv}{\partial x} + \frac{\partial v^2}{\partial y} \right) \end{array} \right. = \begin{array}{l} 0 \\ -g\rho \sin \alpha - \frac{\partial p}{\partial x} + \frac{\partial \tau_{xx}}{\partial x} + \frac{\partial \tau_{xy}}{\partial y} \\ -g\rho \cos \alpha - \frac{\partial p}{\partial y} + \frac{\partial \tau_{yx}}{\partial x} + \frac{\partial \tau_{yy}}{\partial y} \end{array}$$

$$p = -\rho g \cos \alpha (\eta(x, t) - y)$$

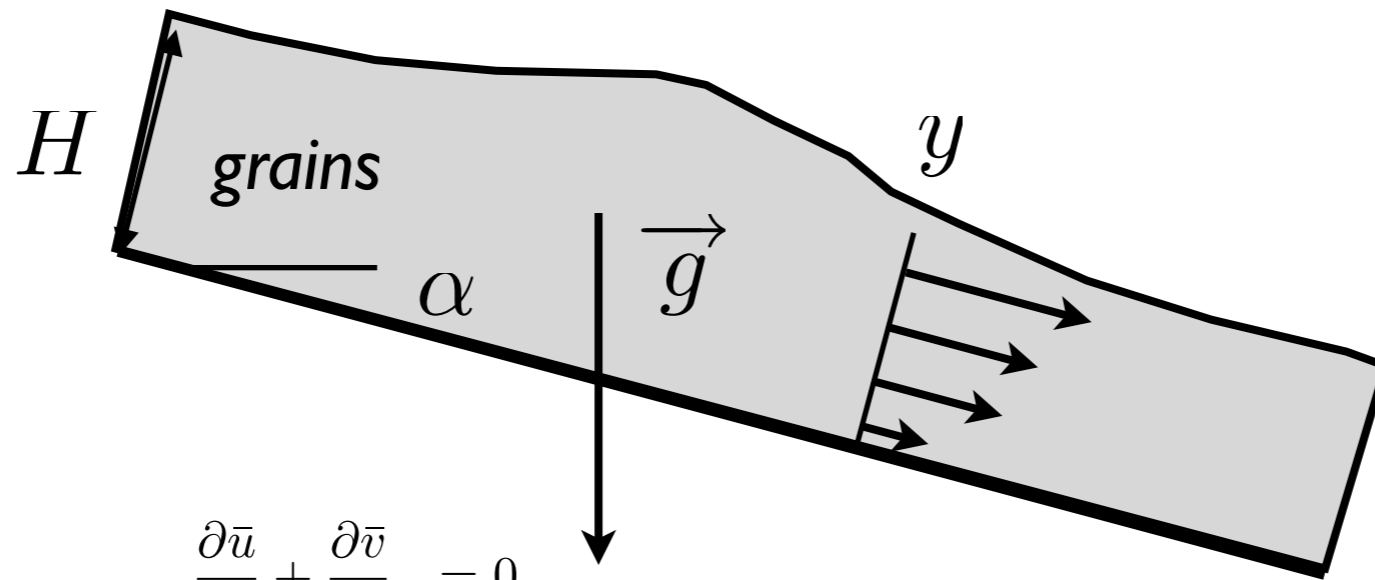
$$\tau_{xy} = \rho g H \sin \alpha \bar{\tau}_{xy}$$

$$\rho U_0^2 / L \longleftrightarrow -\frac{\partial p}{\partial x} = -\rho g \frac{\partial \eta}{\partial x}$$

$$\rho U_0^2 / (H/\varepsilon) = \rho g \varepsilon.$$

$$U_0 = \sqrt{gH}.$$

Saint-Venant Savage Hutter



$$\varepsilon = \dot{H}/L$$

$$\left\{ \begin{array}{l} \frac{\partial \bar{u}}{\partial \bar{x}} + \frac{\partial \bar{v}}{\partial \bar{y}} = 0 \\ \varepsilon \left(\frac{\partial \bar{u}}{\partial \bar{t}} + \frac{\partial \bar{u}^2}{\partial \bar{x}} + \frac{\partial \bar{u}\bar{v}}{\partial \bar{y}} \right) = -\sin \alpha - \varepsilon \frac{\partial \bar{p}}{\partial \bar{x}} + \sin \alpha \varepsilon \frac{\partial \bar{\tau}_{xx}}{\partial \bar{x}} + \sin \alpha \frac{\partial \bar{\tau}_{xy}}{\partial \bar{y}} \\ \varepsilon^2 \left(\frac{\partial \bar{v}}{\partial \bar{t}} + \frac{\partial \bar{u}\bar{v}}{\partial \bar{x}} + \frac{\partial \bar{v}^2}{\partial \bar{y}} \right) = -\cos \alpha - \frac{\partial \bar{p}}{\partial \bar{y}} + \sin \alpha \varepsilon \frac{\partial \bar{\tau}_{yy}}{\partial \bar{x}} + \sin \alpha \frac{\partial \bar{\tau}_{yx}}{\partial \bar{y}} \end{array} \right.$$

$$p = -\rho g \cos \alpha (\eta(x, t) - y)$$

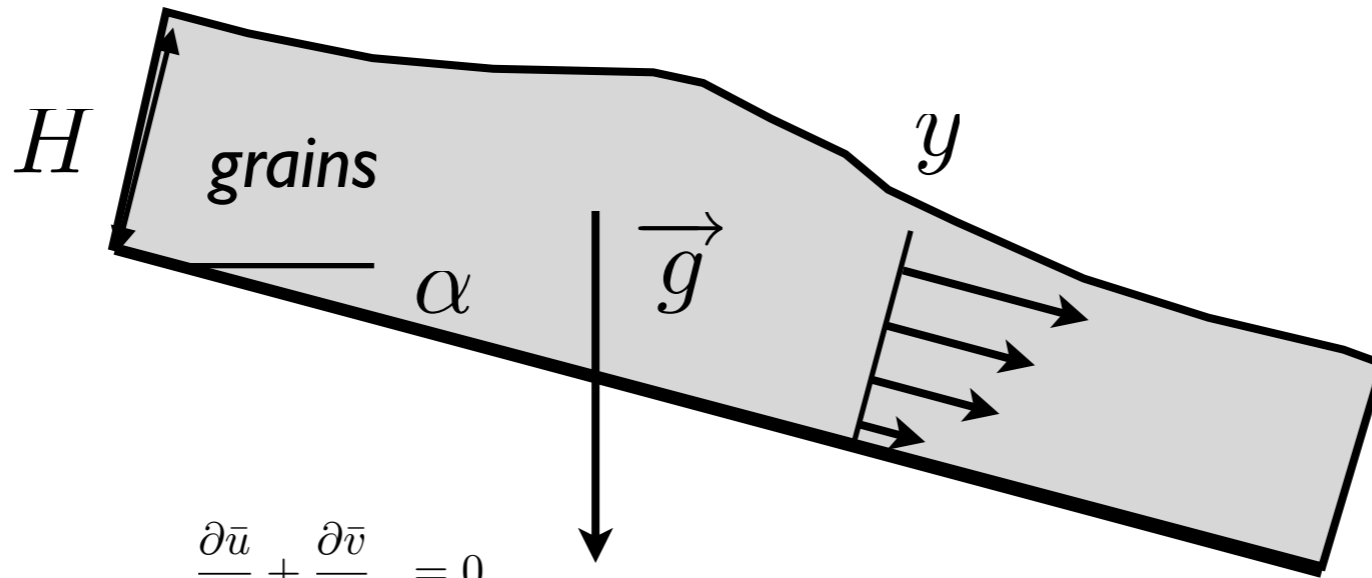
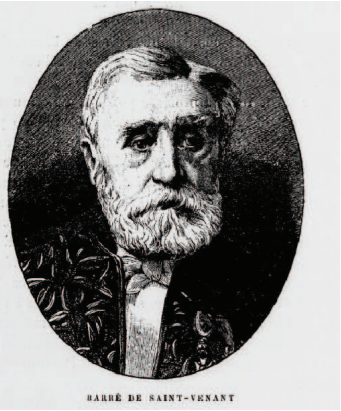
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$$\rho U_0^2 / (H/\varepsilon) = \rho g \varepsilon.$$

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Saint-Venant Savage Hutter



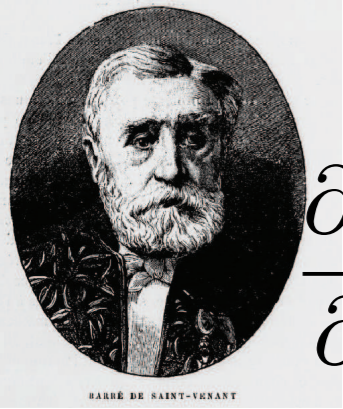
$$\left\{ \begin{array}{l} \frac{\partial \bar{u}}{\partial x} + \frac{\partial \bar{v}}{\partial y} = 0 \\ \varepsilon \left(\frac{\partial \bar{u}}{\partial t} + \frac{\partial \bar{u}^2}{\partial x} + \frac{\partial \bar{u}\bar{v}}{\partial y} \right) = -\sin \alpha - \varepsilon \frac{\partial \bar{p}}{\partial x} + \sin \alpha \frac{\partial \bar{\tau}_{xy}}{\partial y} \\ 0 = -\cos \alpha - \frac{\partial \bar{p}}{\partial y} \end{array} \right.$$

$$\int dy$$

$$\frac{\partial h}{\partial t} + \frac{\partial}{\partial x} \int_f^\eta u(x, y, t) dy = 0$$

$\mu(I)$ at the wall

$$\frac{\partial}{\partial t} \int_f^\eta u(x, y, t) dy + \frac{\partial}{\partial x} \left(\int_f^\eta u(x, y, t)^2 dy + \cos \alpha \frac{g}{2} (h^2) \right) = -gh \sin(\alpha) - \cos \alpha gh \frac{d}{dx} f - \mu(I(0)) gh \cos(\alpha).$$



Saint-Venant Savage Hutter

$$\frac{\partial h}{\partial t} + \frac{\partial(Q)}{\partial x} = 0 \quad \frac{\partial Q}{\partial t} + \frac{\partial}{\partial x} \left(\frac{5Q^2}{4h} + \frac{g}{2}(h^2) \right) = -gh\mu(I) \frac{Q}{|Q|}$$

Integral over the layer of grains



Saint-Venant Savage Hutter *Gerris*

$$\frac{\partial h}{\partial t} + \frac{\partial(Q)}{\partial x} = 0 \quad \frac{\partial Q}{\partial t} + \frac{\partial}{\partial x} \left(\frac{5Q^2}{4h} + \frac{g}{2}(h^2) \right) = -gh\mu(I) \frac{Q}{|Q|}$$

Gerris is a free finite volume code by Stéphane Popinet
one part of the code is a Shallow Water solver

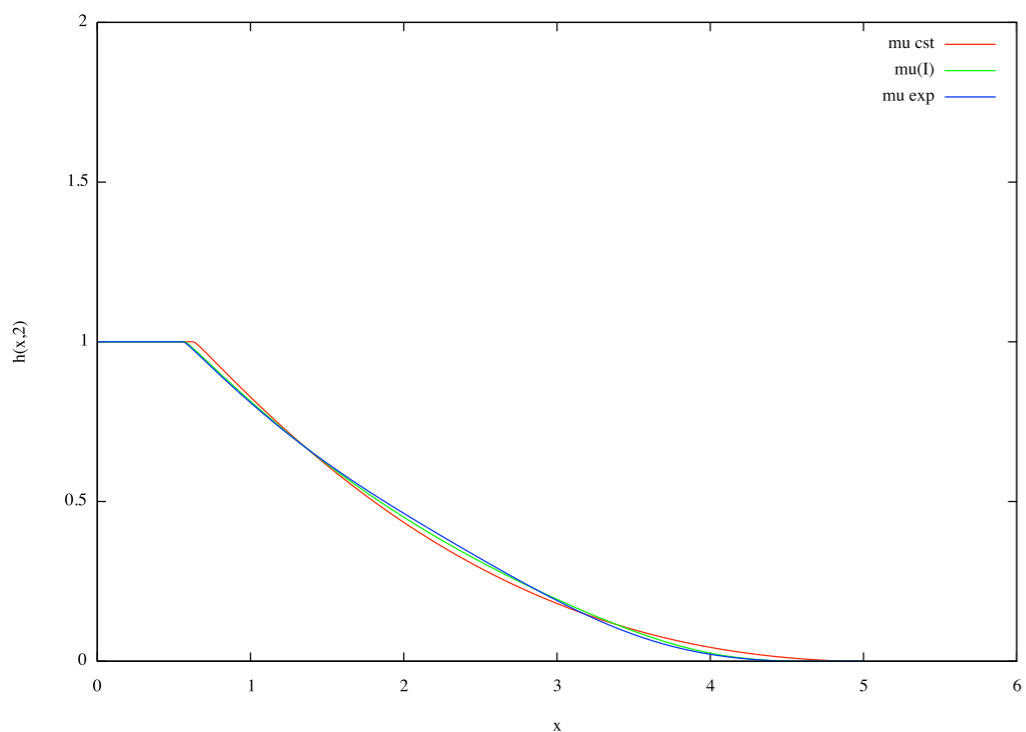
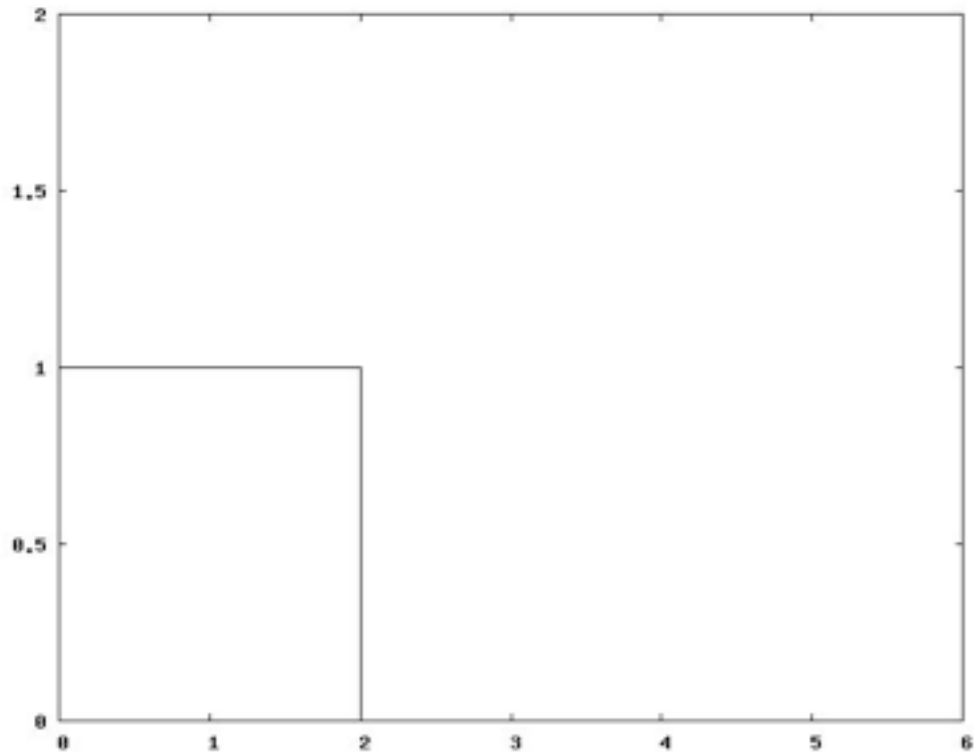
$$\frac{Q^* - Q^n}{\Delta t} + \frac{\partial}{\partial x} \left(\frac{Q^2}{h} + \frac{g}{2}(h^2) \right) = 0 \quad \frac{Q^{n+1} - Q^*}{\Delta t} = -gh^* \mu(I^*) \frac{Q^{n+1}}{|Q^*|}$$

Audusse et al.



Saint-Venant Savage Hutter Gerris

$$\frac{\partial h}{\partial t} + \frac{\partial(Q)}{\partial x} = 0 \quad \frac{\partial Q}{\partial t} + \frac{\partial}{\partial x} \left(\frac{5Q^2}{4h} + \frac{g}{2}(h^2) \right) = -gh\mu(I) \frac{Q}{|Q|}$$



$$\mu(I) = \mu_s$$

$$\mu(I) = \mu_s + \frac{\Delta\mu}{\frac{I_0}{I} + 1}$$

$$\mu(I) = \mu_s + \Delta\mu e^{-\beta/I}$$

$$\mu = 0.45$$

$$\mu = (0.4 + 0.26/(0.4/\ln + 1))$$

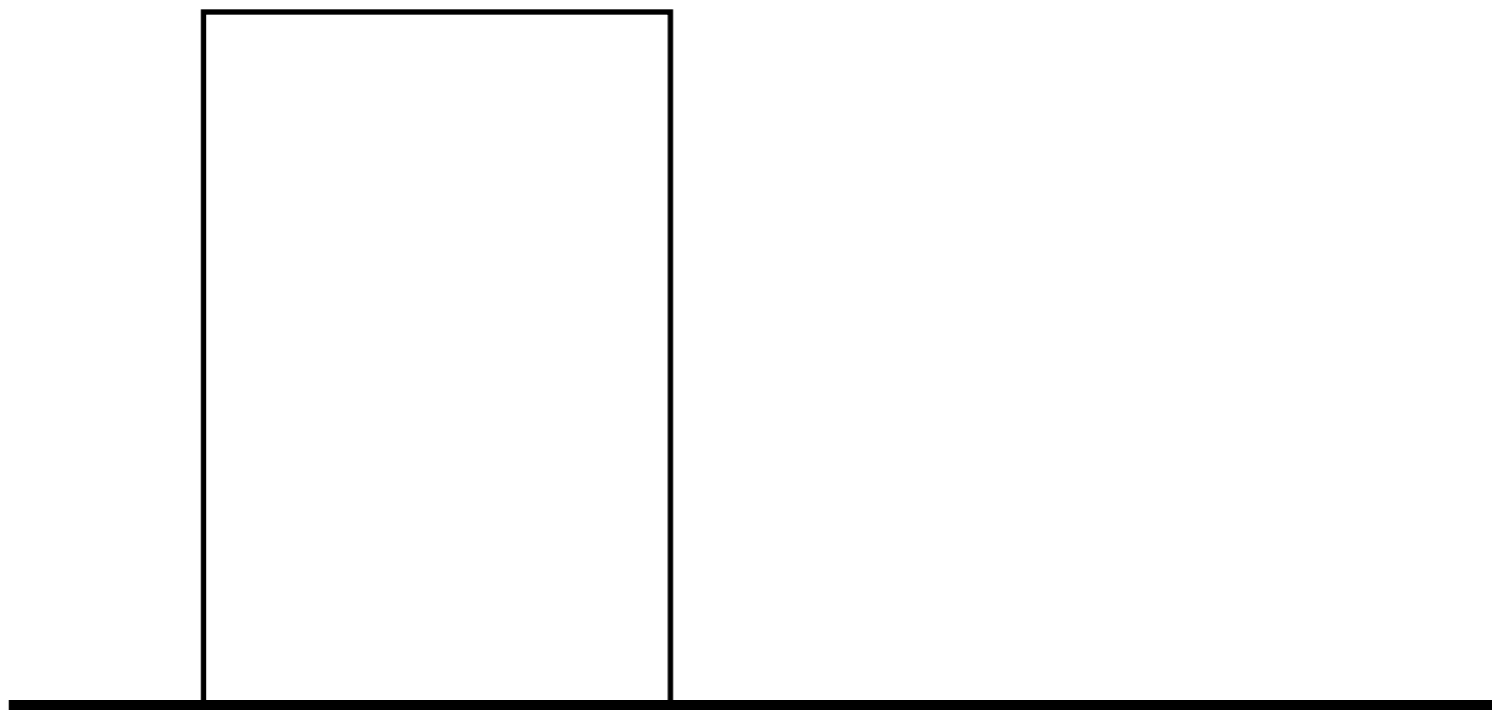
$$\mu = (0.4 + 0.26 * \exp(-0.136/\ln))$$



Saint-Venant Savage Hutter *Gerris*

$$\frac{\partial h}{\partial t} + \frac{\partial(Q)}{\partial x} = 0 \quad \frac{\partial Q}{\partial t} + \frac{\partial}{\partial x} \left(\frac{5Q^2}{4h} + \frac{g}{2}(h^2) \right) = -gh\mu(I) \frac{Q}{|Q|}$$

valid by hypothesis for small aspect ratio



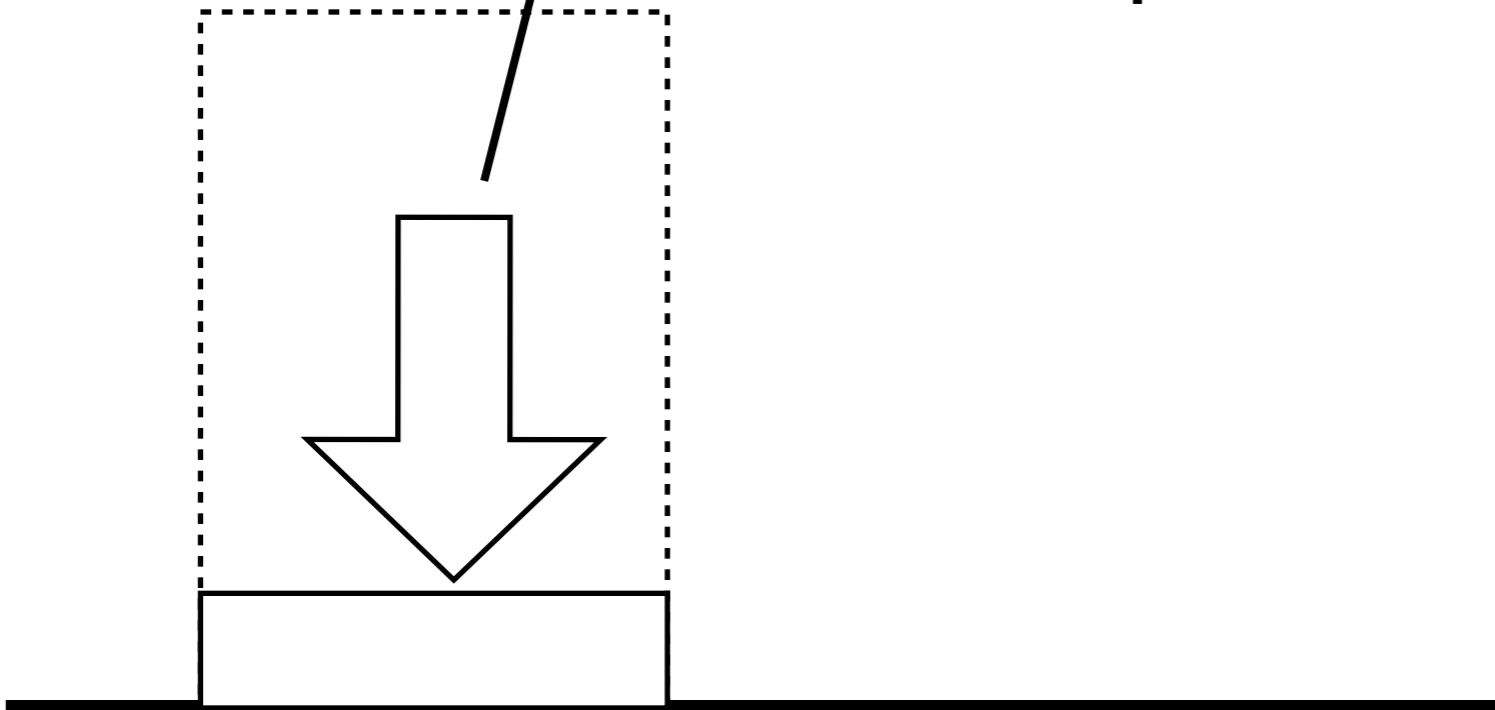


Saint-Venant Savage Hutter *Gerris*

$$\frac{\partial h}{\partial t} + \frac{\partial(Q)}{\partial x} = 0$$

$$\frac{\partial Q}{\partial t} + \frac{\partial}{\partial x} \left(\frac{5Q^2}{4h} + \frac{g}{2}(h^2) \right) = -gh\mu(I) \frac{Q}{|Q|}$$

add a source term corresponding to pluviation

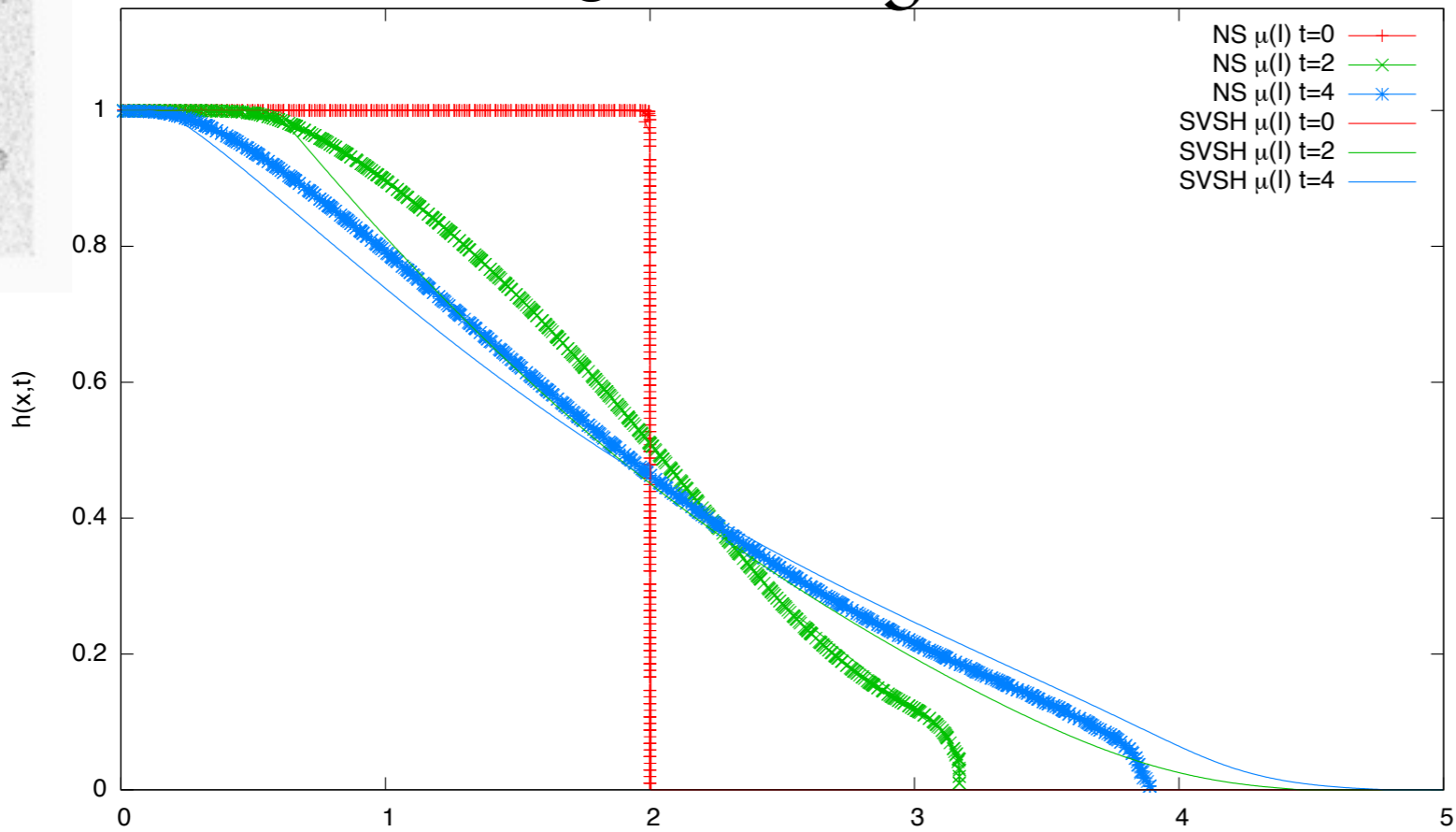




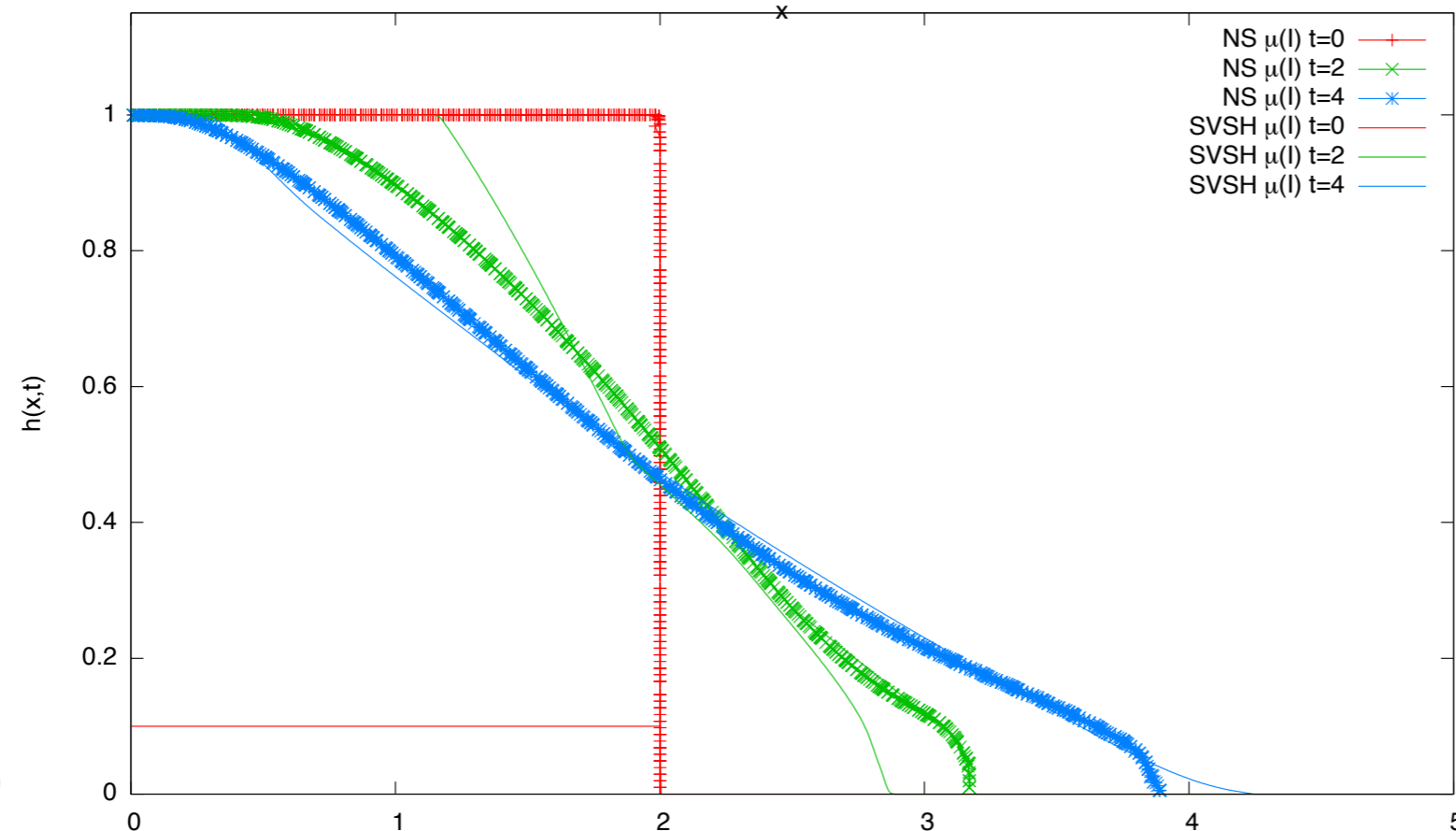
Saint-Venant Savage Hutter *Gerris*

better description of the run-out

no pluviation



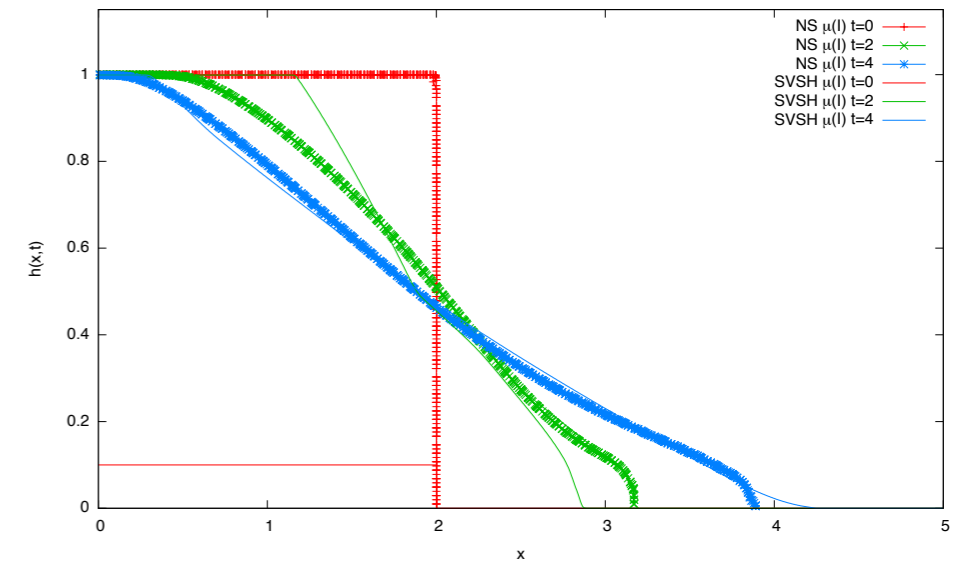
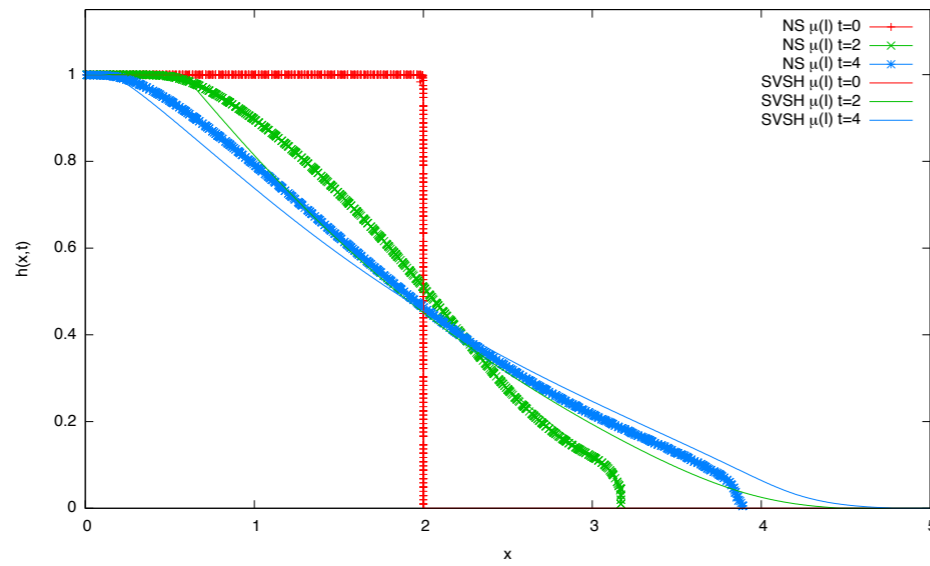
pluviation



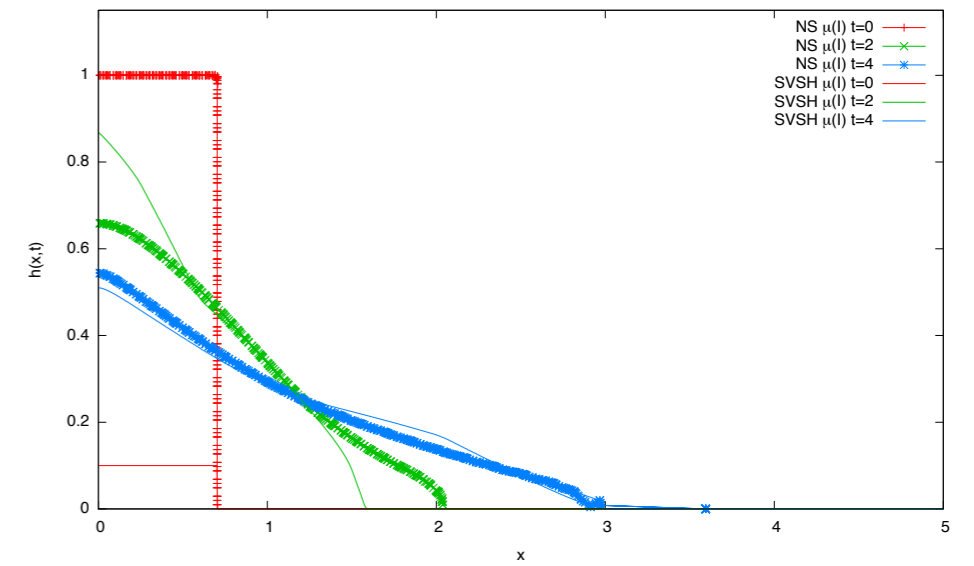
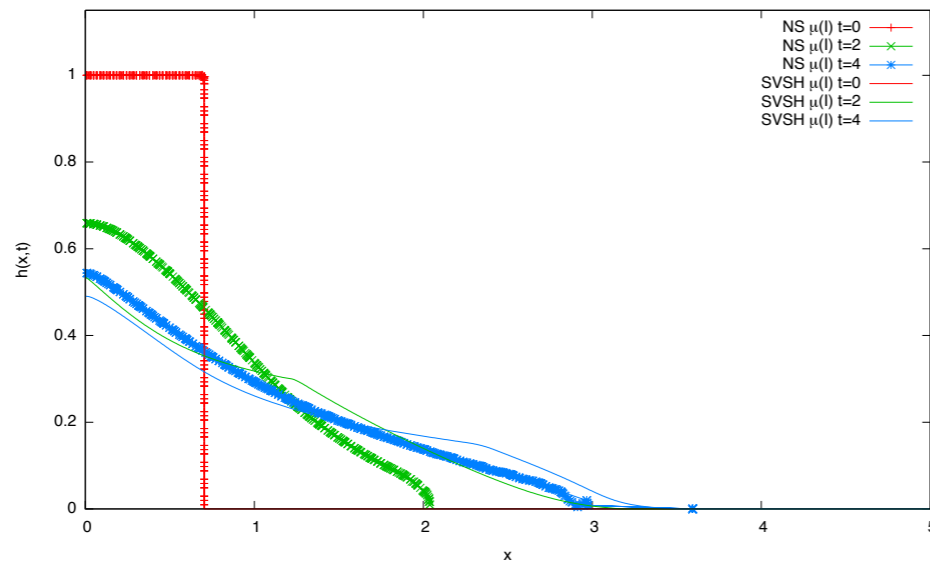


Saint-Venant Savage Hutter *Gerris*

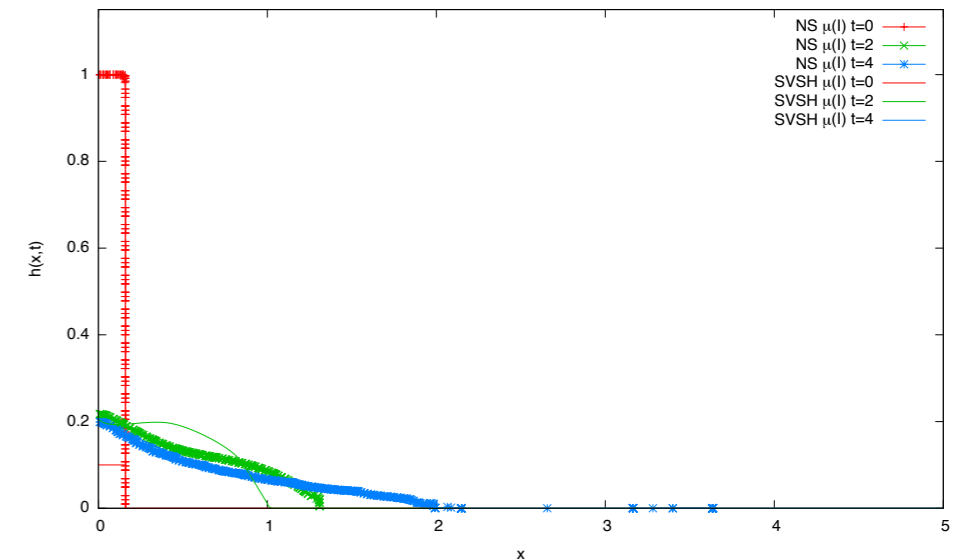
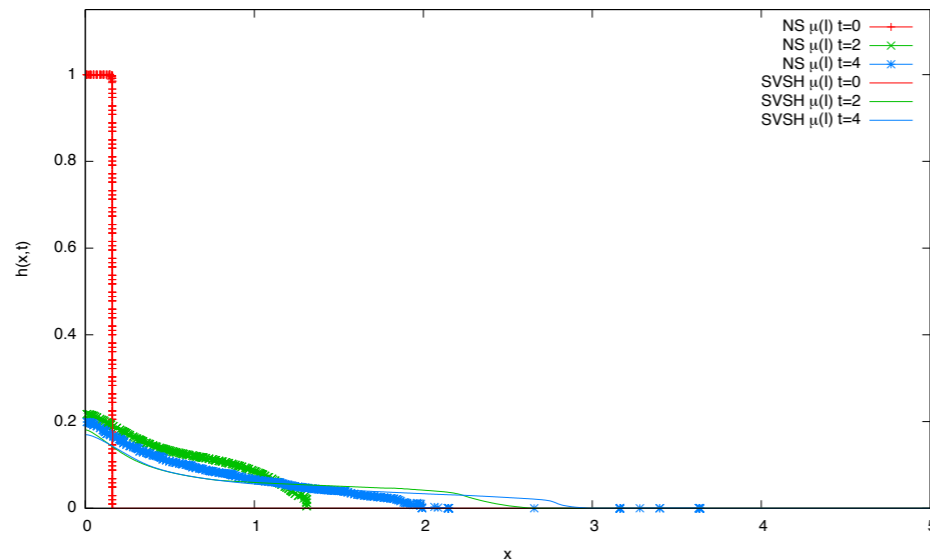
better description of the run-out



no pluviation

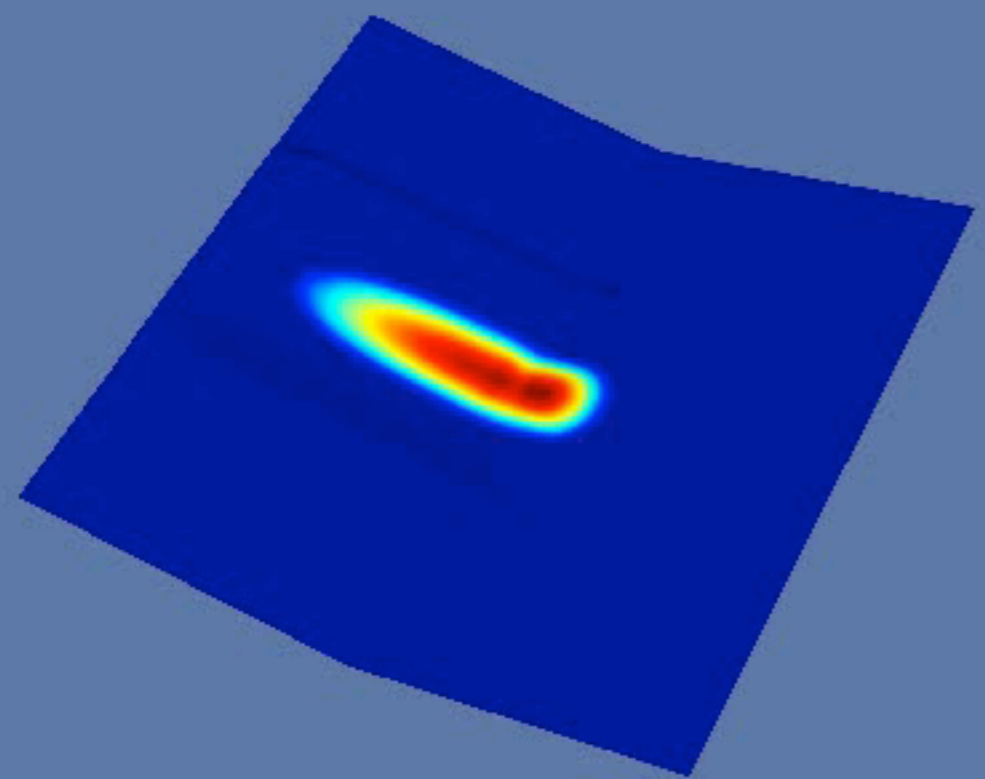
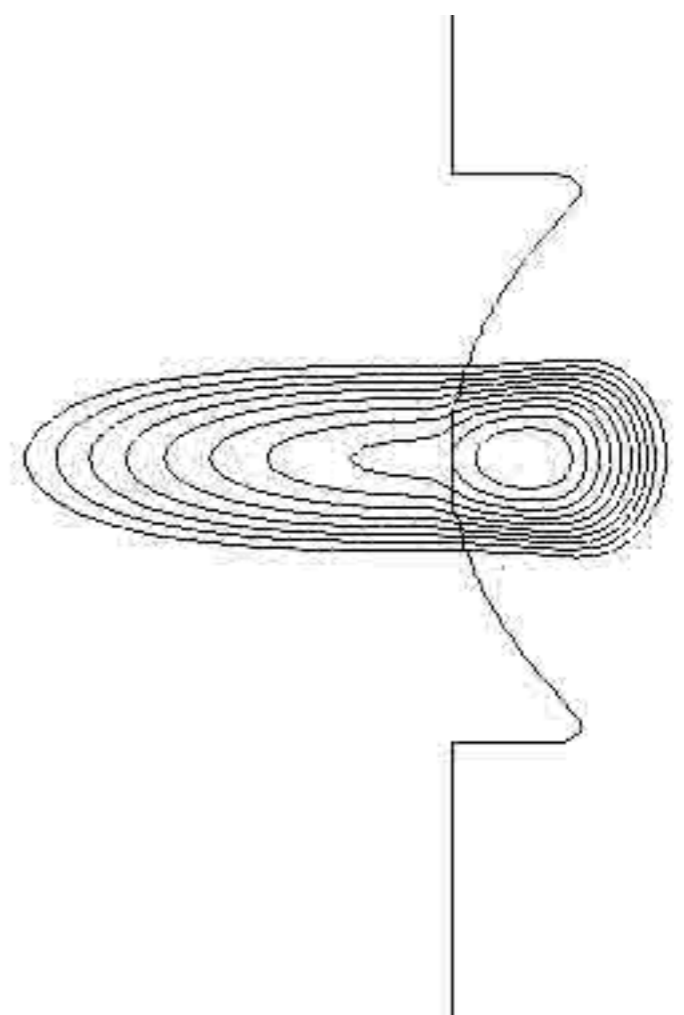


pluviation



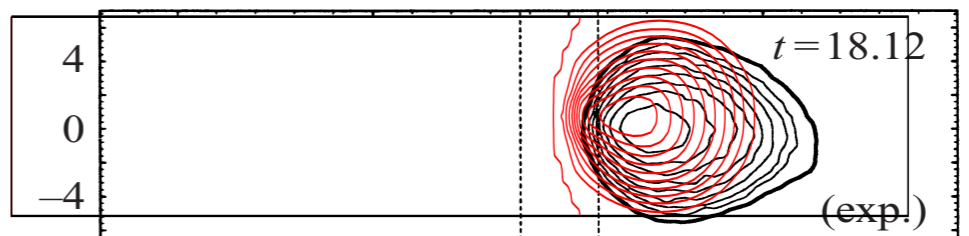
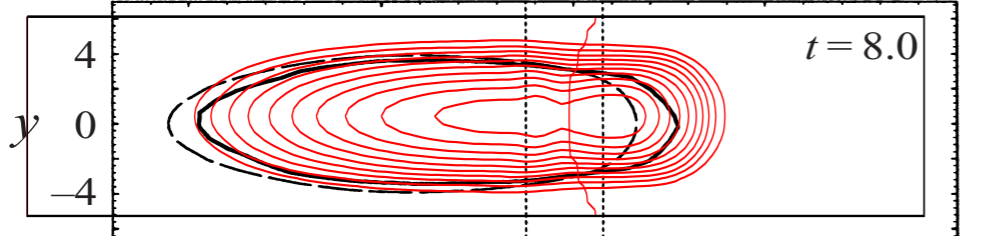
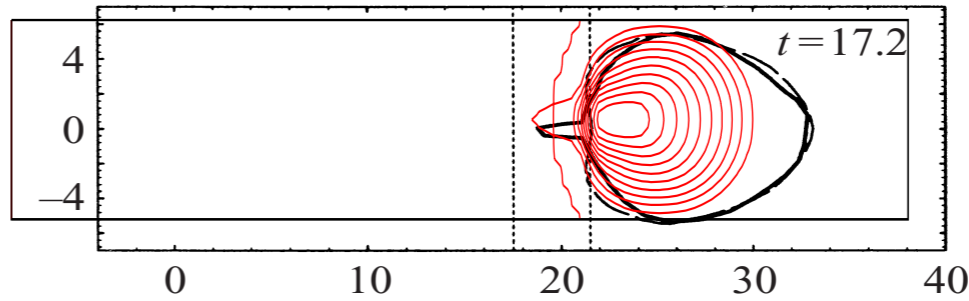
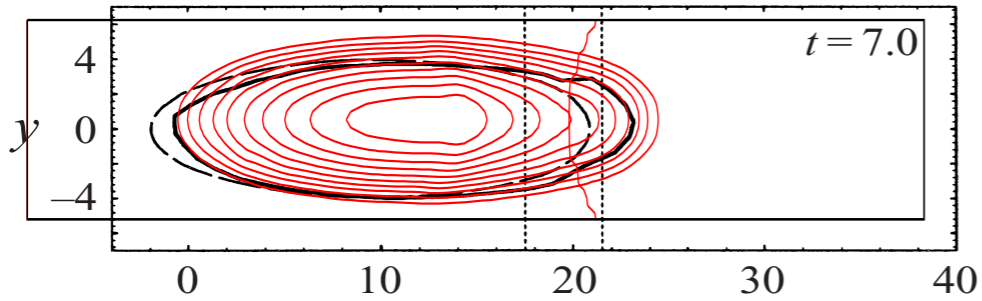
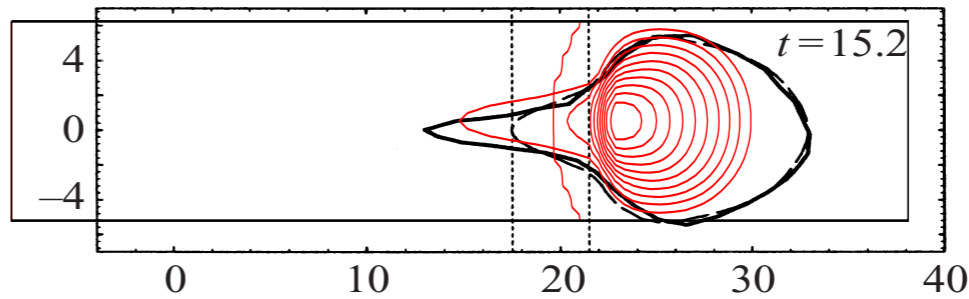
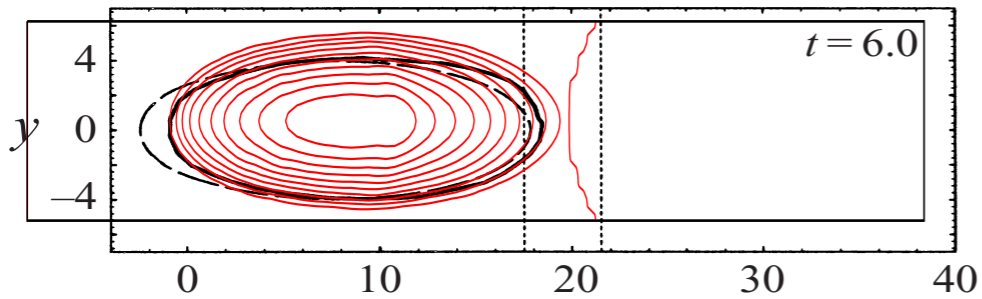
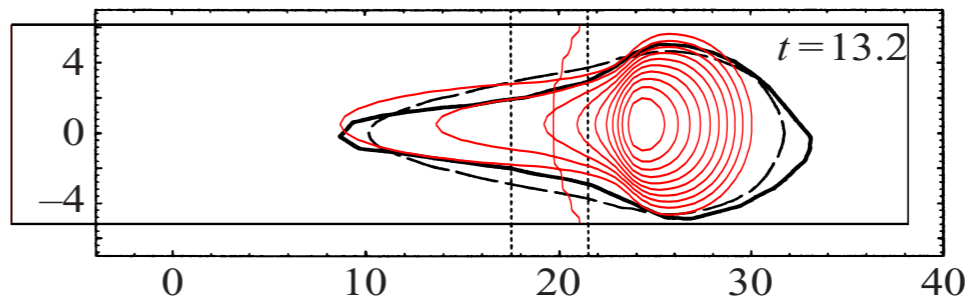
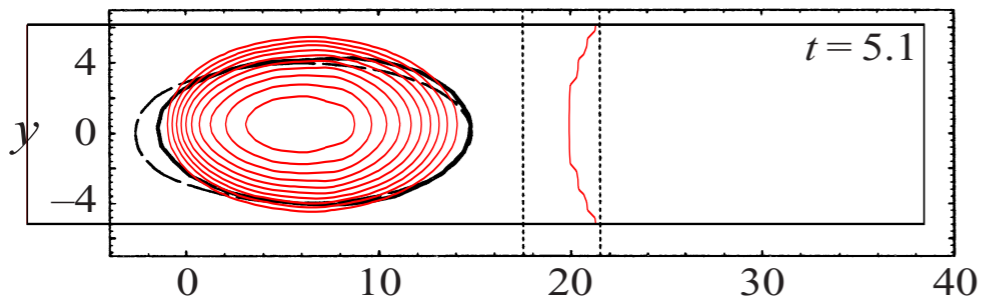
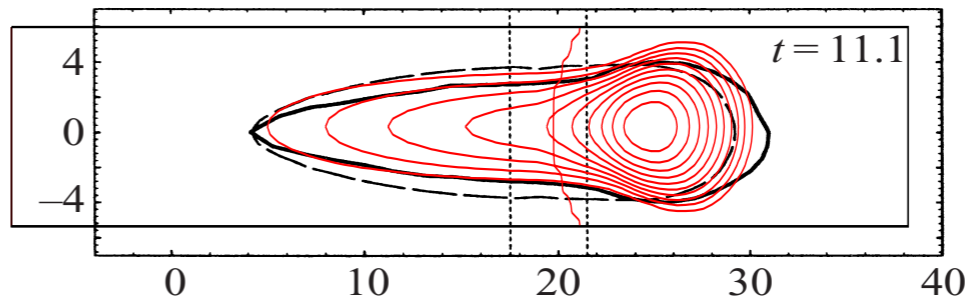
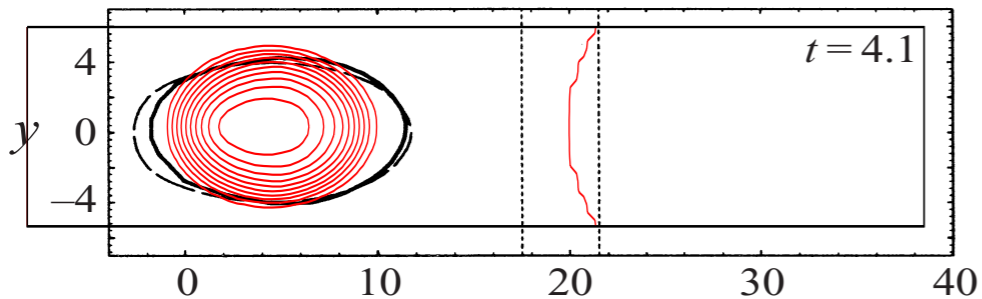
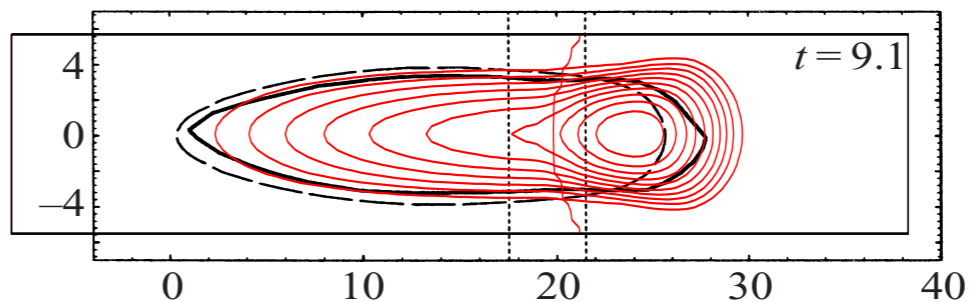
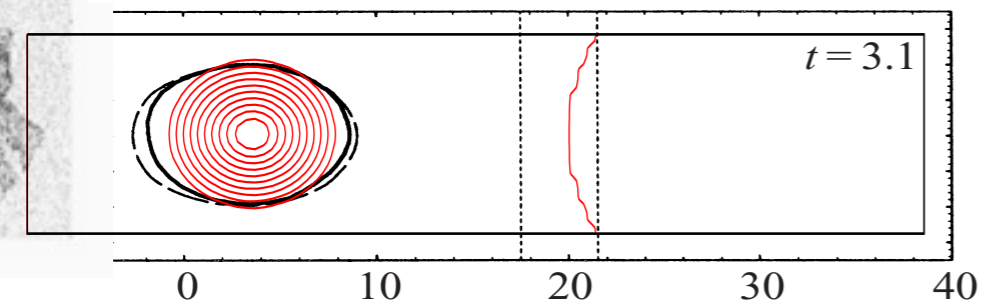


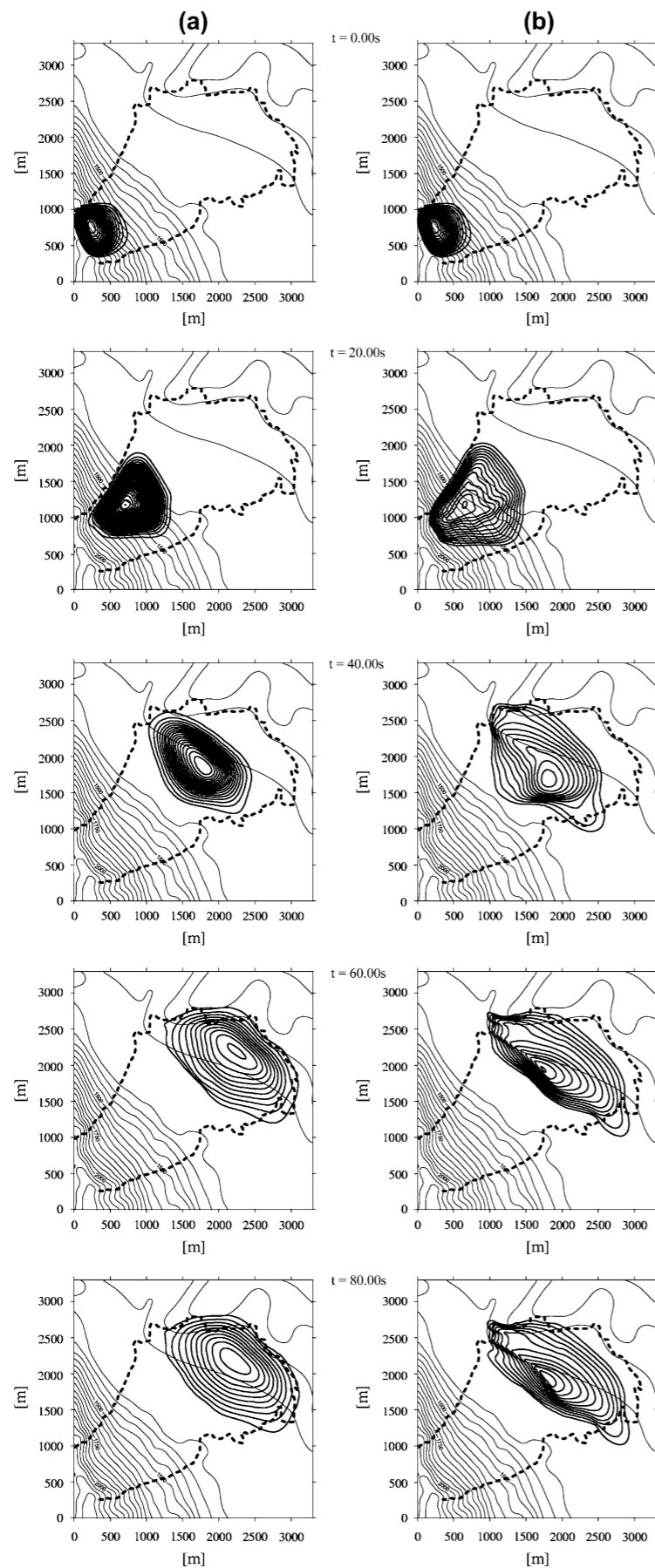
Saint-Venant Savage Hutter *Gerris*



Saint-Venant Savage Hutter *Gerris*

WIELAND, J. M. N. T. GRAY AND K. HUTTER 1999
Marina Pirulli, Marie-Odile Bristeau Anne Mangeney, Claudio Scavia 2006





$H \times L = 1000\text{m} \times 2500\text{m}$

150m initial 16 m for the train



Fig. 24. Analysis of the Frank slide with RASH^{3D}. Plan of the simulated flow position at 20 s intervals (0 s–20 s–40 s–60 s–80 s) in condition of (a) anisotropy and (b) isotropy of normal stresses. The flow depth contours are at 3 m intervals. The sliding surface contours are at 50 m intervals. The dashed line indicated the extent of the real event.



E10001

LUCAS ET AL.: MARTIAN LANDSLIDES SCAR AND DYNAMICS

E10001

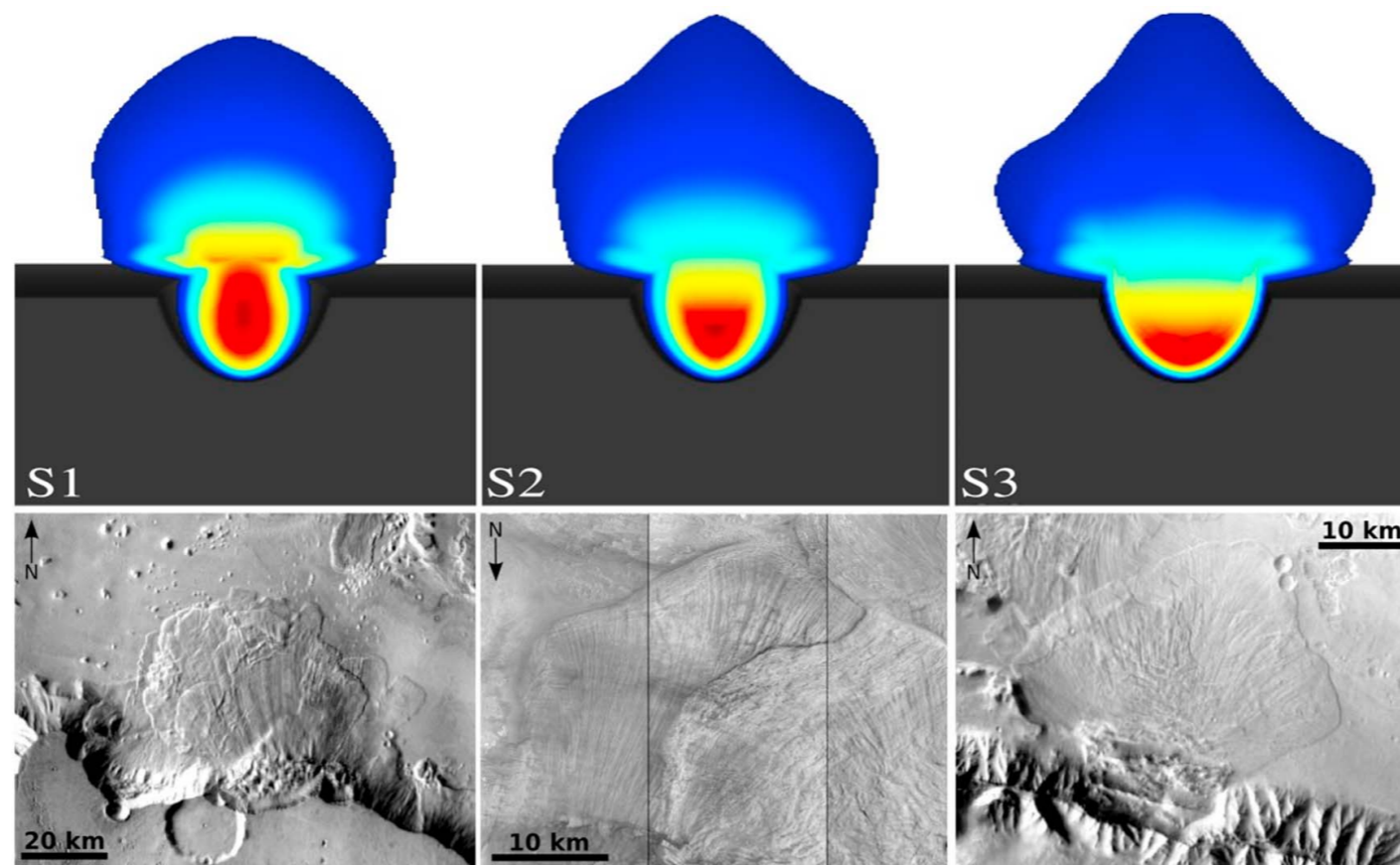
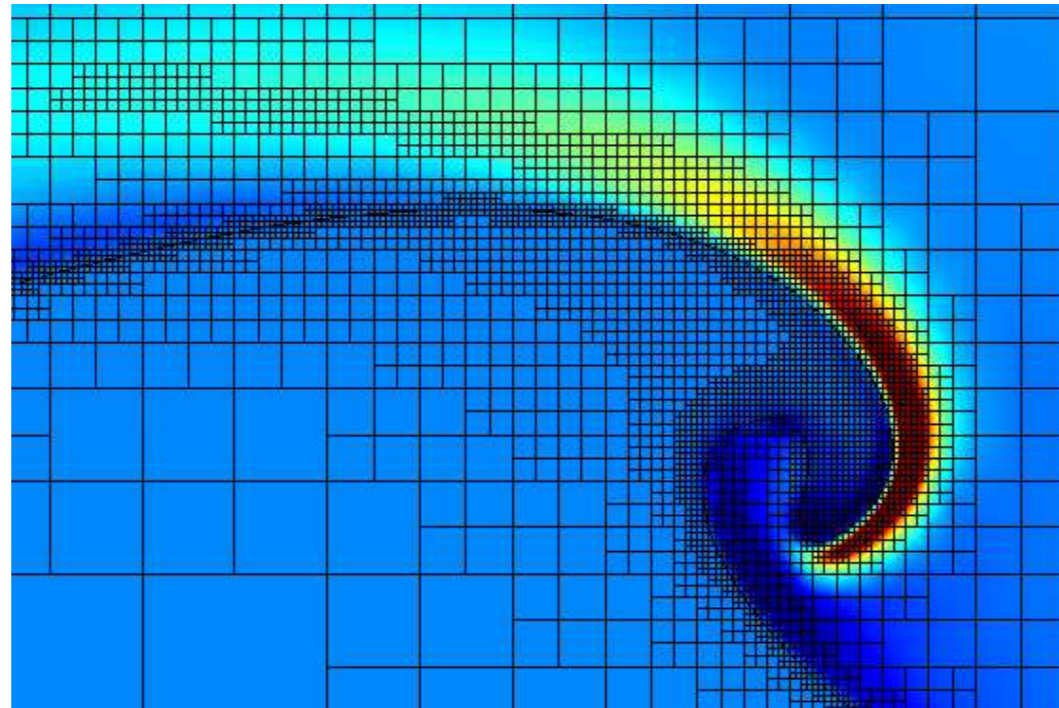
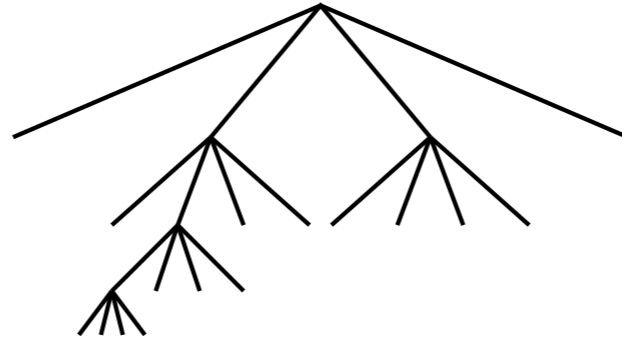
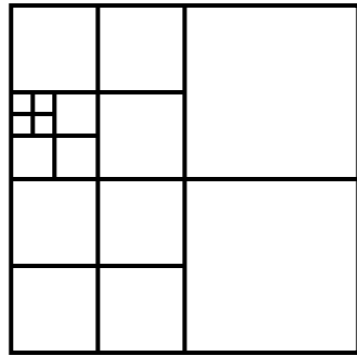


Figure 6. (top) Deposits of simulated landslides obtained for each scar geometry S_i using $\delta = 10^\circ$. (bottom) Martian landslides observed on THEMIS IR present similar deposit shapes, respectively, from left to right: Ganges Chasma, East Ius Chasma, and Coprates Chasma.



outline

- what is a granular fluid? some images
- the $\mu(I)$ friction law obtained from experiments and discrete simulation
- the viscosity associated to the $\mu(I)$ friction law
- the Saint Venant Savage Hutter Hyperbolic model
- implementing the $\mu(I)$ friction law in Navier Stokes
- Examples of flows: focusing on the granular column collapse (limits of Saint Venant Savage Hutter Hyperbolic model)



- *Gerris* is a finite volume code by Stéphane Popinet NIWA
- one part of the code is a Navier Stokes solver
- automatic mesh adaptation
- Volume Of Fluid method for two phase flows
- free on sourceforge



rheology; defining a viscosity

$$\mu(I) = \mu_1 + \frac{\mu_2 - \mu_1}{I_0/I + 1}$$

$$\eta \frac{\partial u}{\partial y} = \mu(I) P$$

local equilibrium

$$\eta = \frac{\mu\left(\frac{d\frac{\partial u}{\partial y}}{\sqrt{P/\rho}}\right) P}{\frac{\partial u}{\partial y}}$$

construction of a viscosity



implementation in *Gerris* flow solver?

$$\mu(I) = \mu_1 + \frac{\mu_2 - \mu_1}{I_0/I + 1}$$

$$D_2 = \sqrt{D_{ij}D_{ij}} \quad D_{ij} = \frac{u_{i,j} + u_{j,i}}{2}$$

construction of a viscosity based on the D_2 invariant and redefinition of I

$$\eta = \min(\eta_{max}, \max\left(\frac{\mu(I)}{\sqrt{2}D_2}p, 0\right)) \quad I = d\sqrt{2}D_2 / \sqrt{(|p|/\rho)}.$$

- the «min» limits viscosity to a large value
- always flow, even slow

Boundary Conditions: no slip and $P=0$ at the interface



implementation in *Gerris* flow solver?

$$\mu(I) = \mu_1 + \frac{\mu_2 - \mu_1}{I_0/I + 1}$$

$$D_2 = \sqrt{D_{ij}D_{ij}} \quad D_{ij} = \frac{u_{i,j} + u_{j,i}}{2}$$

construction of a viscosity based on the D_2 invariant and redefinition of I

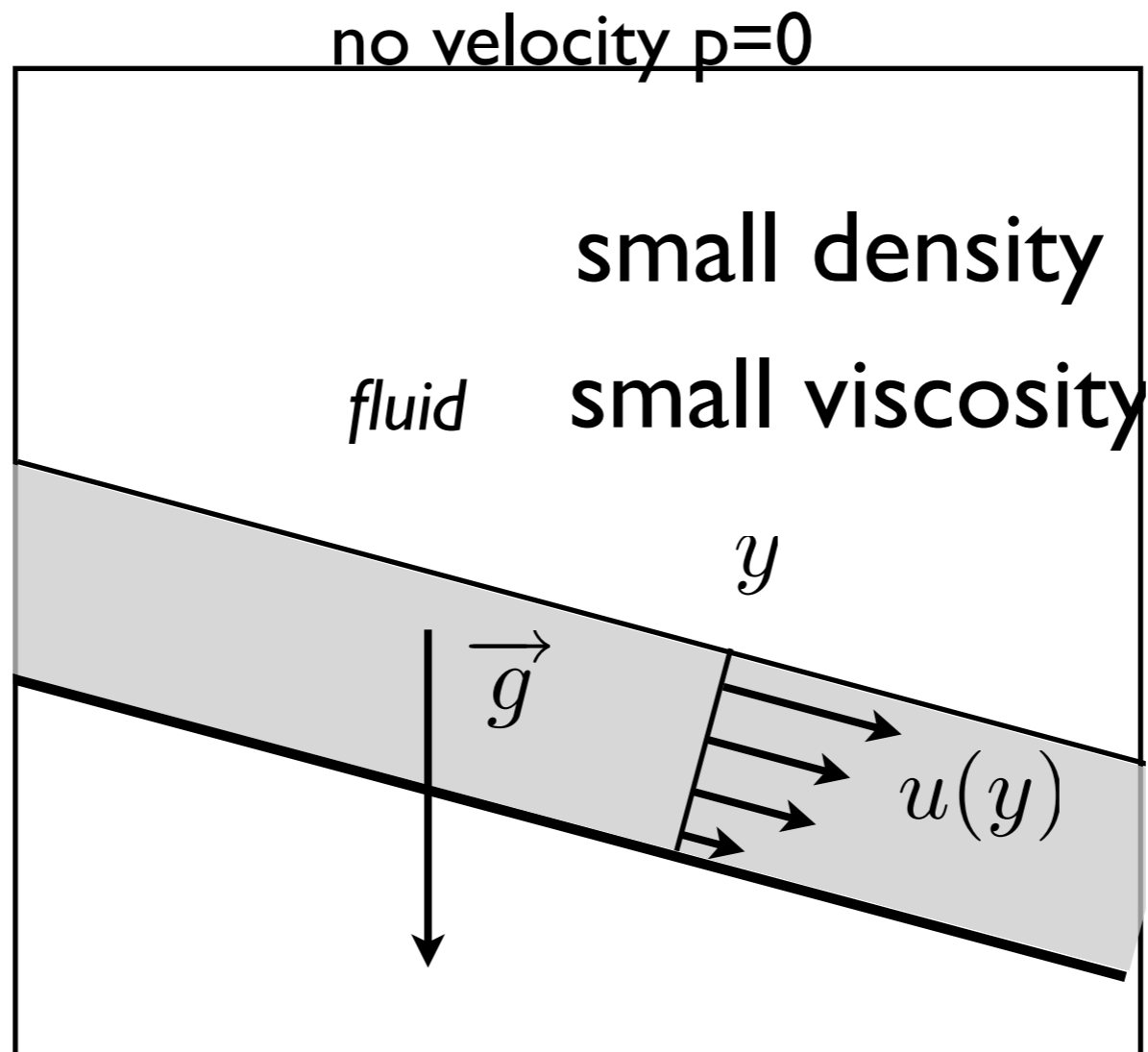
$$\eta = \min(\eta_{max}, \max\left(\frac{\mu(I)}{\sqrt{2}D_2}p, 0\right)) \quad I = d\sqrt{2}D_2 / \sqrt{(|p|/\rho)}.$$

$$\nabla \cdot \mathbf{u} = 0, \quad \rho \left(\frac{\partial \mathbf{u}}{\partial t} + \mathbf{u} \cdot \nabla \mathbf{u} \right) = -\nabla p + \nabla \cdot (2\eta \mathbf{D}) + \rho g,$$

$$\frac{\partial c}{\partial t} + \nabla \cdot (c\mathbf{u}) = 0, \quad \rho = c\rho_1 + (1-c)\rho_2, \quad \eta = c\eta_1 + (1-c)\eta_2$$

The granular fluid is covered by a passive light fluid (it allows for a zero pressure boundary condition at the surface, bypassing an up to now difficulty which was to impose this condition on a unknown moving boundary).

Boundary Conditions: no slip and $P=0$ at the top

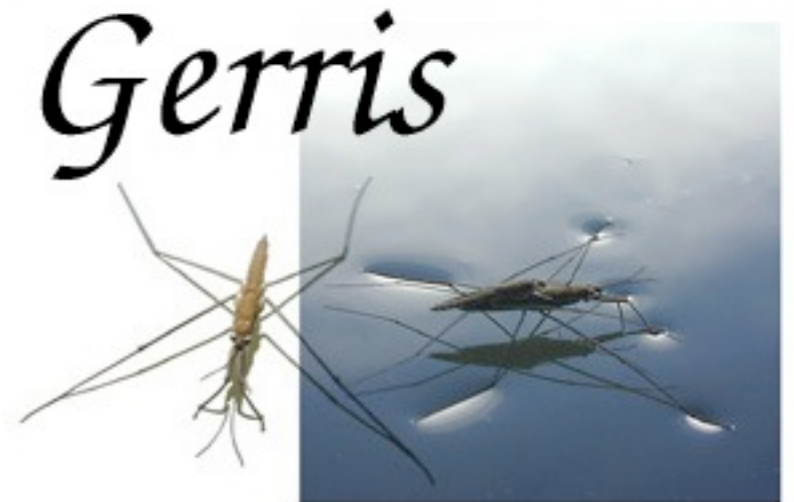


$$\frac{\partial c}{\partial t} + \nabla \cdot (c\mathbf{u}) = 0, \quad \rho = c\rho_1 + (1 - c)\rho_2, \quad \eta = c\eta_1 + (1 - c)\eta_2$$

The granular fluid is covered by a passive light fluid (it allows for a zero pressure boundary condition at the surface, bypassing an up to now difficulty which was to impose this condition on a unknown moving boundary).

Boundary Conditions: no slip and $P=0$ at the top

Projection Method



Gerris

$$\begin{aligned}\rho_{n+\frac{1}{2}} \left(\frac{\mathbf{u}_* - \mathbf{u}_n}{\Delta t} + \mathbf{u}_{n+\frac{1}{2}} \cdot \nabla \mathbf{u}_{n+\frac{1}{2}} \right) &= \nabla \cdot (\eta_{n+\frac{1}{2}} \mathbf{D}_*) - \nabla p_{n-\frac{1}{2}}, \\ \mathbf{u}_{n+1} &= \mathbf{u}_* - \frac{\Delta t}{\rho_{n+\frac{1}{2}}} (\nabla p_{n+\frac{1}{2}} - \nabla p_{n-\frac{1}{2}}), \\ \nabla \cdot \mathbf{u}_{n+1} &= 0.\end{aligned}$$

multigrid solver for Laplacien of pressure

$$\nabla \cdot \left(\frac{\Delta t}{\rho_{n+\frac{1}{2}}} \nabla p_{n+\frac{1}{2}} \right) = \nabla \cdot \left(\mathbf{u}_* + \frac{\Delta t}{\rho_{n+\frac{1}{2}}} \nabla p_{n-\frac{1}{2}} \right)$$

implicit for \mathbf{u}^*

$$\frac{\rho_{n+\frac{1}{2}}}{\Delta t} \mathbf{u}_* - \frac{1}{2} \nabla \cdot (\eta_{n+\frac{1}{2}} \nabla \mathbf{u}_*) = \rho_{n+\frac{1}{2}} \left[\frac{\mathbf{u}_n}{\Delta t} - \mathbf{u}_{n+\frac{1}{2}} \cdot \nabla \mathbf{u}_{n+\frac{1}{2}} \right] - \nabla p_{n-\frac{1}{2}} + \frac{1}{2} \nabla \mathbf{u}_n^T \nabla \eta_{n+\frac{1}{2}}.$$

VOF reconstruction

$$\frac{c_{n+\frac{1}{2}} - c_{n-\frac{1}{2}}}{\Delta t} + \nabla \cdot (c_n \mathbf{u}_n) = 0$$

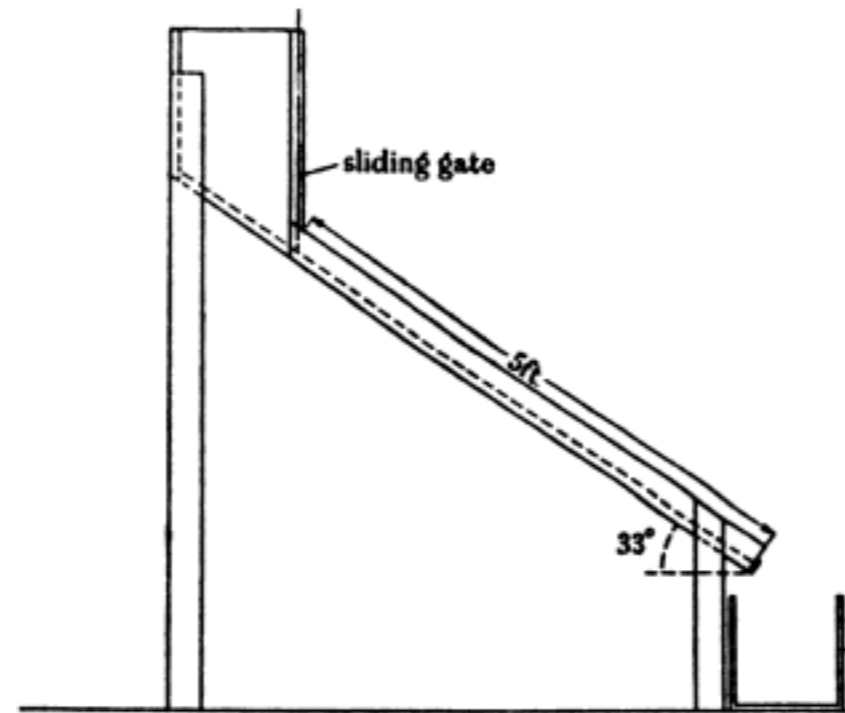


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Test of the code: «Bagnold» avalanche



kind of Nußelt solution

$$T \propto \sigma(\lambda D)^2 (dU/dy)^2$$

$$U = \frac{2}{3} \times 0.165 (g \sin \beta)^{1/2} \frac{y^{3/2}}{D}$$

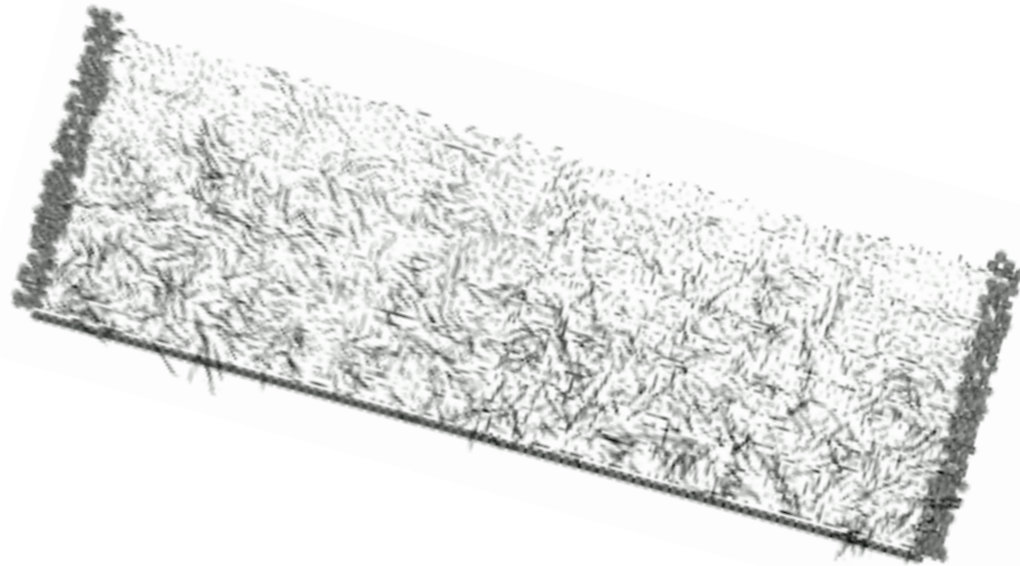
TABLE 1.

flow height Y (cm)	measured speed (cm/sec)	speed, from (9) (cm/sec)	ratio
0.5	17.2	26.4	1.53
0.65	27.5	38.8	1.41
0.75	30.0	48.0	1.6
0.9	39.0	63.0	1.61

Bagnold 1954



Test of the code: «Bagnold» avalanche

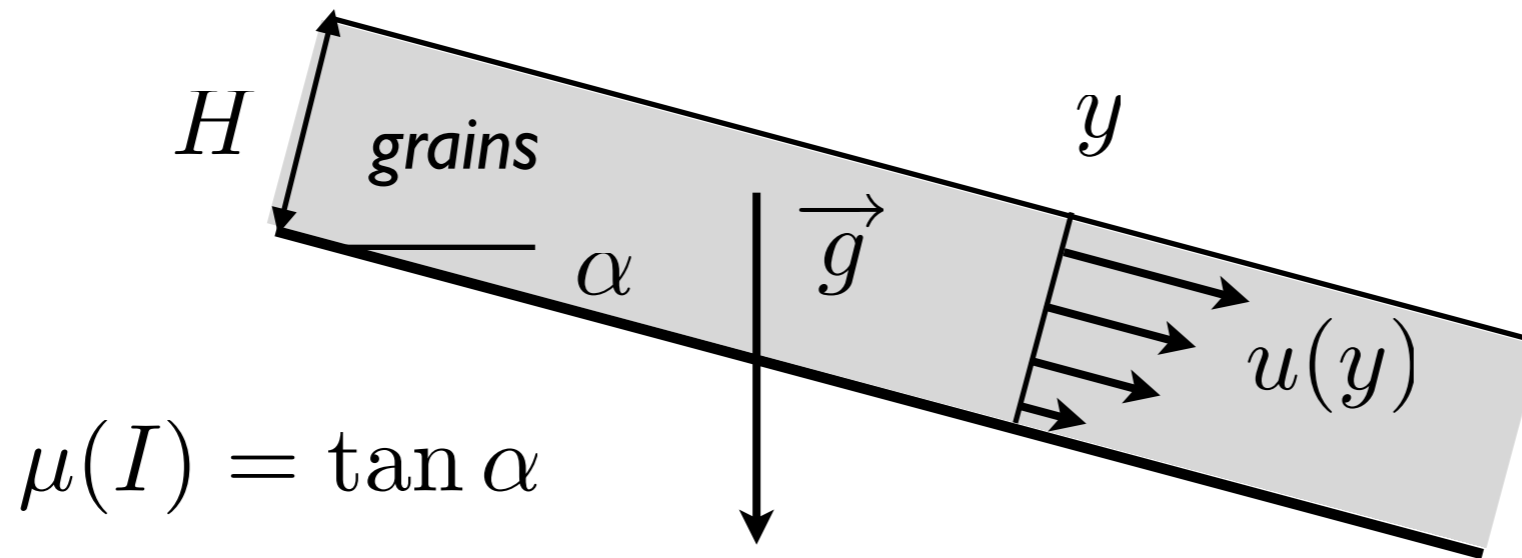


kind of Nußelt solution

Contact Dynamic
simulation Lydie Staron

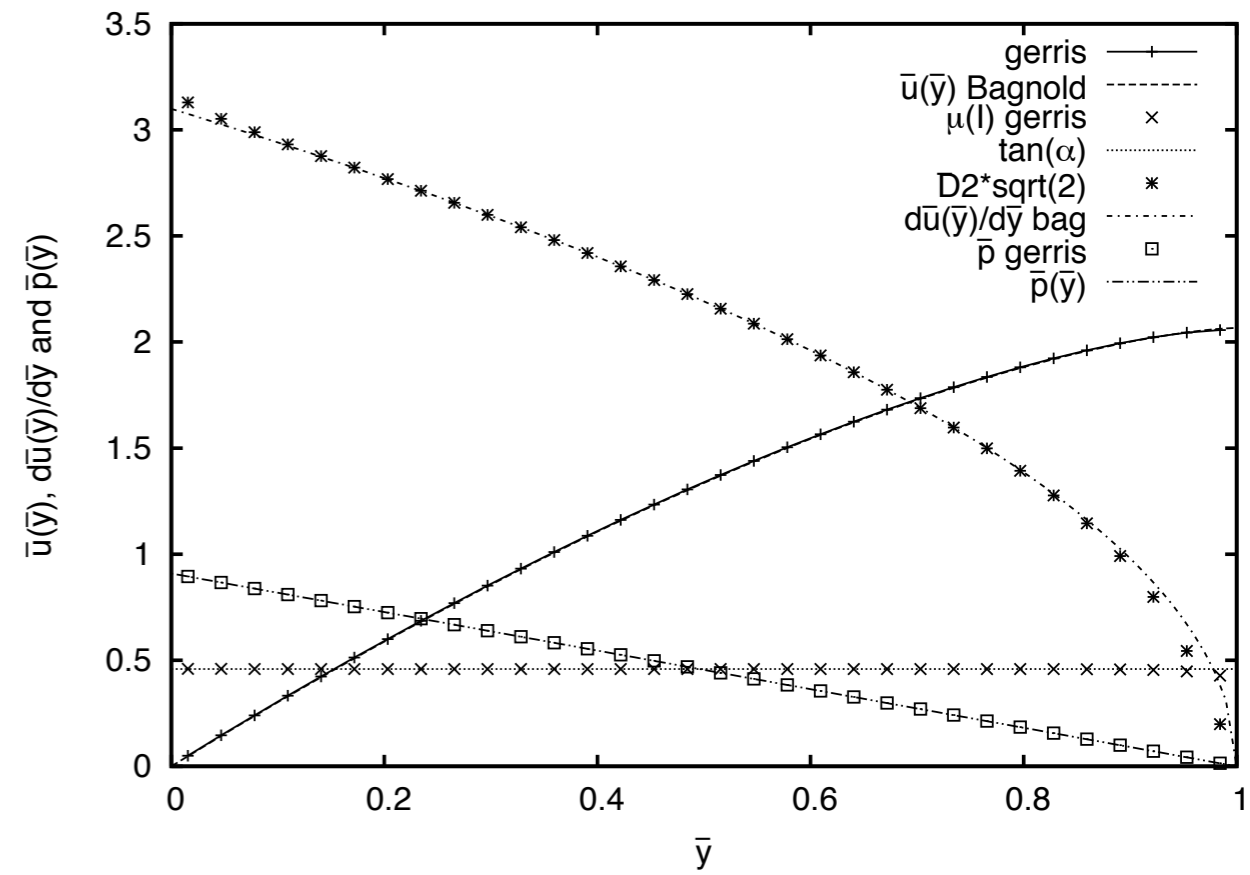


Test of the code: «Bagnold» avalanche



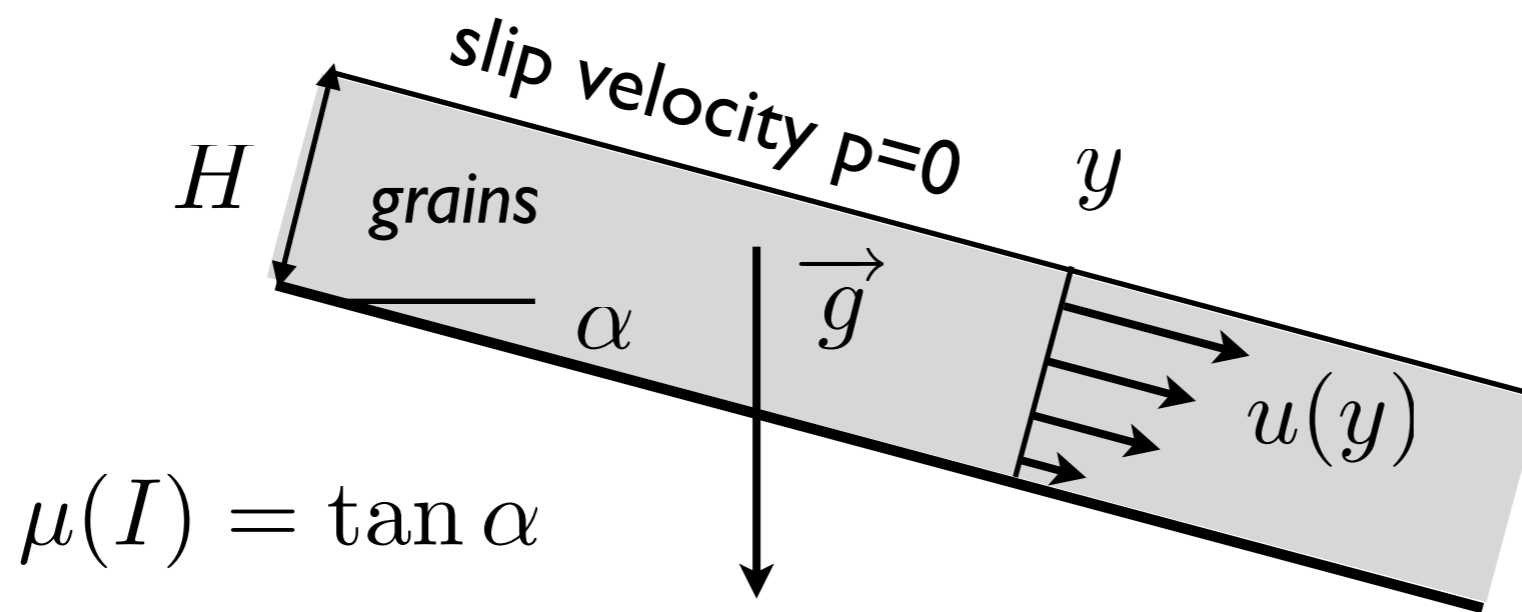
$$u = \frac{2}{3} I_{\alpha} \sqrt{gd \cos \alpha} \frac{H^3}{d^3} \left(1 - \left(1 - \frac{y}{H} \right)^{3/2} \right),$$

$$v = 0, \quad p = \rho g H \left(1 - \frac{y}{H} \right) \cos \alpha.$$



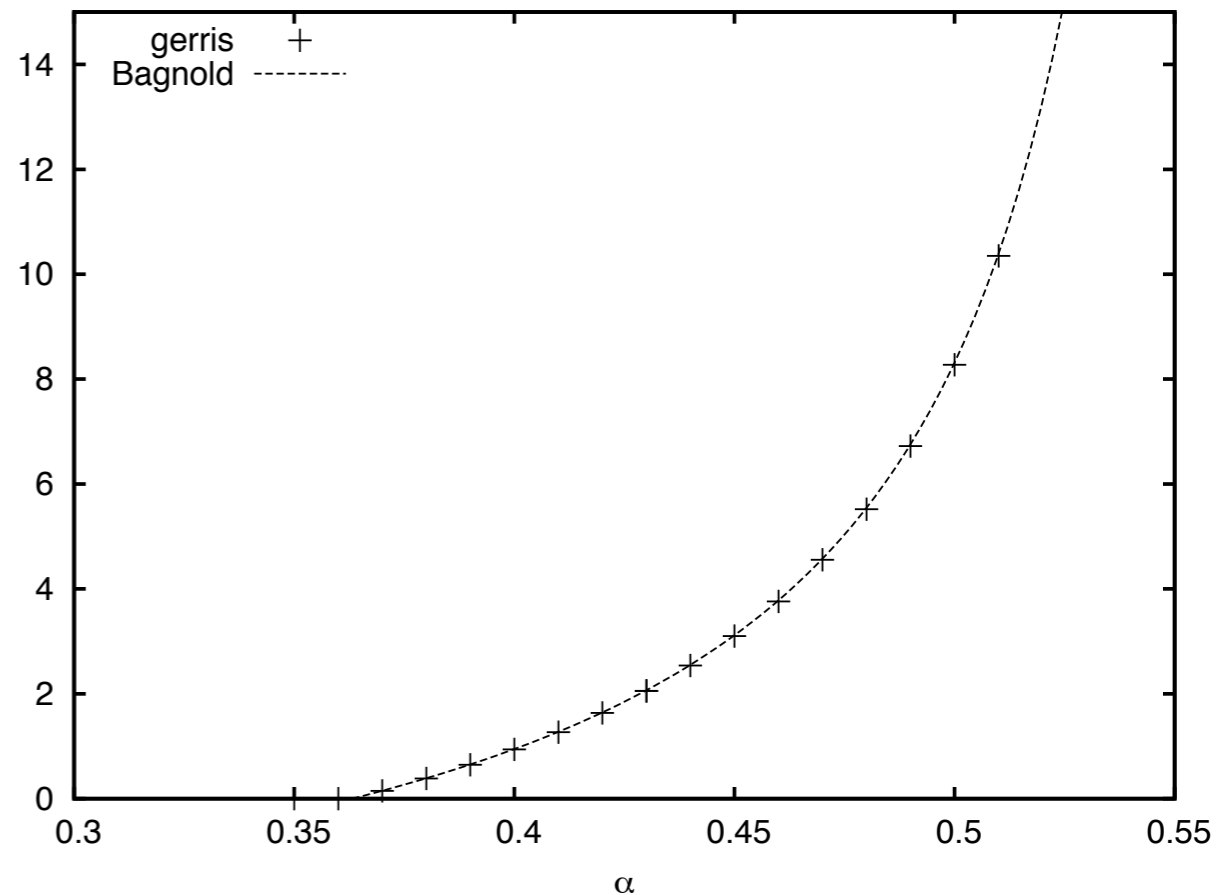


Test of the code: «Bagnold» avalanche



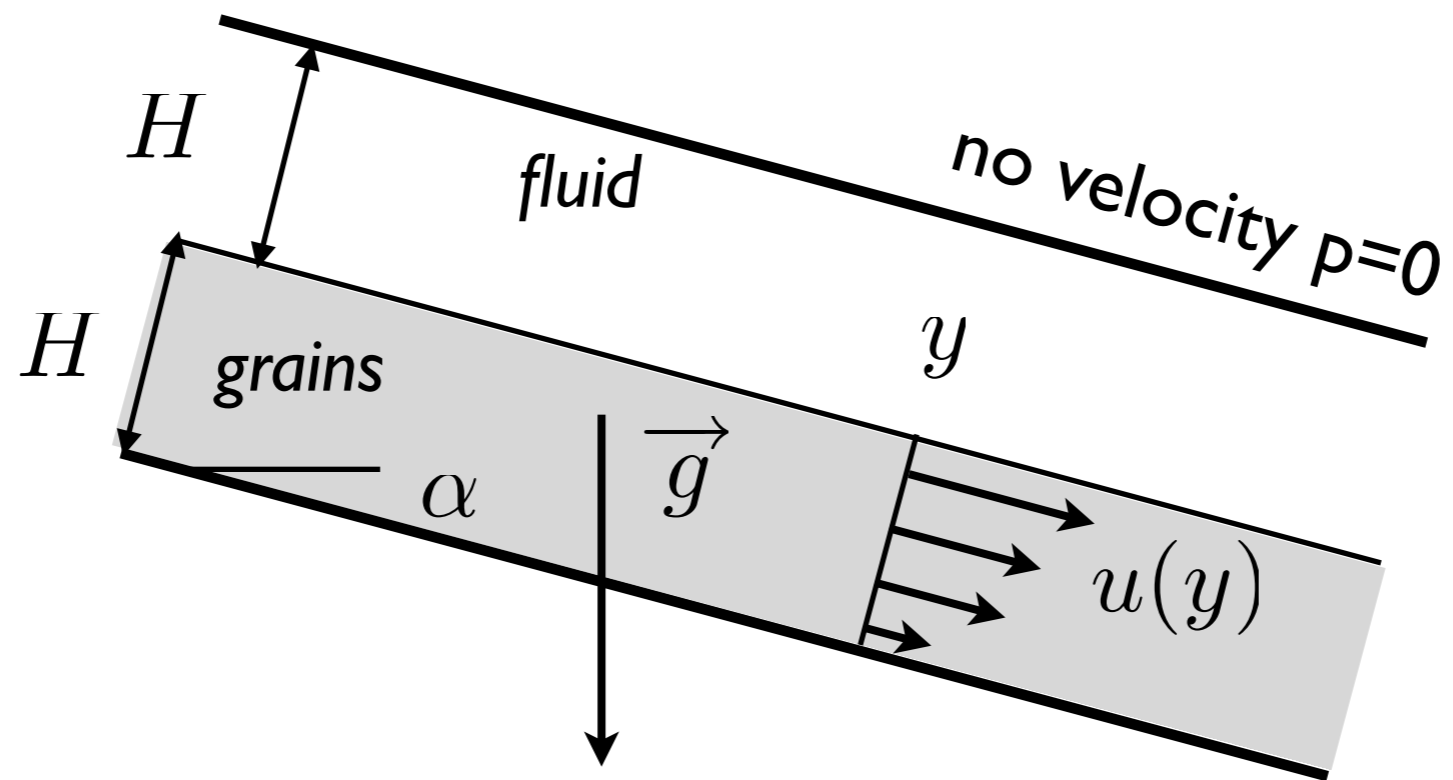
$$u = \frac{2}{3} I_{\alpha} \sqrt{gd \cos \alpha} \frac{H^3}{d^3} \left(1 - \left(1 - \frac{y}{H} \right)^{3/2} \right), \quad \bar{u}(1)$$

$$v = 0, \quad p = \rho g H \left(1 - \frac{y}{H} \right) \cos \alpha.$$





Test of the code: «Bagnold» avalanche



$$u(y) = \frac{(2H - y) (\rho_f g y \sin \alpha - 2\tau_0)}{2\mu_f}$$

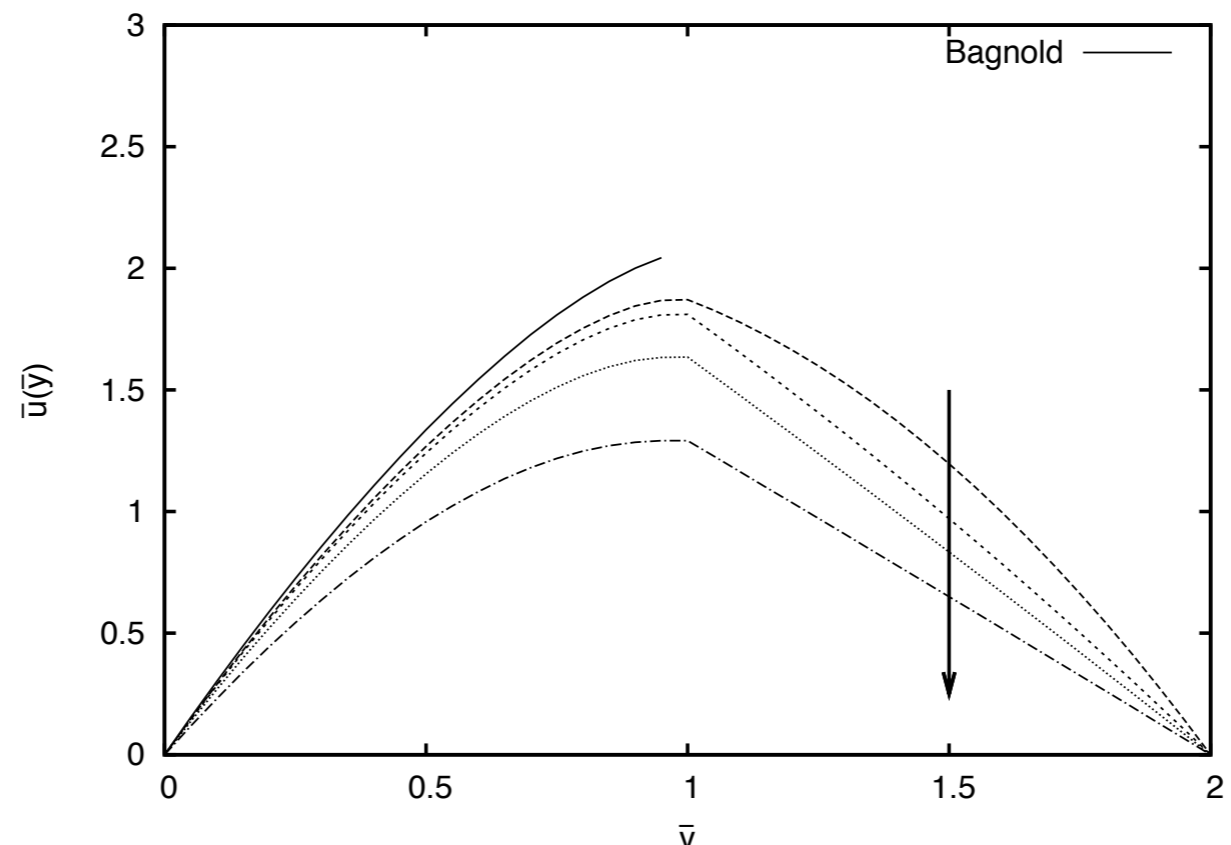
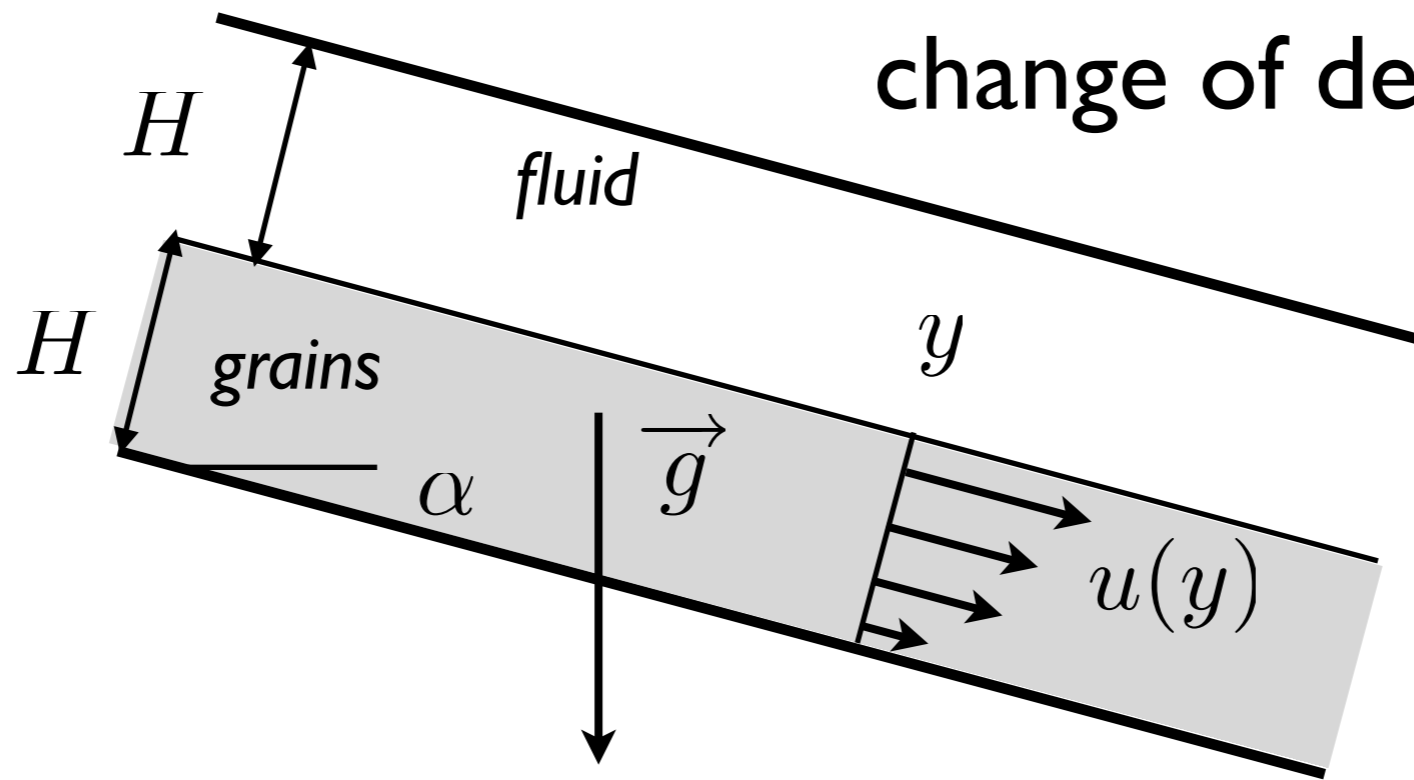
$$d\frac{\partial u}{\partial y} = \max \left[\sqrt{p_0/\rho + gH \left(1 - \frac{y}{H}\right) \cos \alpha} \times \mu^{-1} \left(\frac{\tau_0 + \rho g H \left(1 - \frac{y}{H}\right) \sin \alpha}{p_0 + \rho g H \left(1 - \frac{y}{H}\right) \cos \alpha} \right), 0 \right].$$

$$u(H^-) - u(H^+) = 0$$



Test of the code: «Bagnold» avalanche

change of density of the fluid

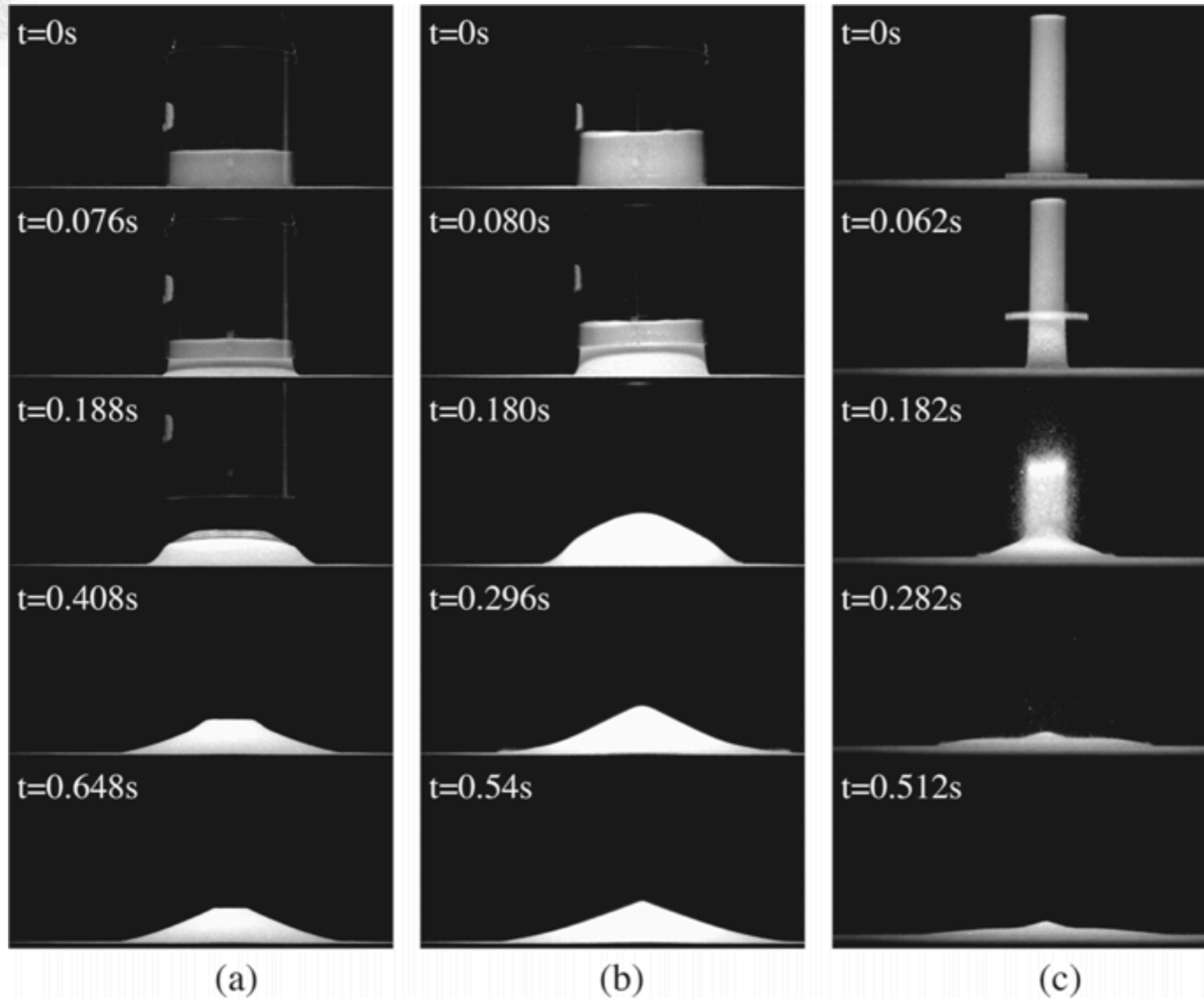




The sand pit problem: quickly remove the bucket of sand

Granular Column Collapse

E. Lajeunesse A. Mangeney-Castelnau and J. P. Vilotte PoF 204

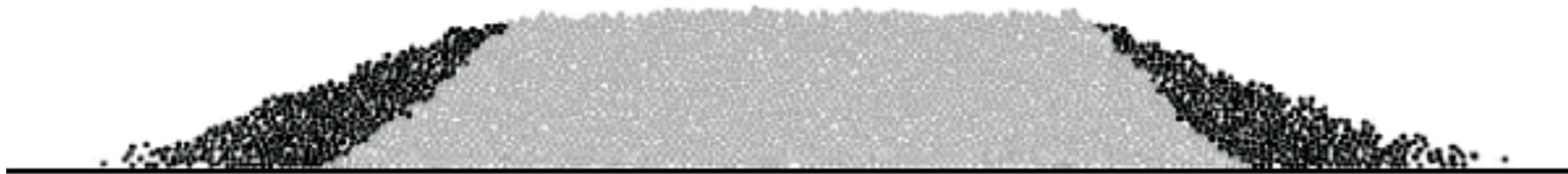


The sand pit problem: quickly remove the bucket of sand



Collapse of columns

$a=0.37$



Contact Dynamic
simulation Lydie Staron



Collapse of columns

$a=0.90$



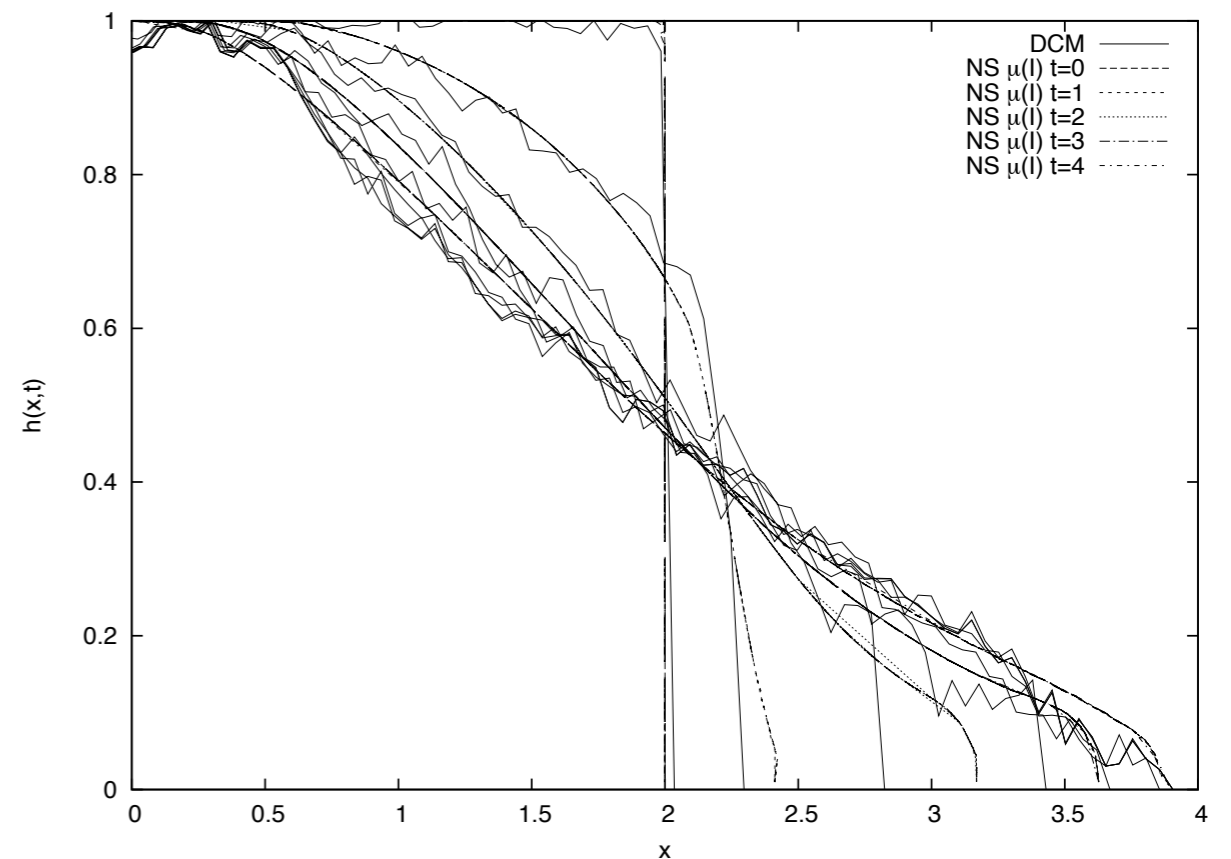
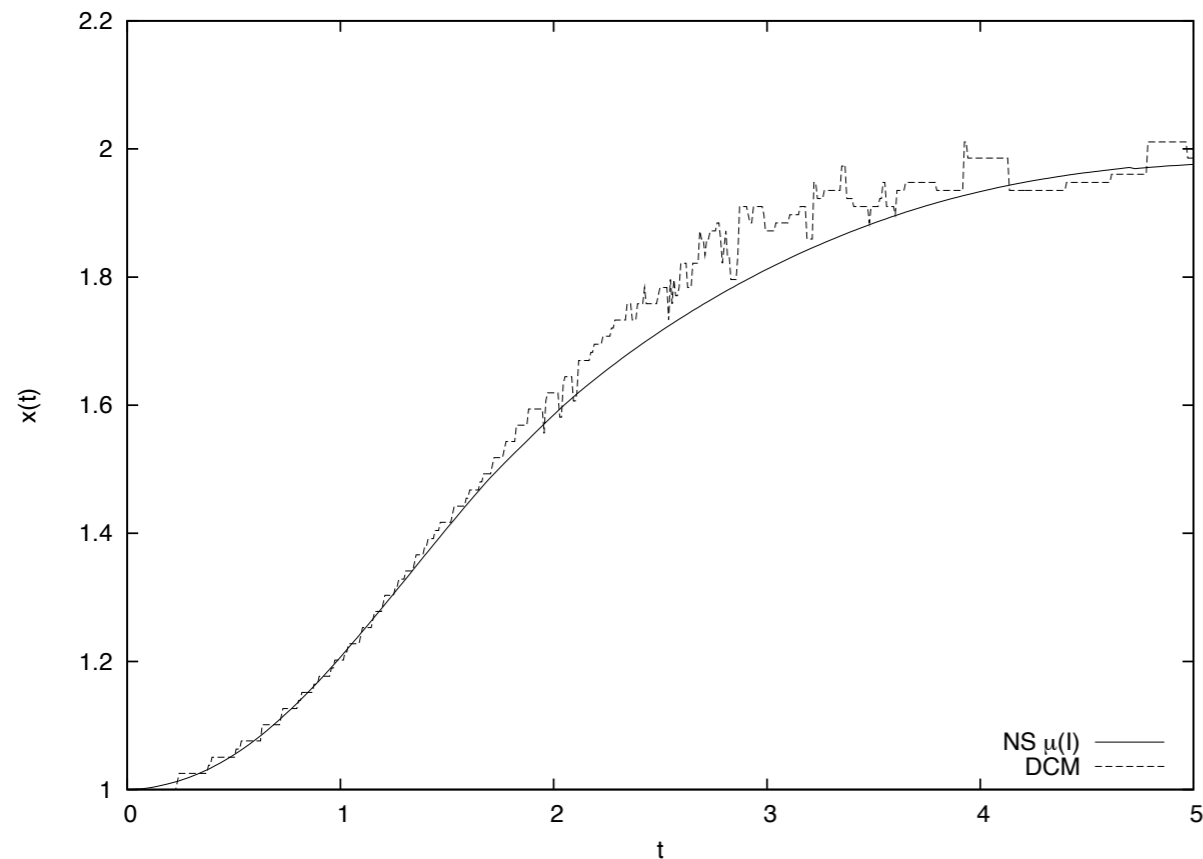
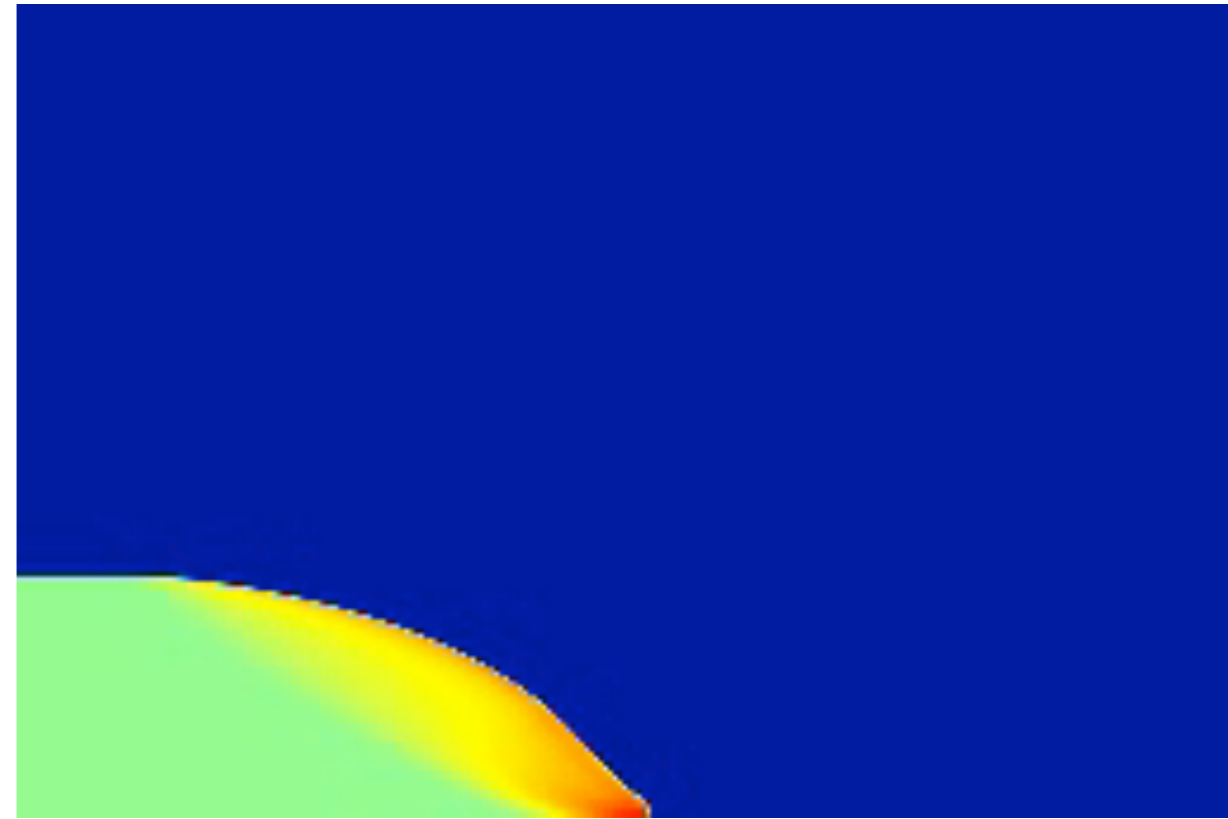
Contact Dynamic
simulation Lydie Staron



Collapse of columns simulation *Gerris* $\mu(l)$



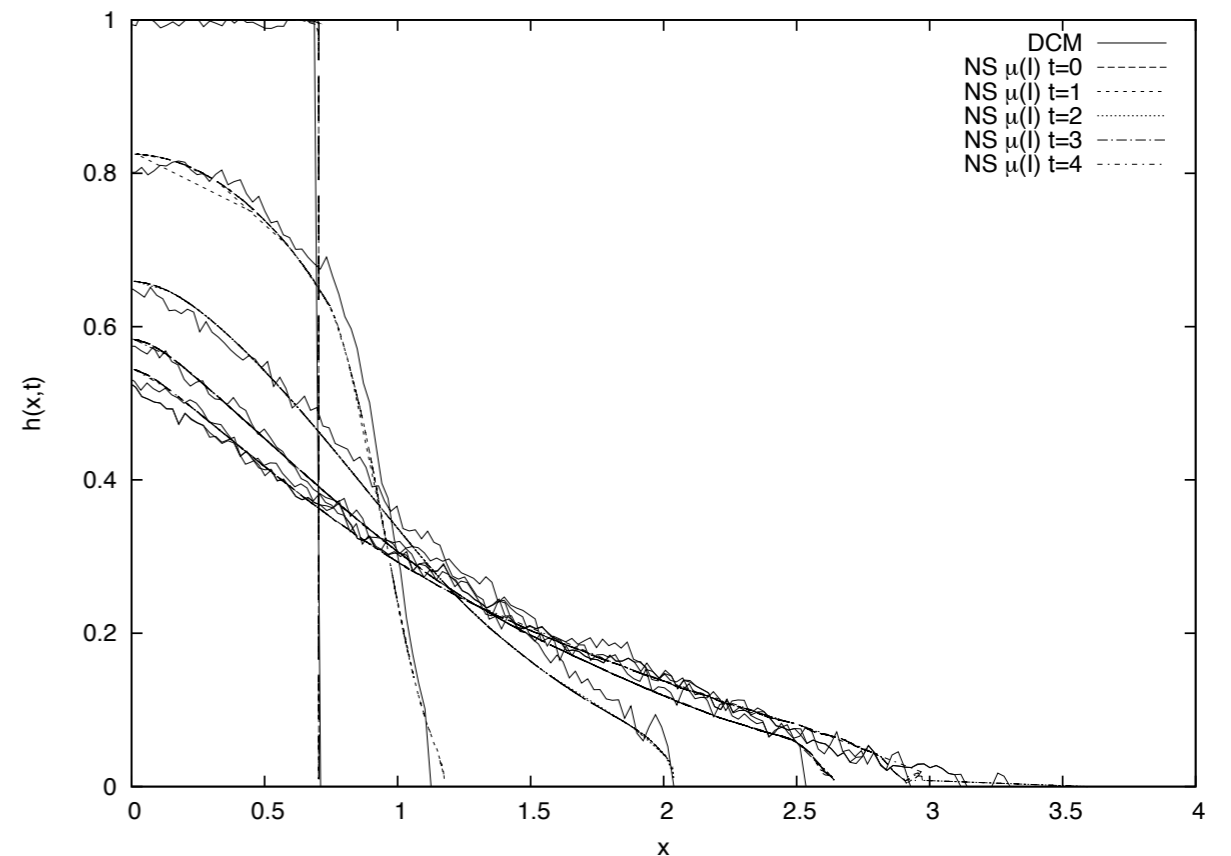
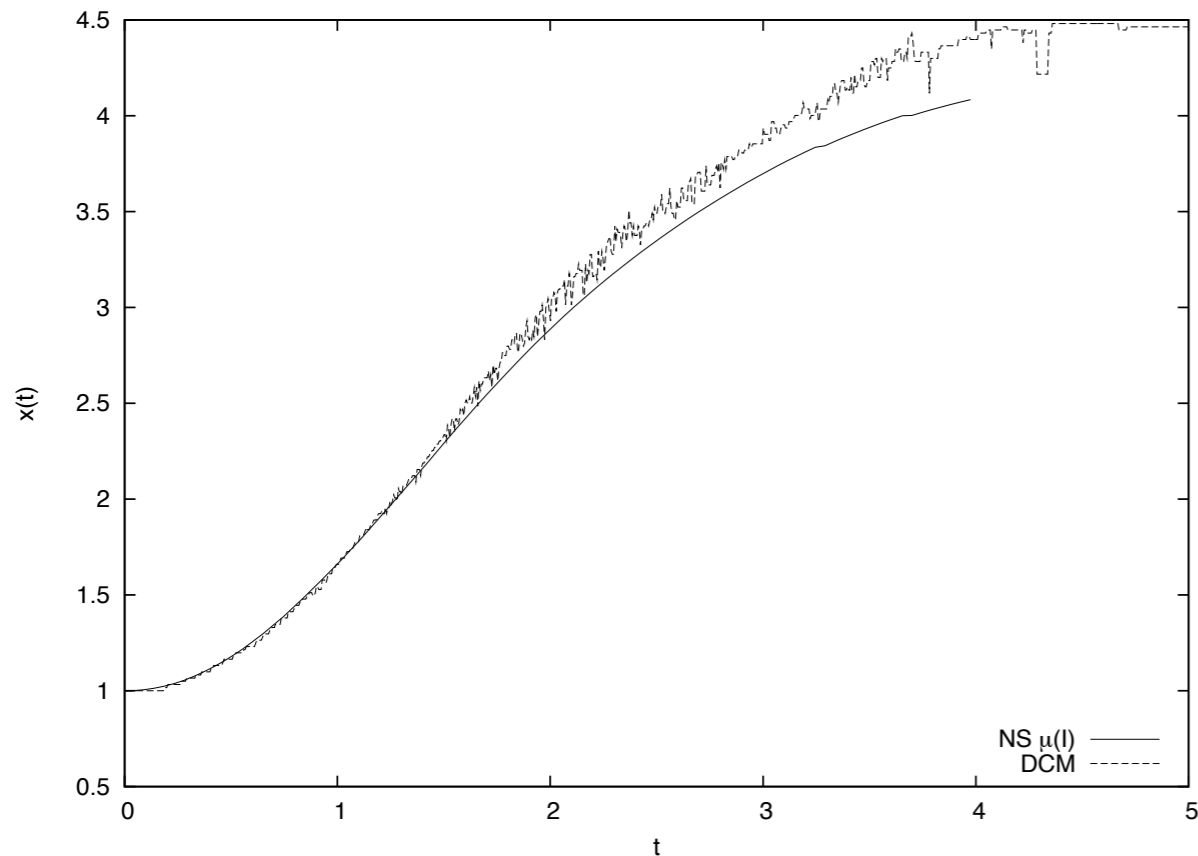
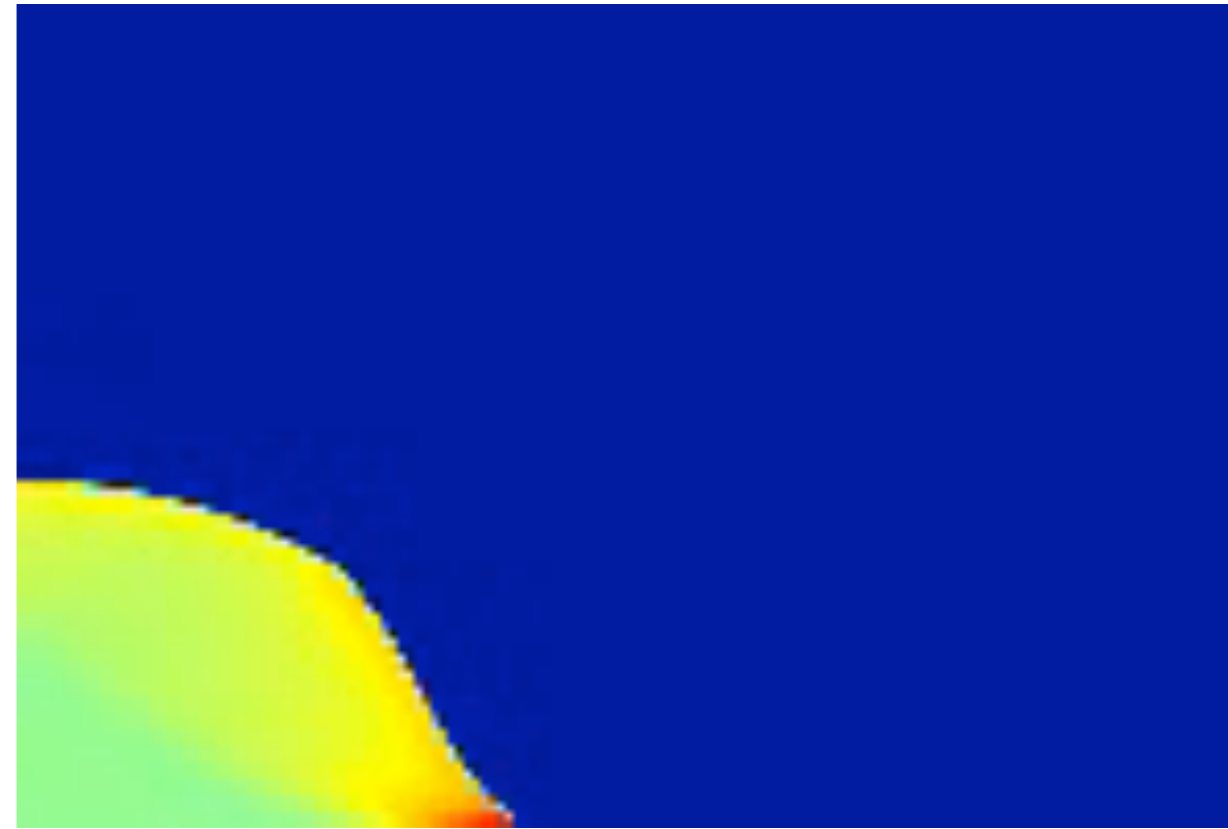
Collapse of columns of aspect ratio 0.5
comparison of Discrete Simulation Contact
Method and Navier Stokes gerris, shape at time
0, 1, 2, 3, 4 and position of the front of the
avalanche as function of time (time measured
with $\sqrt{H_0/g}$ and space with aH_0)



Collapse of columns simulation *Gerris* $\mu(l)$



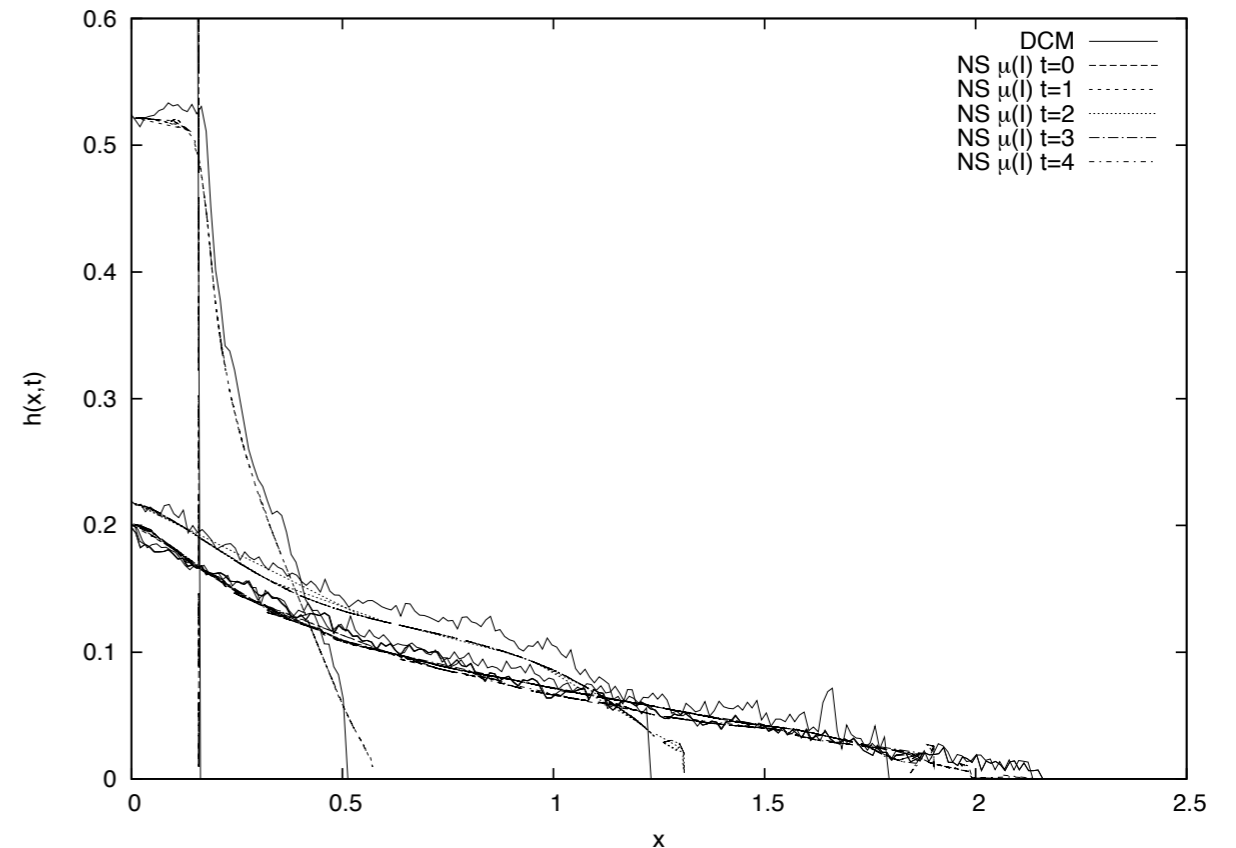
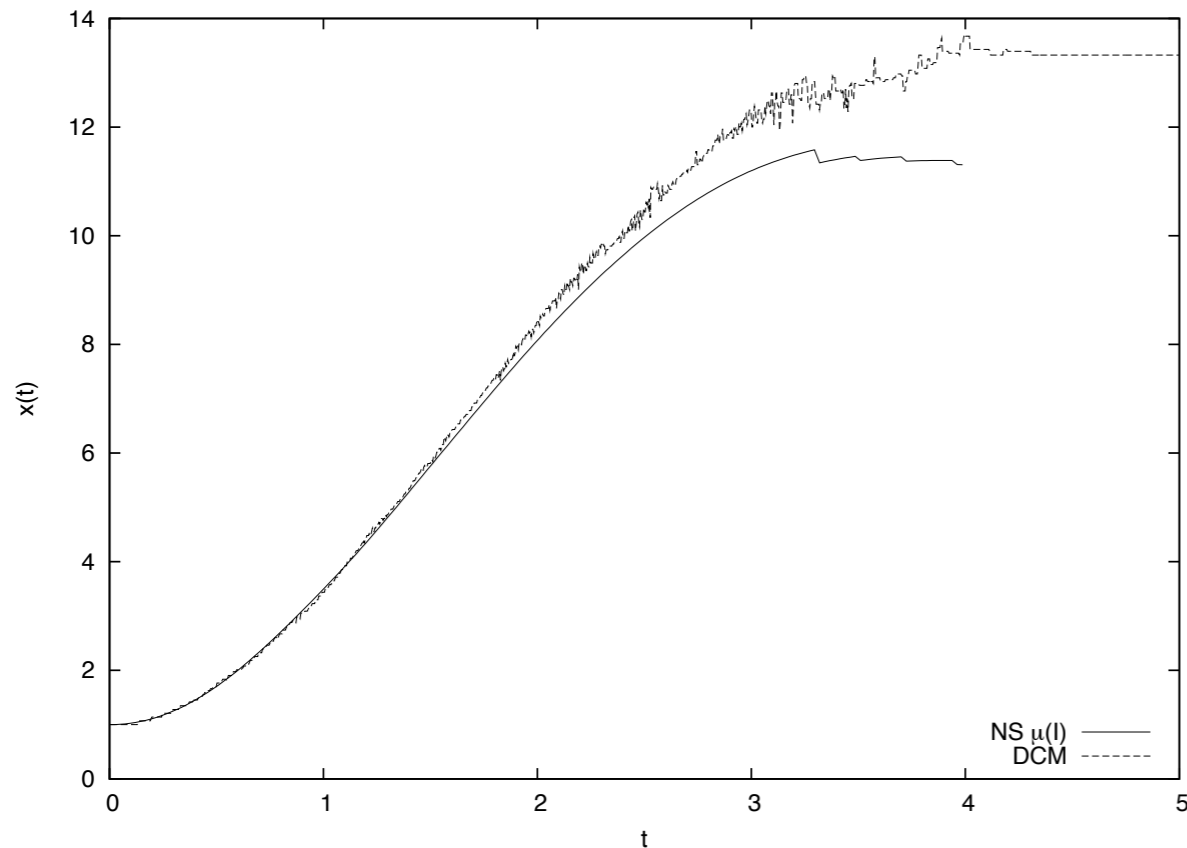
Collapse of columns of aspect ratio 1.42
comparison of Discrete Simulation Contact
Method and Navier Stokes gerris, shape at time
0, 1, 2, 3, 4 and position of the front of the
avalanche as function of time (time measured
with $\sqrt{H_0/g}$ and space with aH_0)



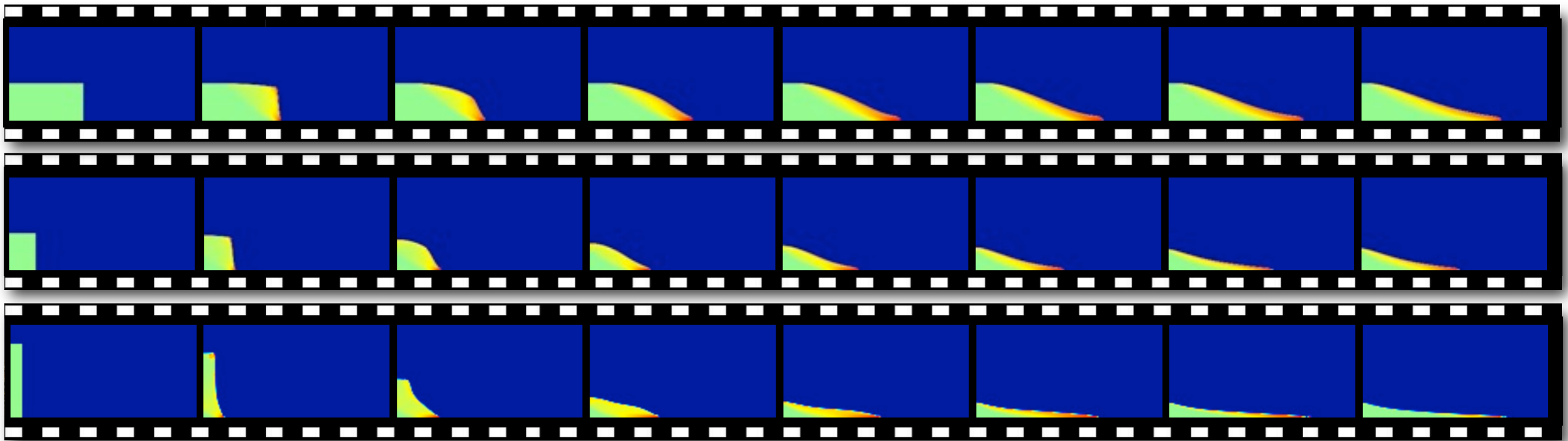
Collapse of columns simulation *Gerris* $\mu(l)$



Collapse of columns of aspect ratio 6.26
comparison of Discrete Simulation Contact
Method and Navier Stokes gerris, shape at time
0, 1, 2, 3, 4 and position of the front of the
avalanche as function of time (time measured
with $\sqrt{H_0/g}$ and space with aH_0)



Collapse of columns simulation *Gerris* $\mu(l)$



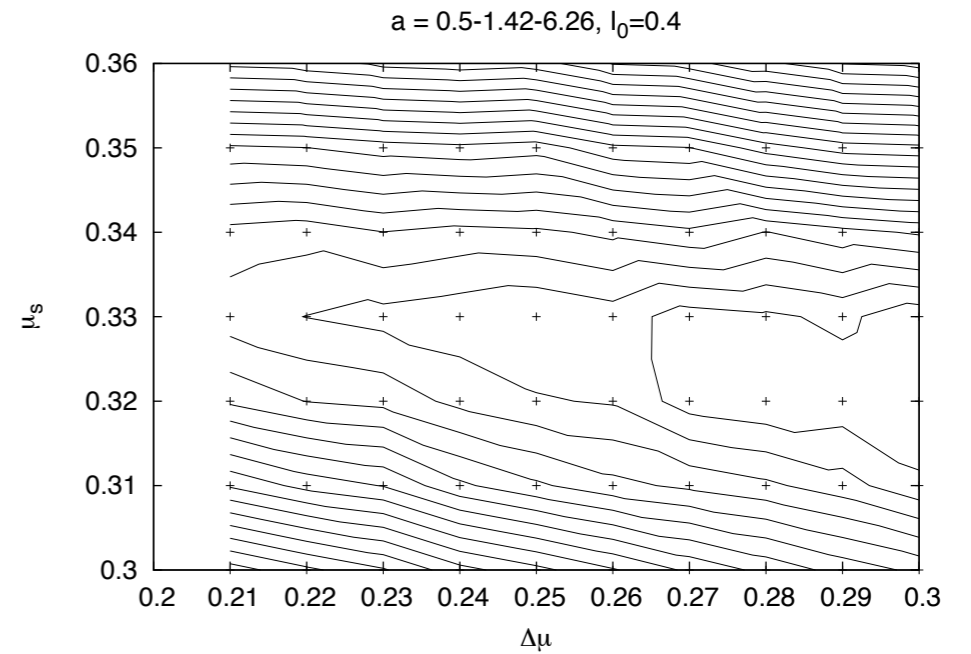
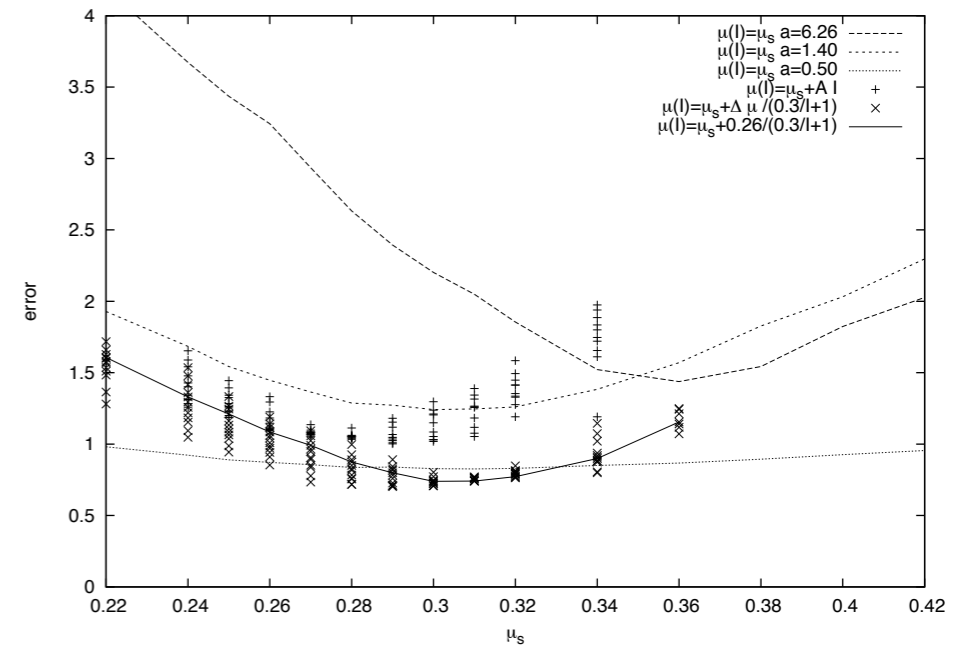
Snapshots of collapse of three columns of aspect ratio 0.5 1.42 and 6.26 (top to bottom)



Collapse of columns simulation *Gerris* $\mu(I)$

optimisation

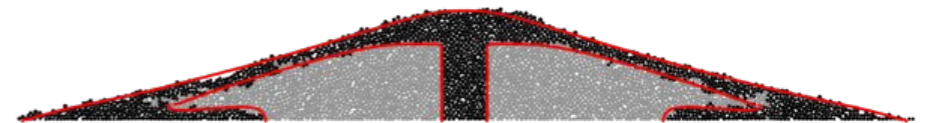
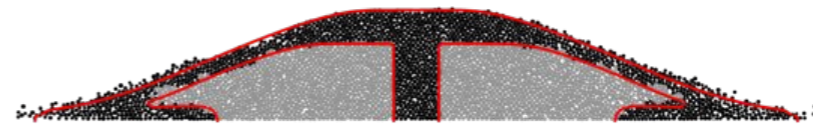
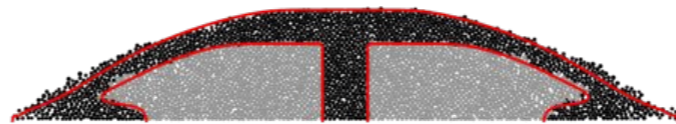
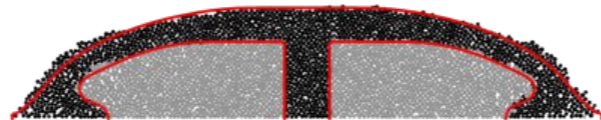
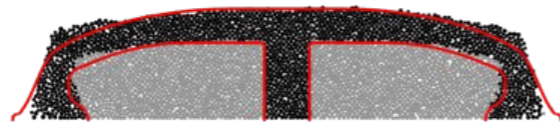
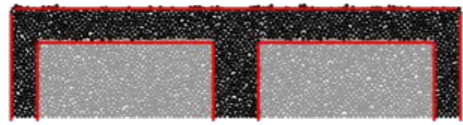
$$\mu(I) = \mu_s + \frac{\Delta\mu}{\frac{I_0}{I} + 1}$$



final values

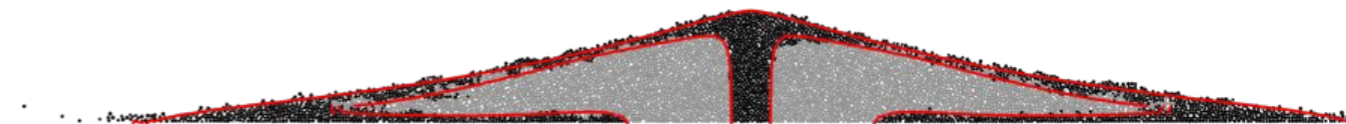
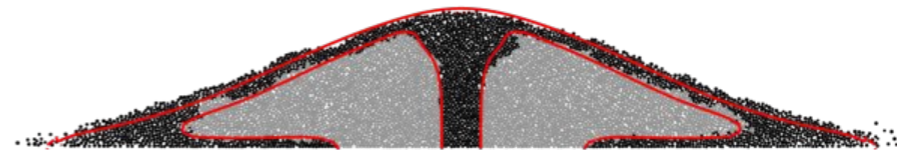
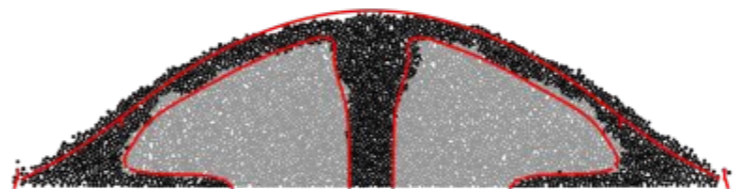
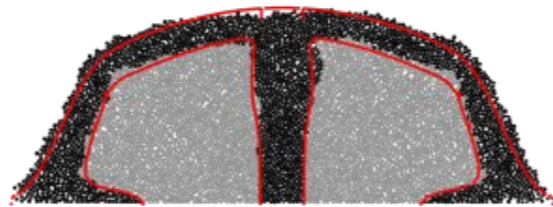
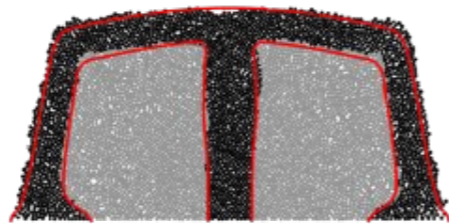
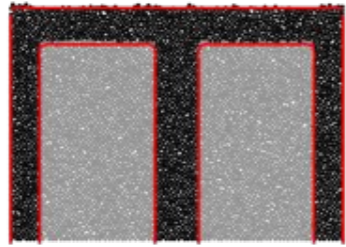
$$\mu_s = 0.32 \quad \Delta\mu = 0.28 \quad I_0 = 0.4$$

Collapse of columns simulation *Gerris* $\mu(l)$



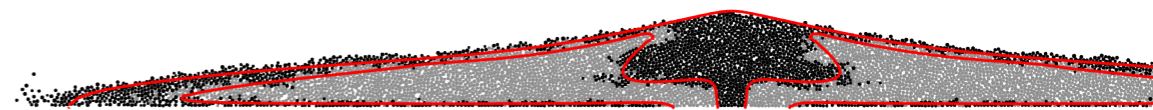
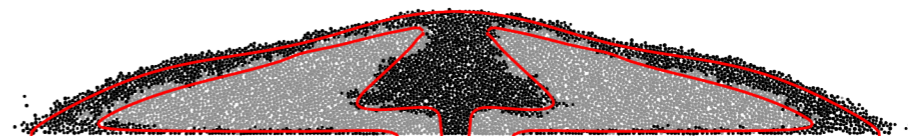
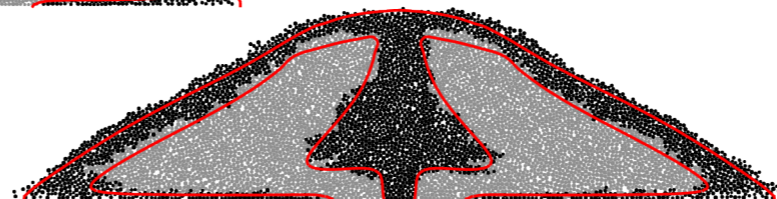
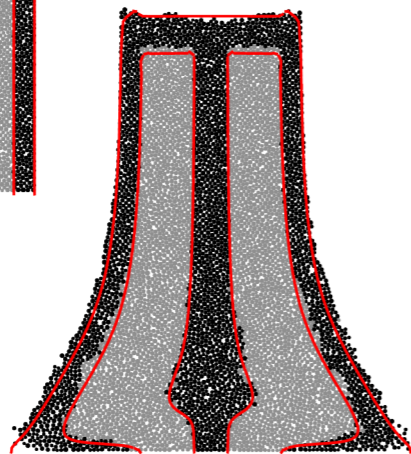
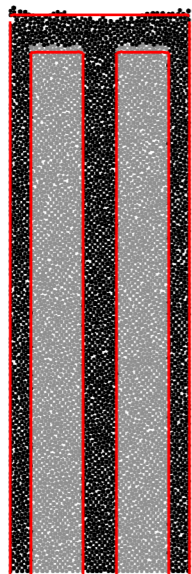
$a = 0.5$ DCM vs *Gerris* $\mu(l)$

Collapse of columns simulation *Gerris* $\mu(l)$



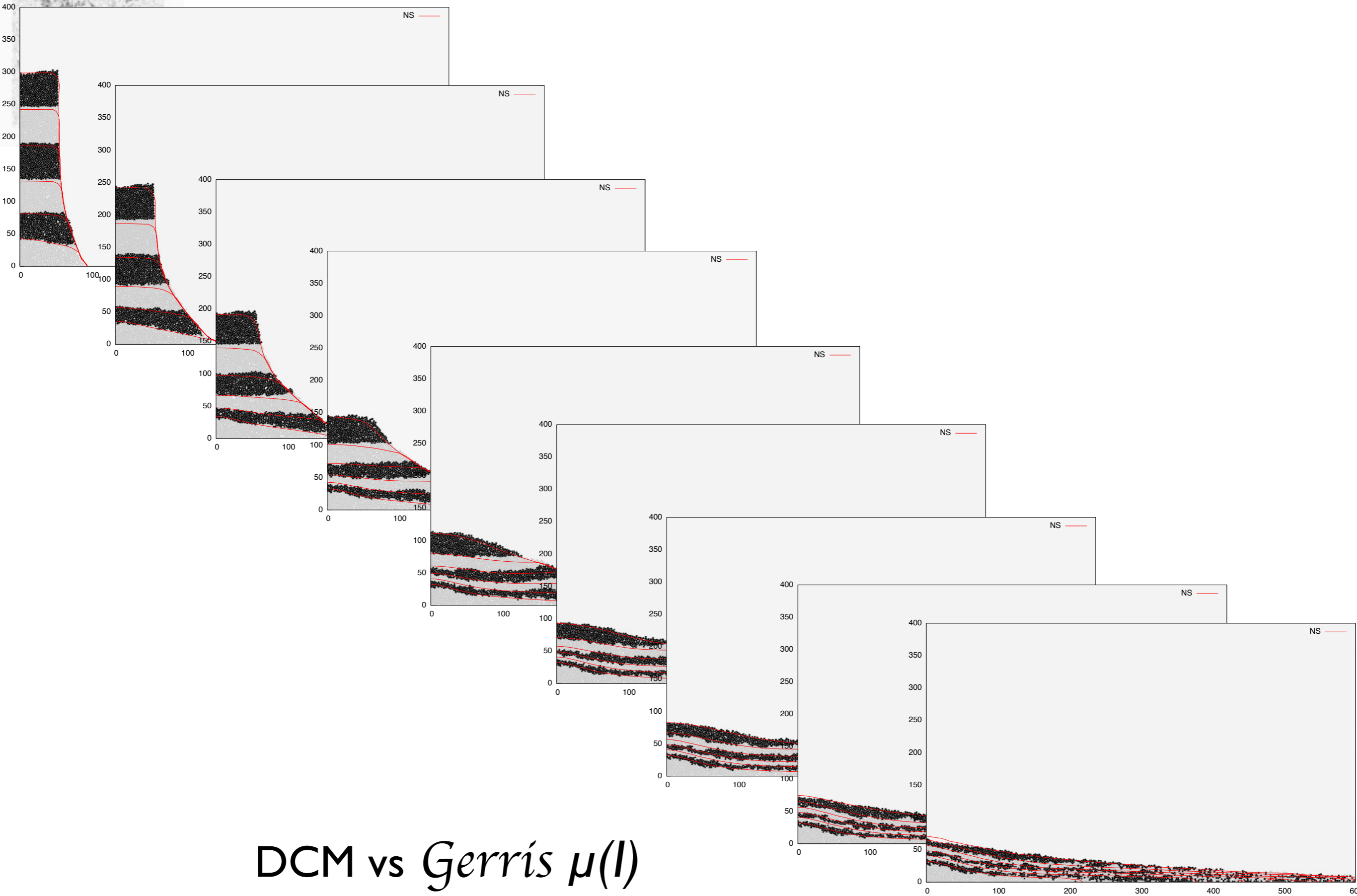
$a = 1.42$ DCM vs *Gerris* $\mu(l)$

Collapse of columns simulation *Gerris* $\mu(l)$



$a = 6.6$ DCM vs *Gerris* $\mu(l)$

Collapse of columns simulation *Gerris* $\mu(l)$

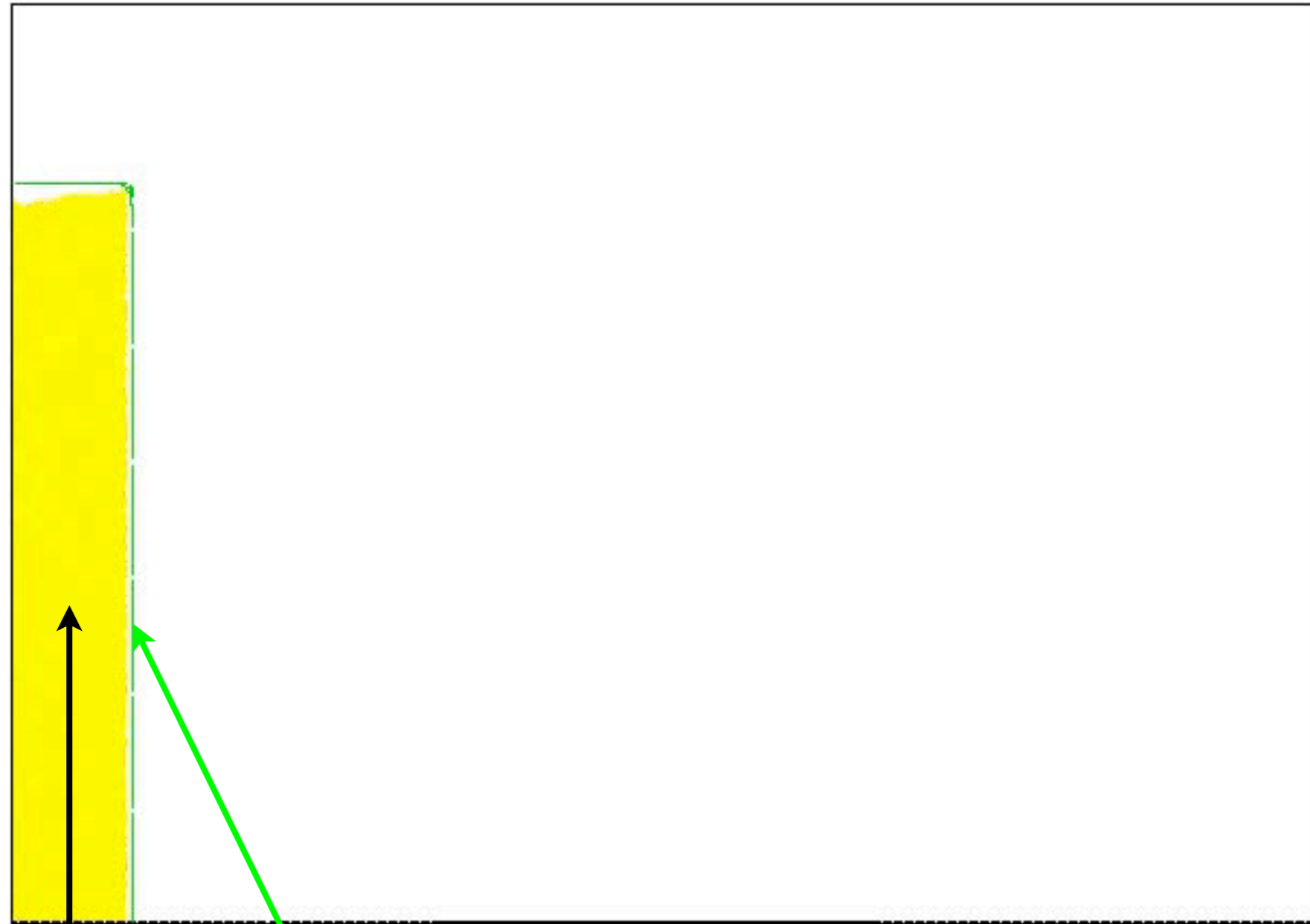


DCM vs *Gerris* $\mu(l)$

Collapse of columns simulation *Gerris* $\mu(l)$



NS/CD t=0.0190

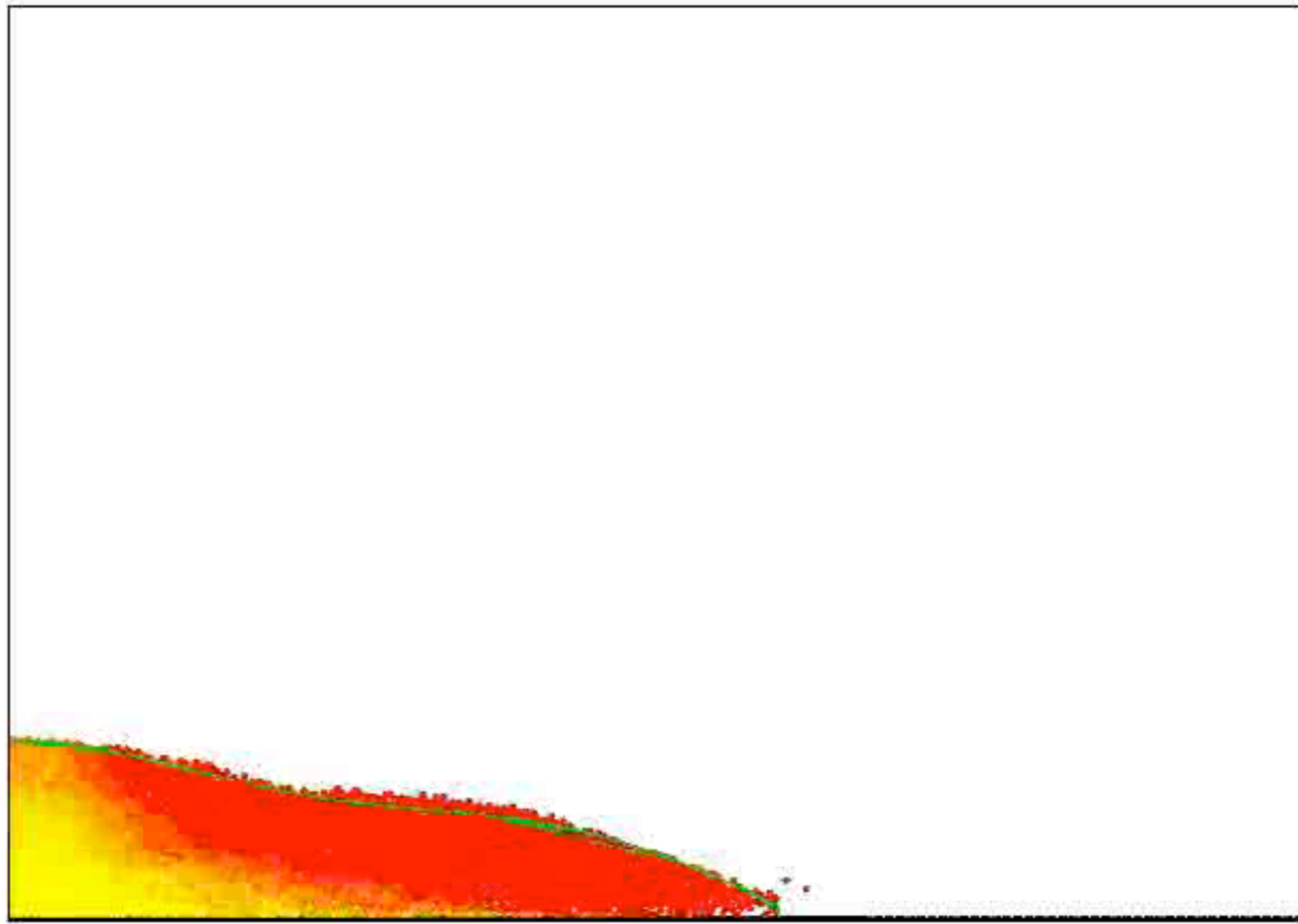


DCM vs *Gerris* $\mu(l)$

Collapse of columns simulation *Gerris* $\mu(l)$



NS/CD t=1.6720



DCM vs *Gerris* $\mu(l)$

Collapse of columns simulation *Gerris* $\mu(l)$



NS/CD t=0.0075

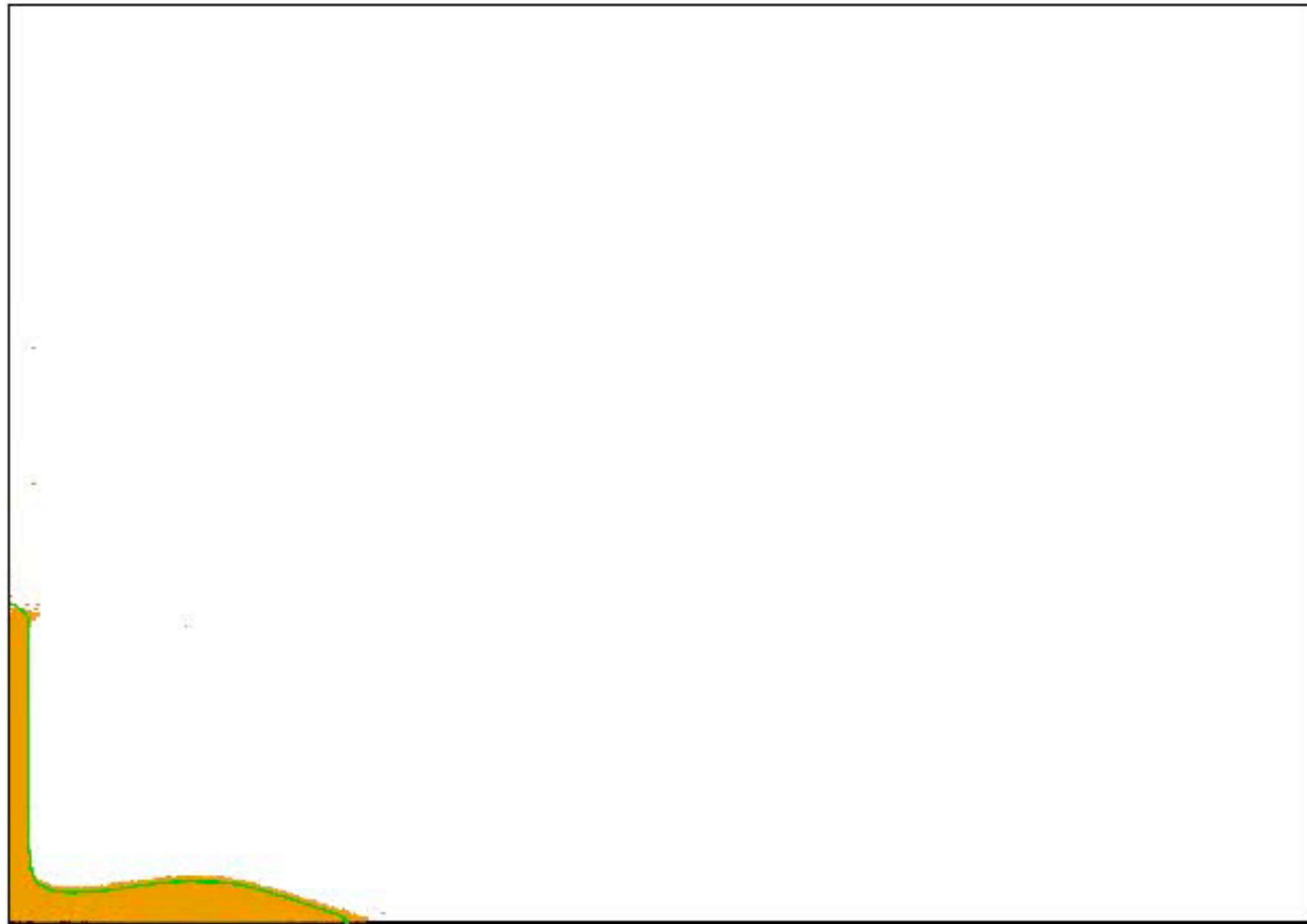


DCM vs *Gerris* $\mu(l)$

Collapse of columns simulation *Gerris* $\mu(l)$



NS/CD t=1.1400



DCM vs *Gerris* $\mu(l)$

Collapse of columns simulation *Gerris* $\mu(l)$

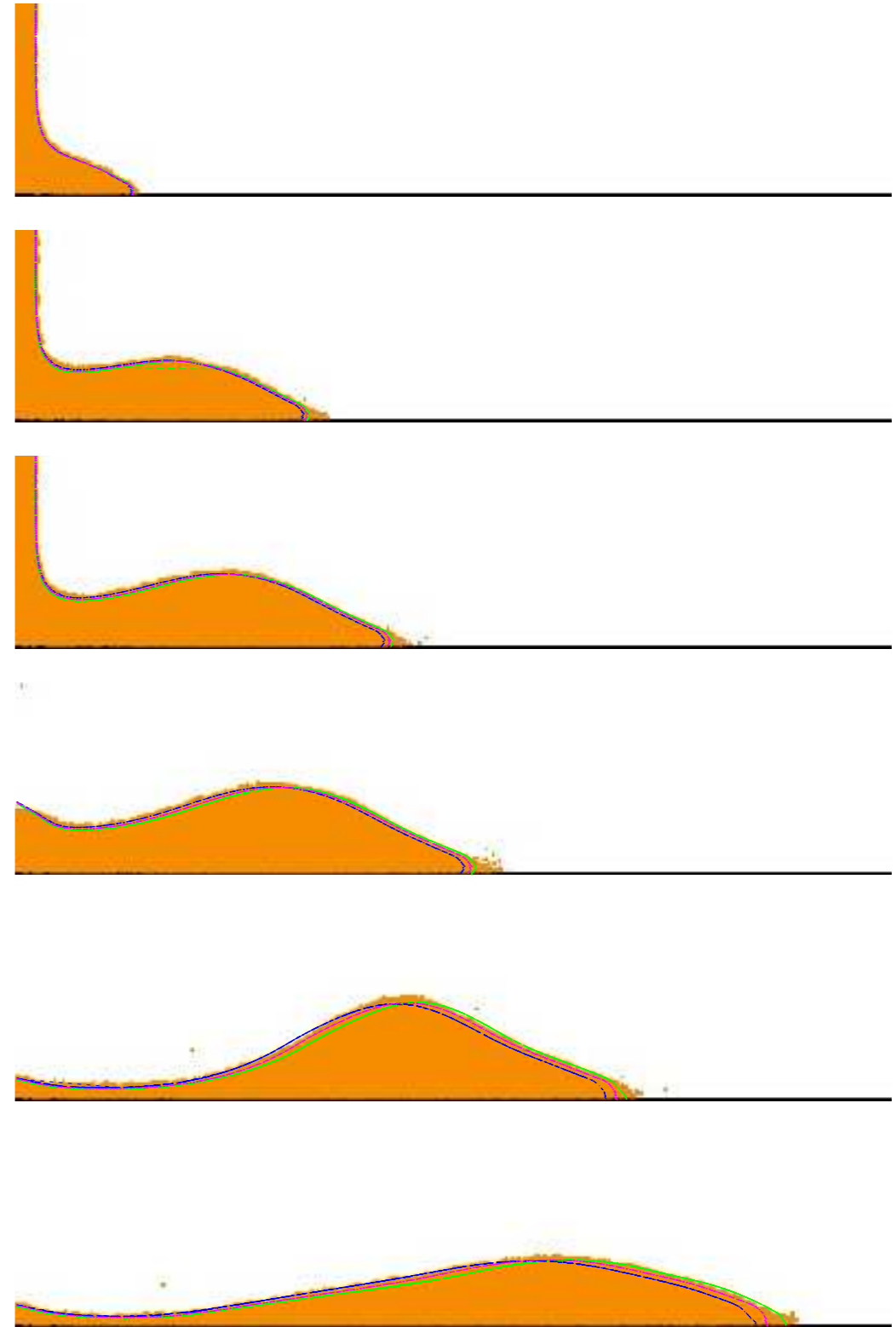
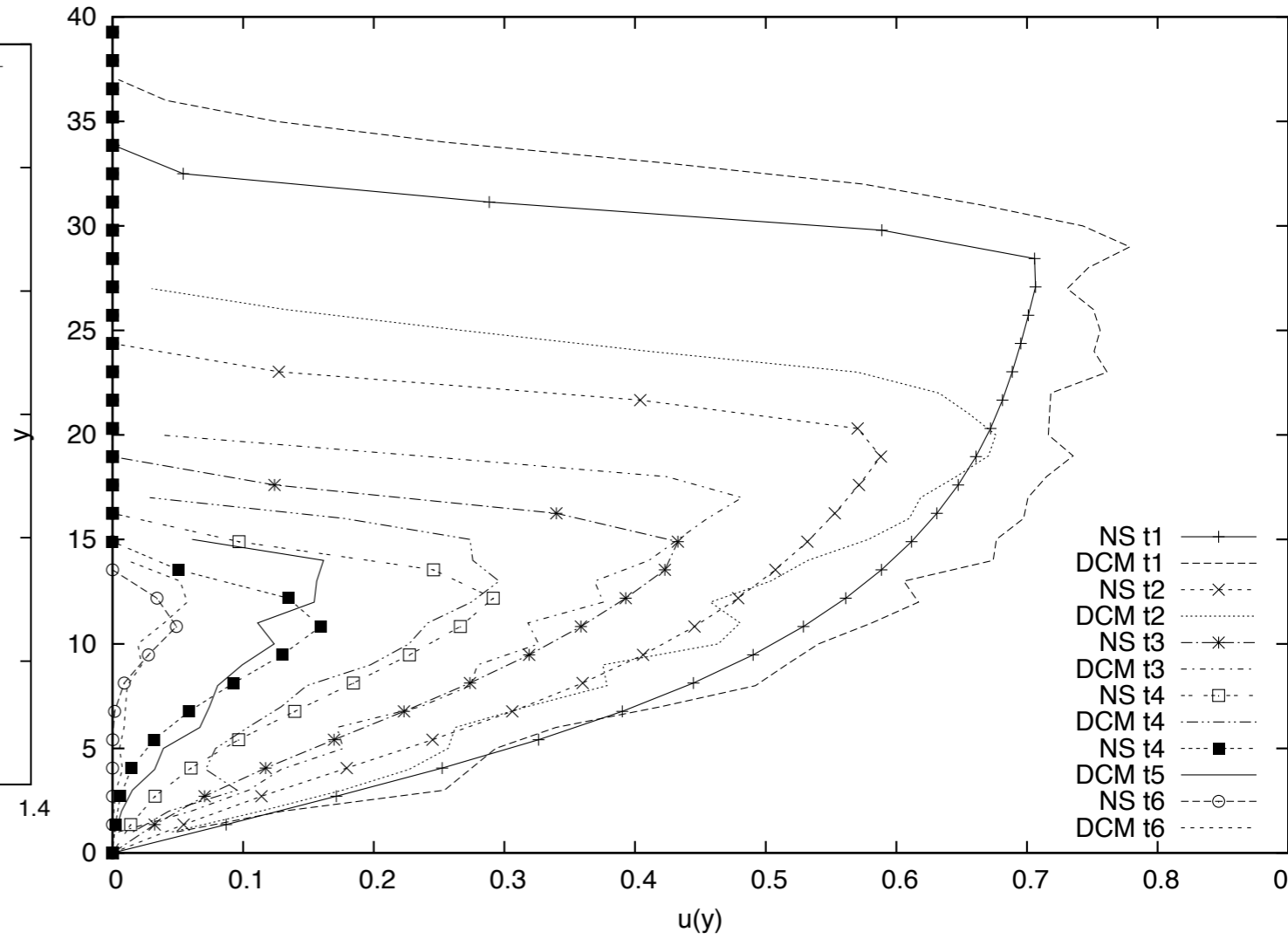
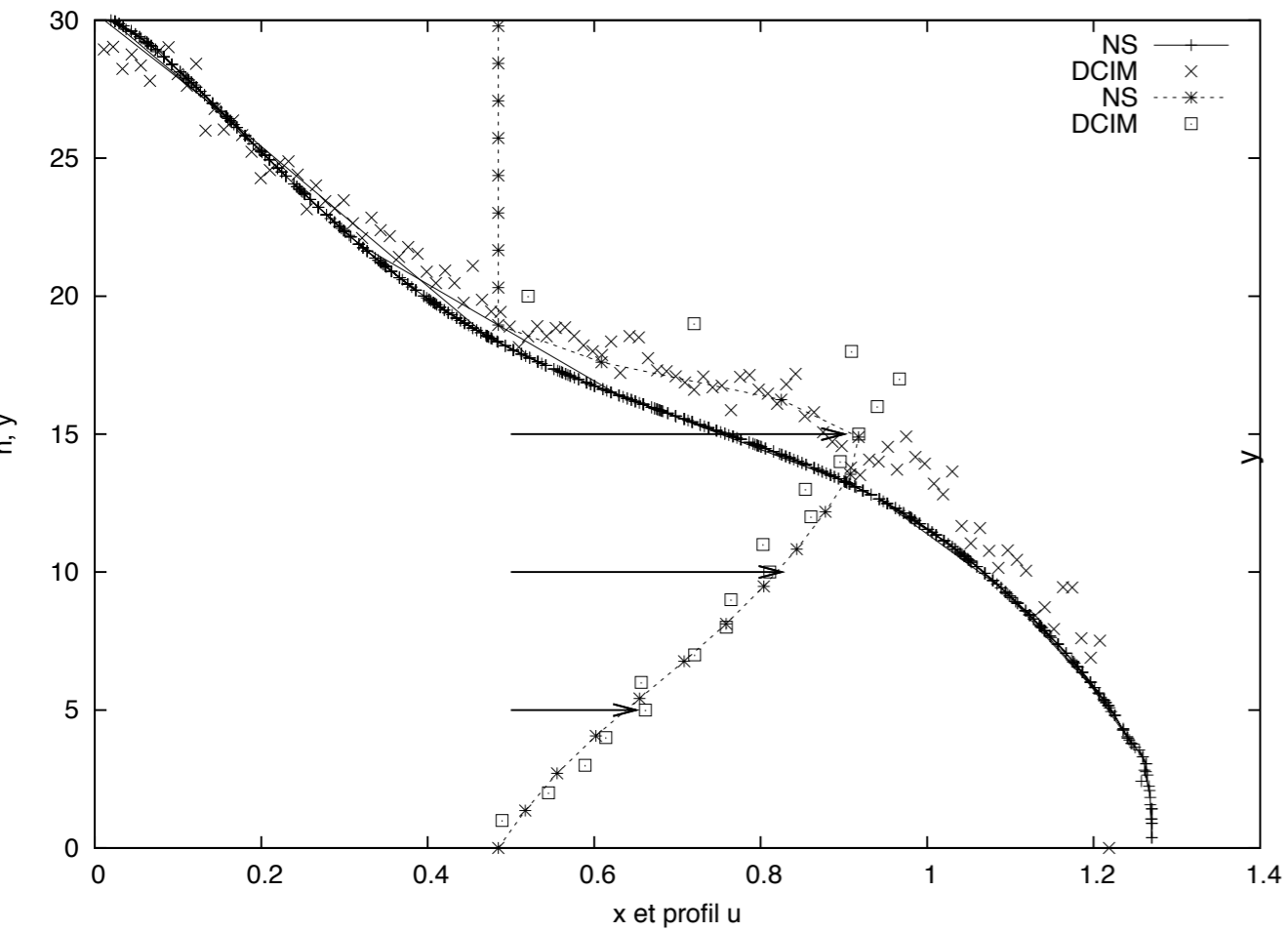


Figure 10: Strip representing a series of snapshots ($t = 0.5, 1.0, 1.2, 1.4, 1.7,$ and 2.0) of a column collapse with aspect ratio $a = 68$. The most advanced curve (in green) corresponds to $\mu_s = 0.3$ $\Delta mu = 0.26$ and $I_0 = 0.30$. the less advanced (in blue) $\mu_s = 0.32$ $\Delta mu = 0.28$ and $I_0 = 0.30$ fits better the end of the heap. The curve in between (in cyan) corresponds to $\mu_s = 0.32$ $\Delta mu = 0.28$ and $I_0 = 0.40$ and fits better the top of the surge.

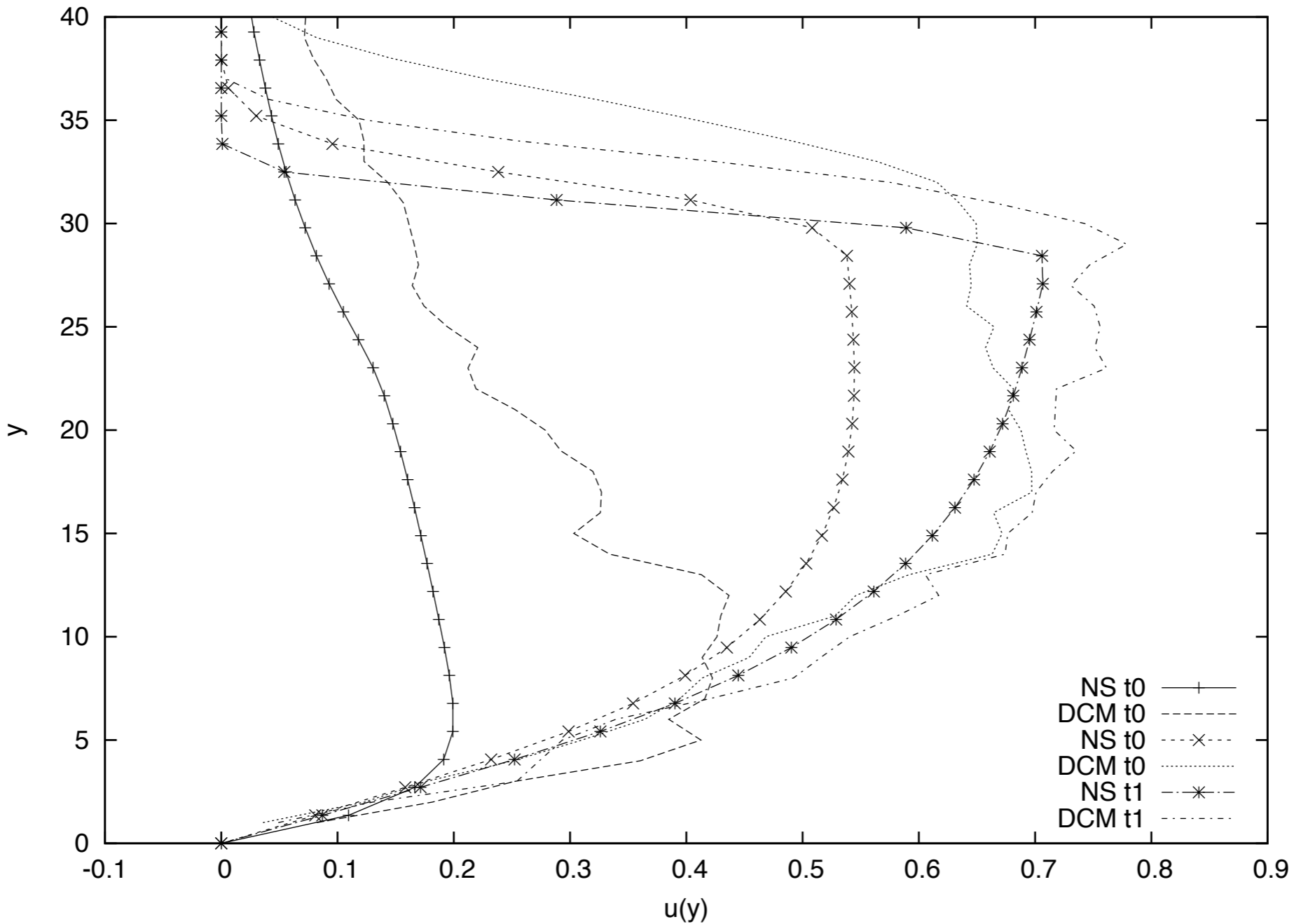
DCM vs *Gerris* $\mu(l)$

● comparaisn of velocity profiles



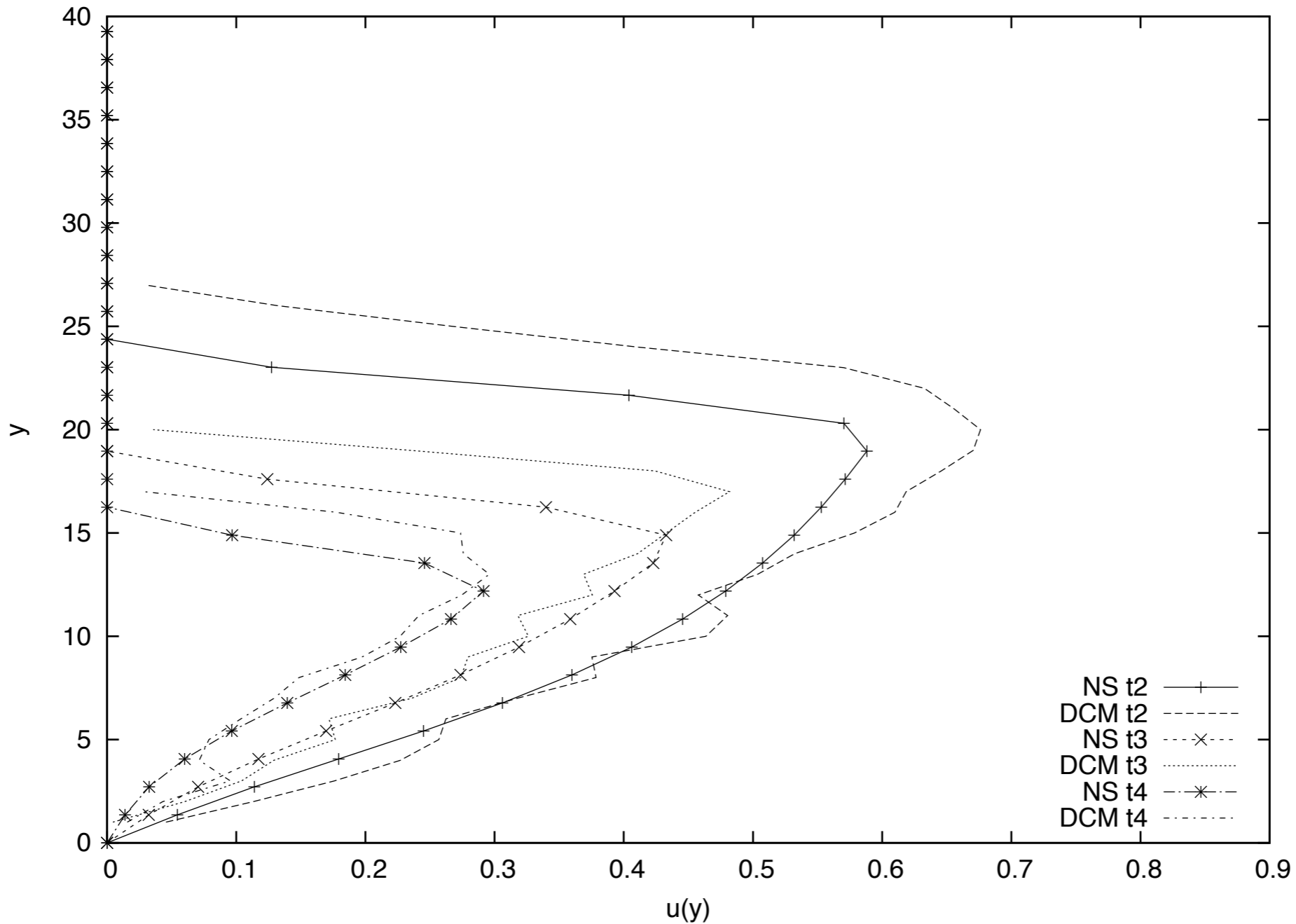
DCM vs *Gerris* $\mu(l)$

● comparaisn of velocity profiles



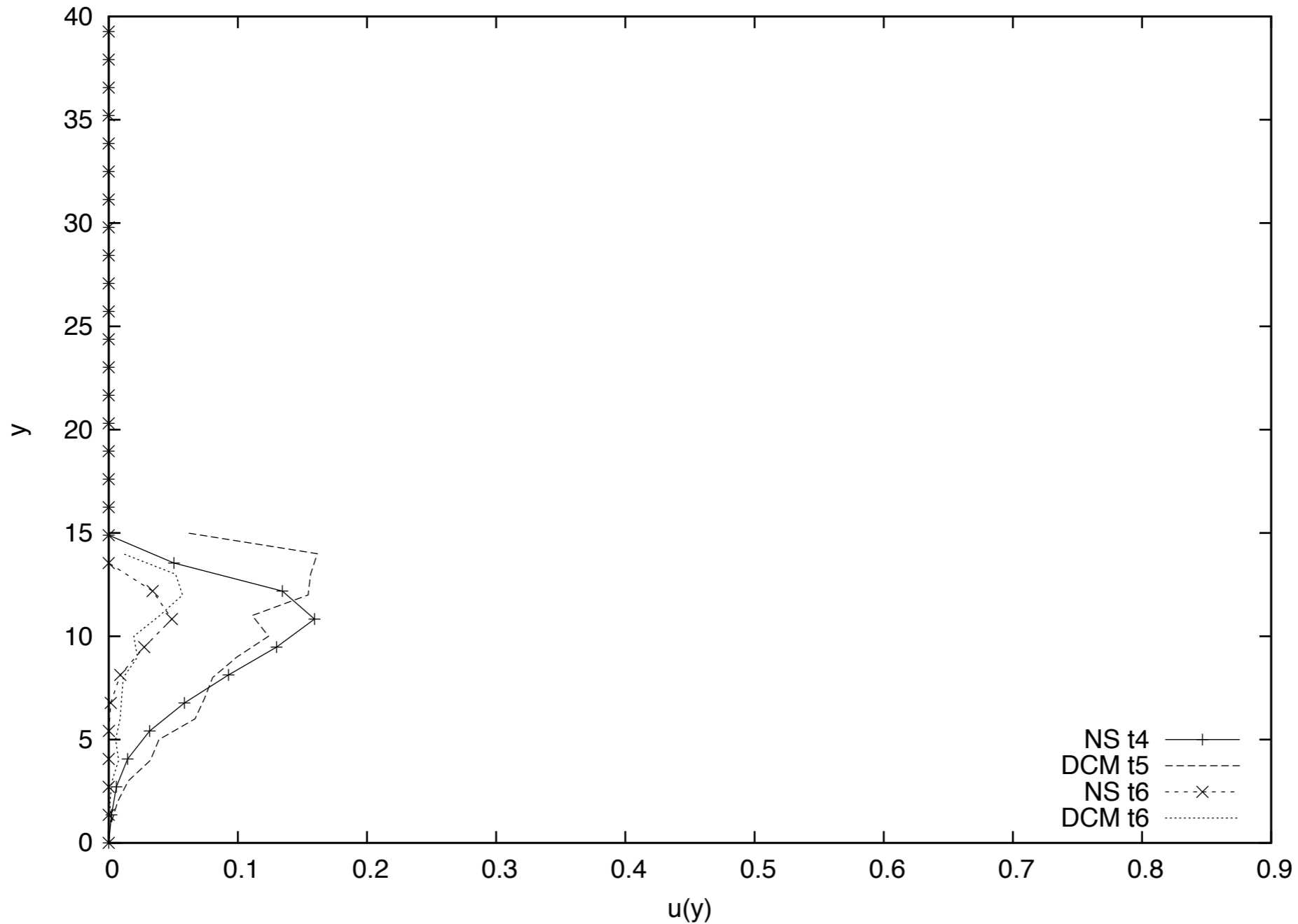
DCM vs *Gerris* $\mu(l)$

● comparaisn of velocity profiles



DCM vs *Gerris* $\mu(l)$

● comparaisn of velocity profiles

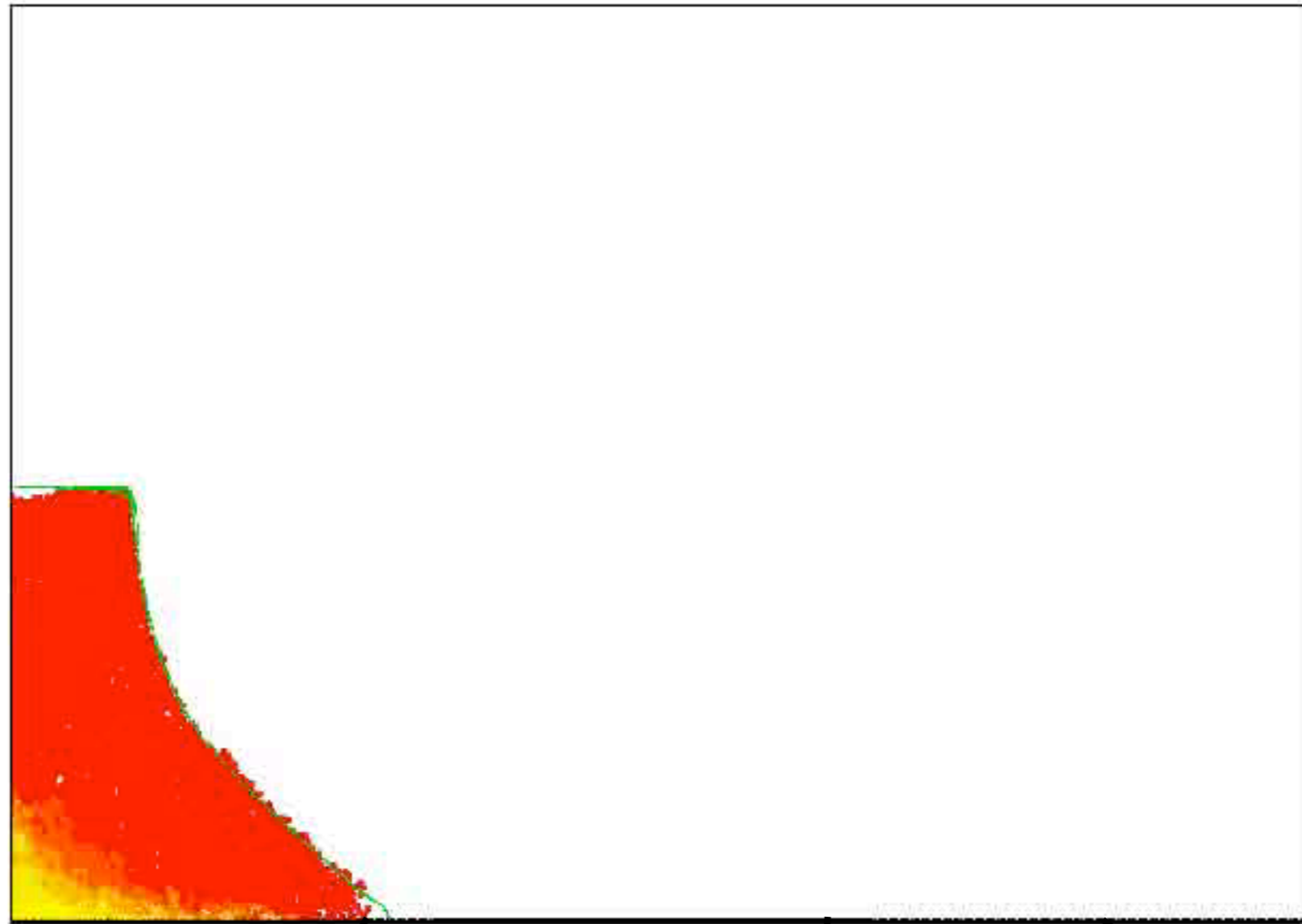


DCM vs *Gerris* $\mu(l)$

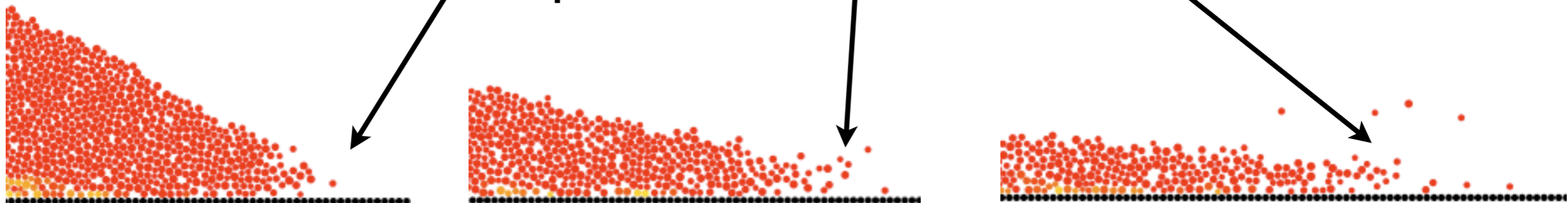
Collapse of columns simulation *Gerris* $\mu(l)$



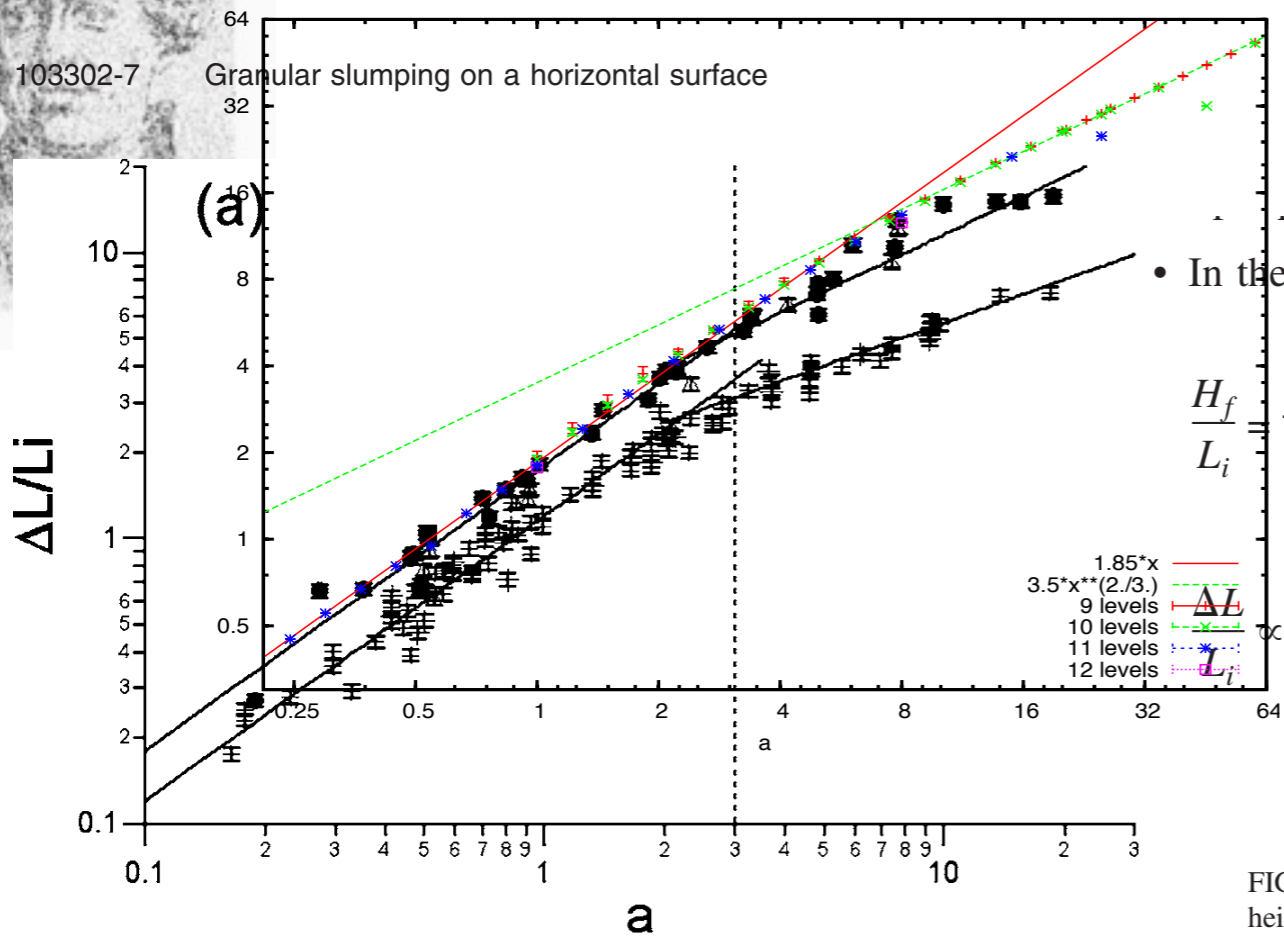
NS/CD t=0.9310



at the tip, $a=6.6$ $t=1.33$ 2 2.66



DCM vs *Gerris* $\mu(l)$



• In the axisymmetric geometry

$$\frac{H_f}{L_i} \propto \begin{cases} a & a \lesssim 0.74, \\ 0.74 & a \gtrsim 0.74, \end{cases}$$

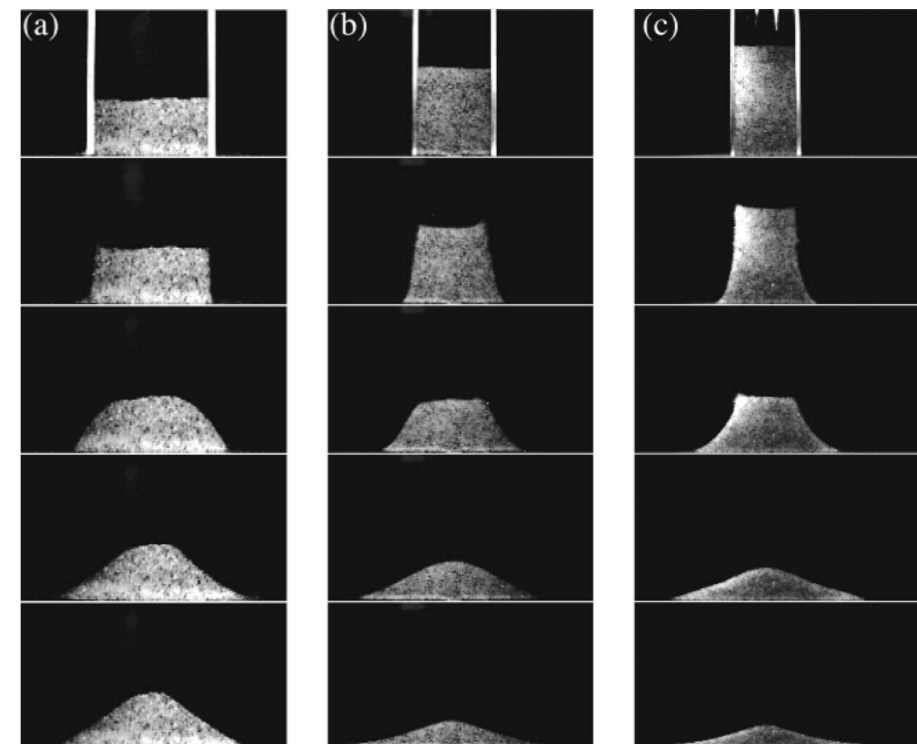
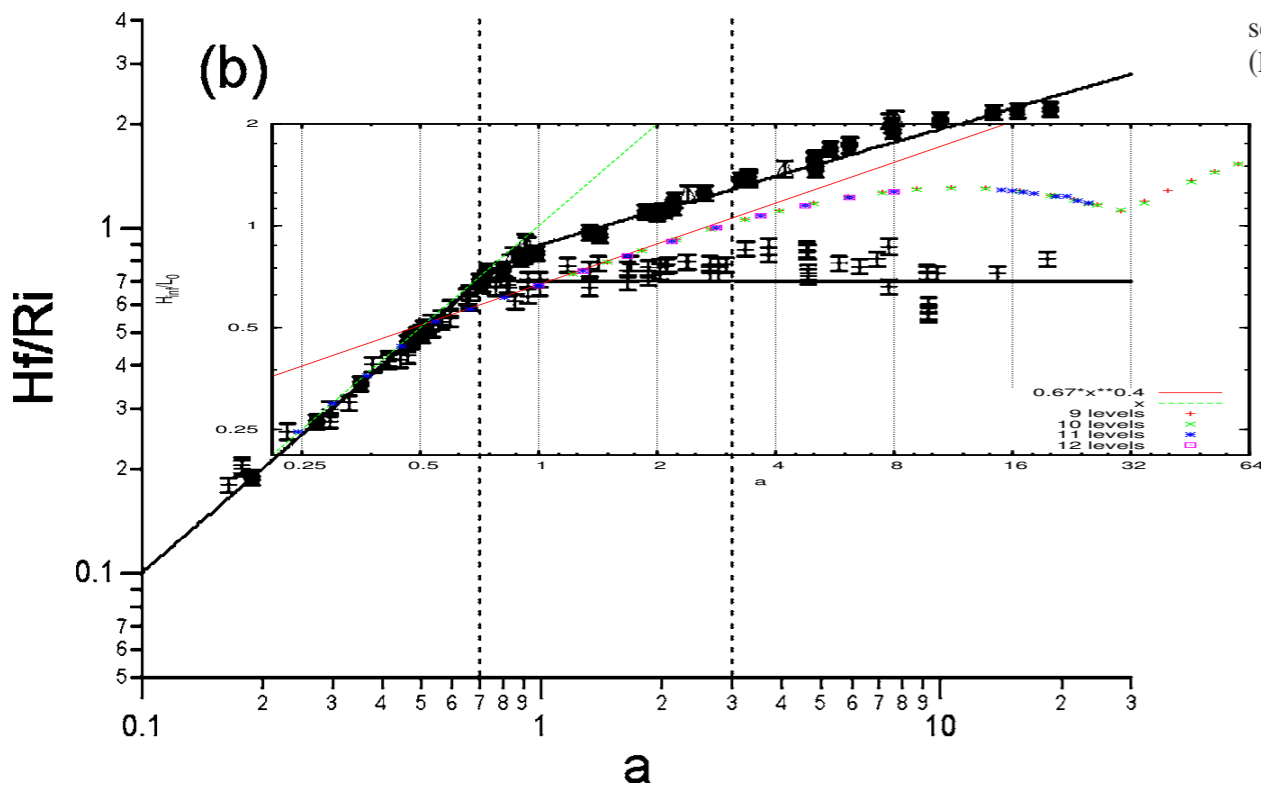
$$\frac{\Delta L}{L_i} \propto \begin{cases} a & a \lesssim 3, \\ a^{1/2} & a \gtrsim 3. \end{cases}$$

• In the rectangular channel:

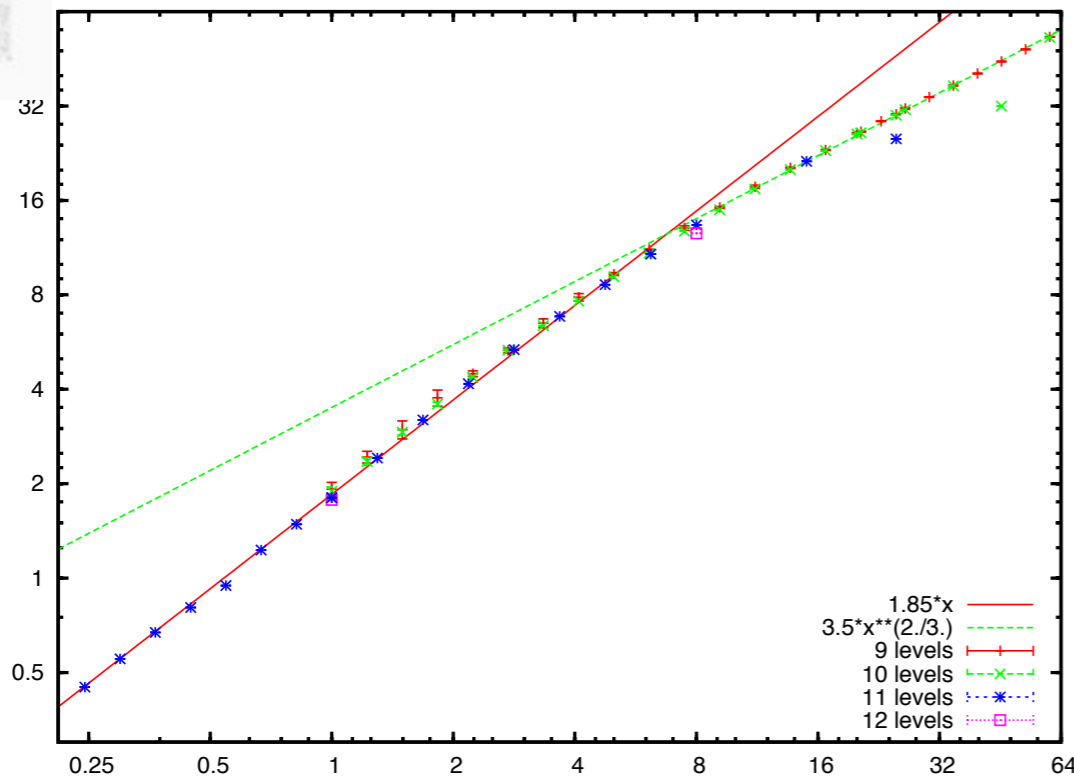
$$\frac{H_f}{L_i} \propto \begin{cases} a & a \lesssim 0.7, \\ a^{1/3} & a \gtrsim 0.7, \end{cases}$$

$$\frac{\Delta L}{L_i} \propto \begin{cases} a & a \lesssim 3, \\ a^{2/3} & a \gtrsim 3. \end{cases}$$

FIG. 6. Scaled runout $\Delta L/L_i$ (a) and scaled deposit height H_f/L_i (b) as functions of a . Circles and triangles correspond to experiments performed in the 2D channel working respectively with glass beads of diameter $d = 1.15$ mm or $d = 3$ mm. Crosses correspond to the data set of axisymmetric collapses from Lajeunesse *et al.* (Ref. 10).

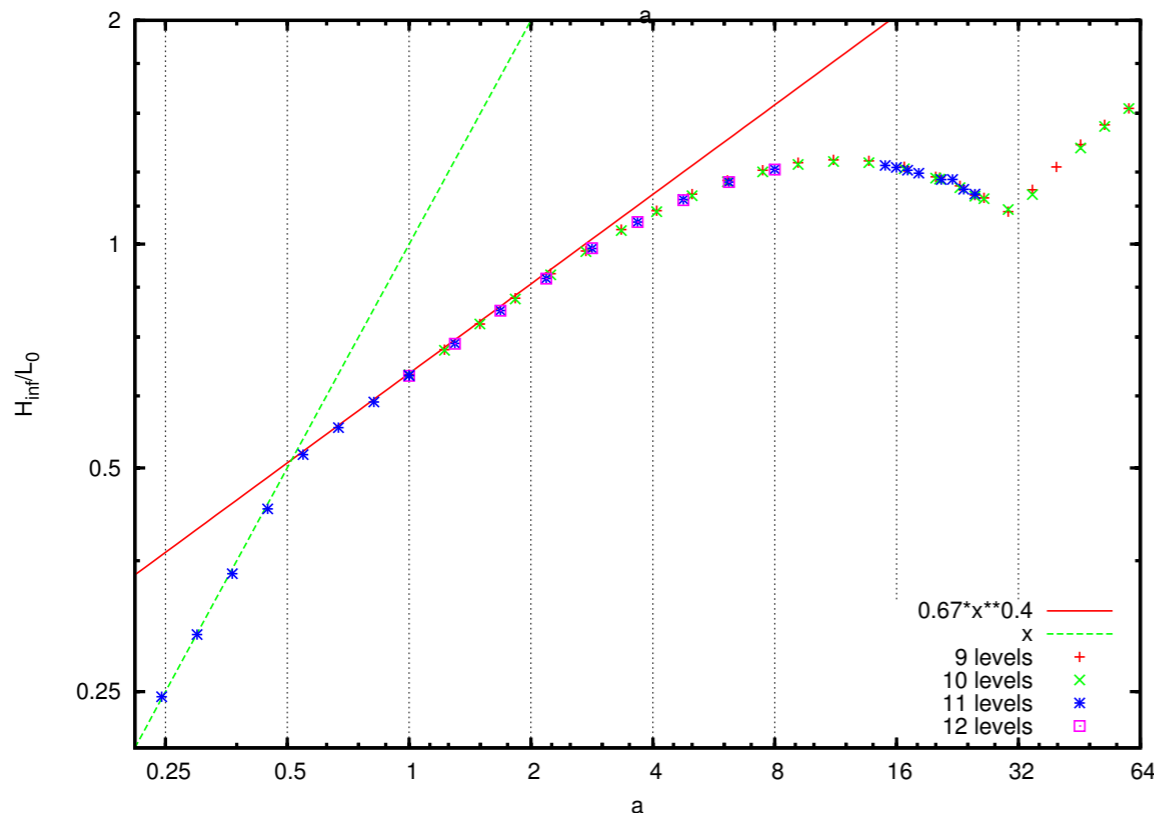


Collapse of columns simulation *Gerris* $\mu(l)$



Normalised final deposit extent as a function of aspect ratio a .

Well-defined power law dependencies with exponents of 1 and 2/3 respectively.



We recover the experimental scaling [Lajeunesse et al. 04] and [Staron et al. 05]. Differences between the values of the prefactors are due to the difficulties to obtain the run out length: friction in the Navier Stokes code tends to underestimate it, whereas direct simulation shows that the tip is very gaseous, it can no longer explained by a continuum mechanic description.

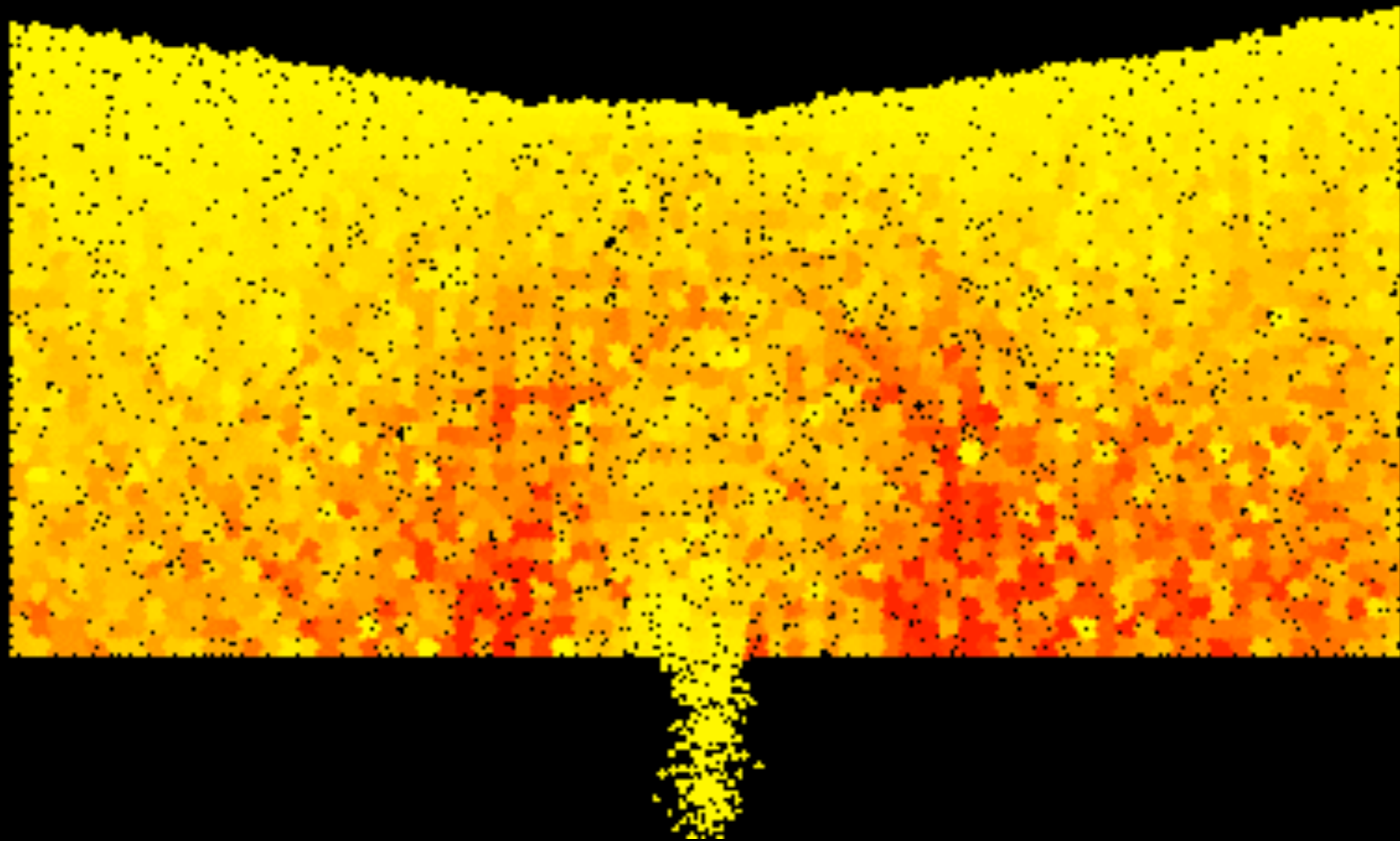


perspectives

- Other examples under investigation

- Silo, hopper



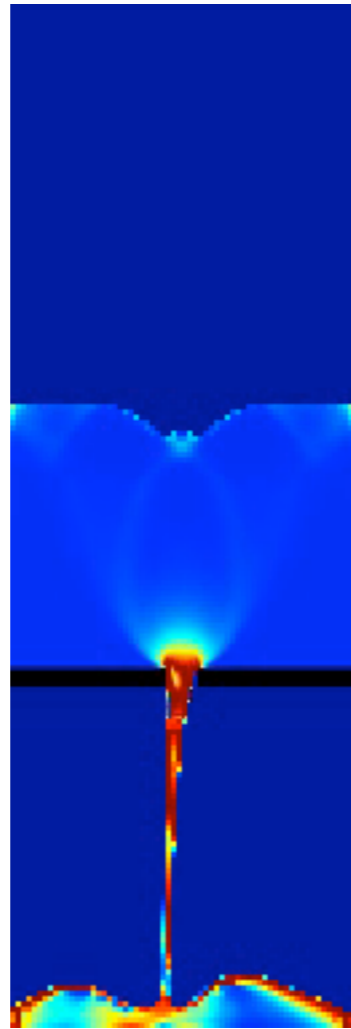


- Vidange de Silo

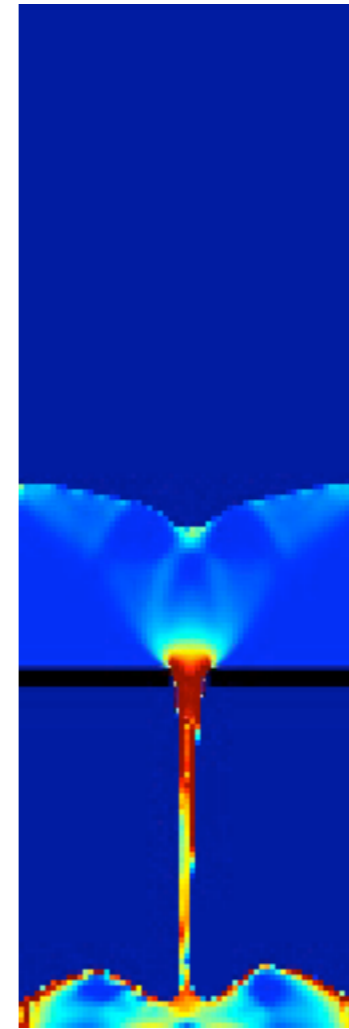




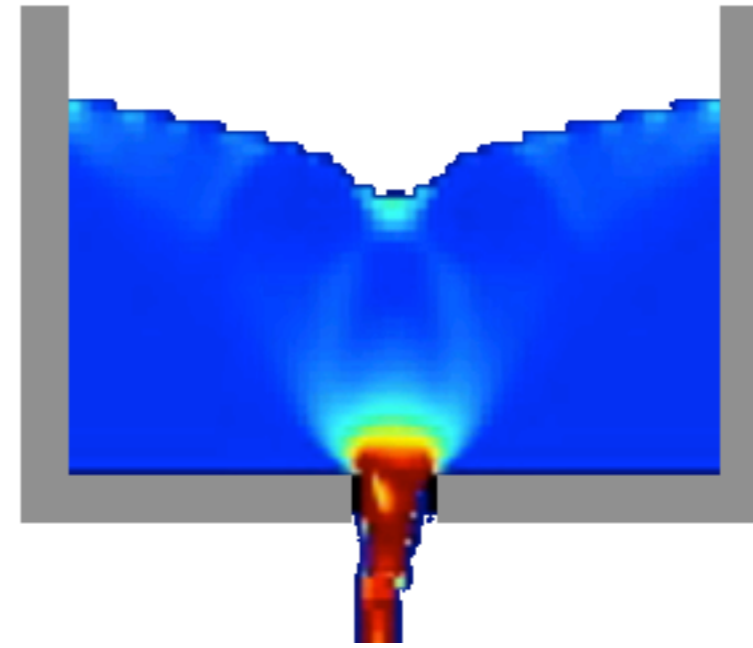
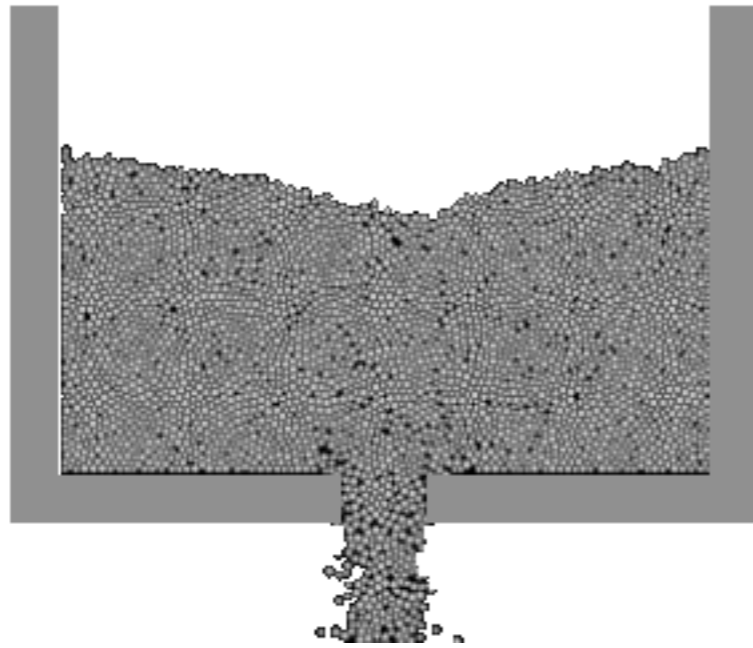
- Vidange de Silo



12 grains

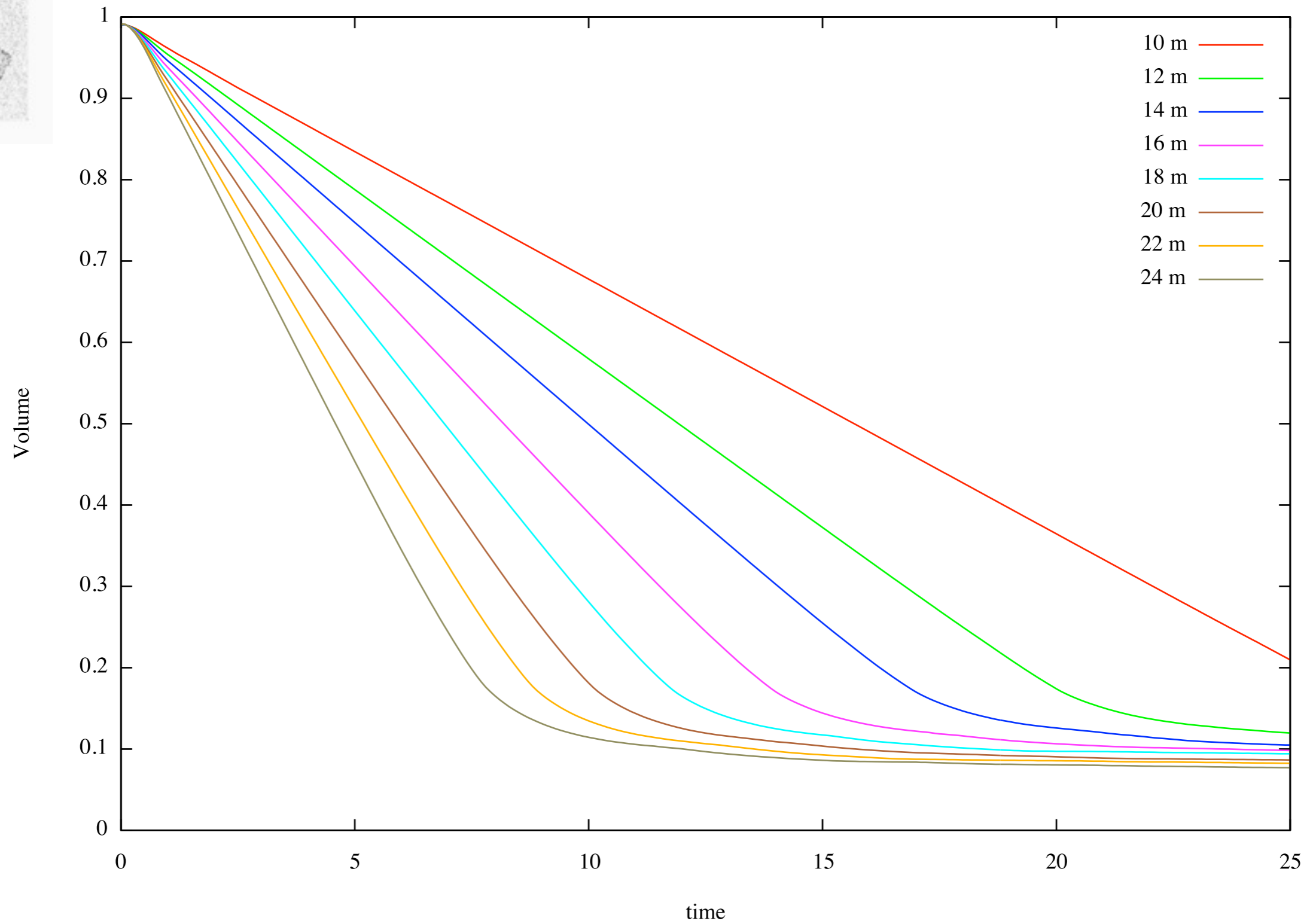


14 grains





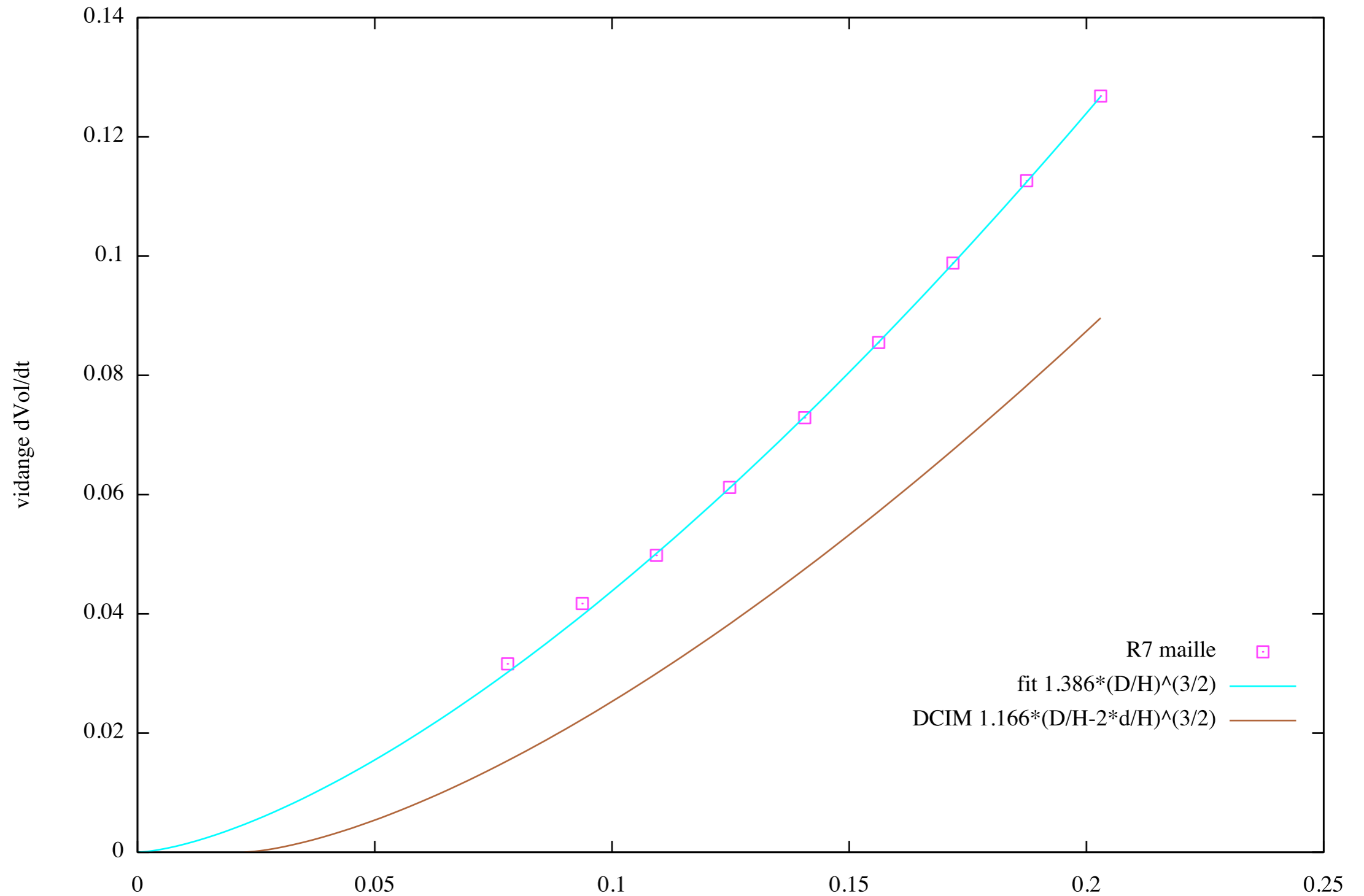
● Vidange de Silo





● Vidange de Silo

$$\frac{d\bar{V}}{dt} = 1.166\left(\frac{D}{L} - k\frac{d}{L}\right)^{3/2}$$



loi de Beverloo Hagen



conclusion

- $\mu(I)$ obtained from experimental flows of dry granular flows [Jop et al. 06], implemented it in *Gerris*
- test case: analytical solution of steady avalanche (Bagnold solution)
- collapse of granular columns (shape as function of time compared to Discrete Simulations).
- The experimental trends of the scaling of the run out are reobtained
- Saint Venant Savage Hutter to be compared with.
- complete spectra: discrete grains/ Saint Venant/ Navier Stokes and plasticity

This opens the door to systematic studies of granular flows using this continuum approach.



références:

P. Jop, Y. Forterre, O. Pouliquen, (2006) "A rheology for dense granular flows", Nature 441, pp. 727-730 (2006)

P.-Y. Lagrée, L. Staron and S. Popinet (2011) "The granular column collapse as a continuum: validity of a two-dimensional Navier-Stokes model with a $\mu(I)$ -rheology", JFM pp1-31 , doi:10.1017/jfm.2011.335

E. Lajeunesse, A. Mangeney-Castelnau, and J.-P. Vilotte, (2004) "Spreading of a granular mass on an horizontal plane», Phys. Fluids, 16(7), 2371-2381.

L. Staron & E. J. Hinch (2005) "Study of the collapse of granular columns using two-dimensional discrete-grain simulation", J. Fluid Mech. (2005), vol. 545, pp. 1-27.