Non linear instability of the incompressible Euler with gravity equations in the ICF context

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## 1 Physical context

Ablation front in the inertial confinement fusion of Deuterium and Tritium

Study of the propagation of a defect in the target (which is not a perfect sphere) :

we need to obtain the critical density for the beginning of the thermonuclear reaction, and growth of defects can prevent this

#### Model presented here :

the ablative Rayleigh-Taylor system in a strip with constant gravity and continuous transition between the high density  $\rho_a$  and the vacuum  $(\rho \to 0)$ ; A (noncomplete)list of references :

J.W. Strutt (Lord Rayleigh) 1883

G. Taylor 1950

S. Chandrasekhar1961

H.J. Kull and S.I. Anisimov 1986

Takabe et al 1996

V. Goncharov 1998

E. Grenier 2000, L 2000

C. Cherfils, L, P.A. Raviart : 2001

J. GARNIER, C. CHERFILS, P.A. HOLSTEIN 2003

S. Cordier, E. Grenier and Y. Guo 2003

Y. GUO AND H.J. HWANG 2003

B. Helffer and L 2003, 2007

P. CLAVIN AND L. MASSE 2004

B. Desjardins and E. Grenier 2006

R. Poncet 2007

- Mixing of fluids : modeled by a variable density in the medium.
- Presence of gravity.
- Euler equations.
- Energy equation : thermal conduction model.
- Quasi-isobaric model with Fourier law :

 $\operatorname{div}(C_p \rho T \vec{u} + \kappa(T) \nabla T) = 0, \rho T = \operatorname{cst}, \kappa(T) = \kappa_0 T^{\nu}, p = \rho T$ 

Model system 0 (introducing  $Z(\rho) = \frac{\rho_a^{\nu+1}}{(\nu+1)\rho^{\nu+1}}$ ) :

$$\begin{cases} \partial_t \rho + \operatorname{div}(\rho \vec{u}) = 0\\ \partial_t(\rho \vec{u}) + \operatorname{div}(\rho \vec{u} \otimes \vec{u} + pId) = \rho \vec{g} \\ \operatorname{div}(\vec{u} + L_0 V_a \nabla Z(\rho)) = 0. \end{cases}$$
(1)

Stationary solution  $\rho_0(x) = \rho_a \xi(\frac{x}{L_0}), u_0(x) = -\frac{V_a}{\xi}, p_0$  solution of  $(p - \frac{\rho_a V_a^2}{\xi})' = \rho_a g\xi$ , where  $\xi$  satisfies

$$\dot{\xi} = \xi^{\nu+1}(1-\xi).$$

Concentrate on  $\rho_0$  only. and forget this system and solution.

# 2 Euler equations with gravity around a general profile

Profile  $(\rho_0(x), \vec{0}, p_0(x))$  with  $\nabla p_0 = \rho_0 \vec{g}$ . Introduce  $k_0(x) = \frac{\rho'_0(x)}{\rho_0(x)}$ . Euler in 2d :  $x \in (-\infty, +\infty), z \in [0, a]$ .

System of Euler equations in the unknowns  $\vec{u}, T = \frac{\rho_0(x)}{\rho(x,y,t)}, q = \frac{p-p_0(x)}{\rho_0(x)}$ :  $\begin{cases} \partial_t T + (\vec{u}.\nabla)T = k_0 uT \\ \partial_t \vec{u} + (\vec{u}.\nabla)\vec{u} + (T-1)\vec{g} + \nabla q + k_0(x)q\vec{e_1} = 0 \\ \operatorname{div} \vec{u} = 0 \end{cases}$ 

(2)

Particular solution  $(\vec{0}, 1, 0)$ .

Linearized system (operator Emod) :

$$\begin{cases} \partial_t T_1 - k_0 u_1 = 0 \\ \partial_t \vec{u}_1 + T_1 \vec{g} + \nabla q_1 + k_0(x) q_1 \vec{e}_1 = 0 \\ \operatorname{div} \vec{u}_1 = 0 \end{cases}$$
(3)

Linearized system with left hand side

$$\begin{cases} \partial_t T_1 - k_0 u_1 = S_0 \\ \partial_t \vec{u}_1 + T_1 \vec{g} + \nabla q_1 + k_0(x) q_1 \vec{e}_1 = \vec{S} \\ \operatorname{div} \vec{u}_1 = 0 \end{cases}$$
(4)

Rayleigh equation (1) : apply  $\partial_t \partial_z^2$  on the equation on  $u_1$  and use  $\partial_t T_1$  and  $\partial_z^z q_1 = \partial_t \partial_x u_1$ :

$$\partial_t^2 \partial_z^2 u_1 + g k_0(x) \partial_z^2 u_1 + \rho_0^{-1} \partial_x (\rho_0 \partial_x \partial_t^2 u_1)$$
  
=  $\partial_t \partial_z^2 S_1 - g \partial^2 z S_0 - \rho_0^{-1} \partial_x (\rho_0 \partial_t \partial_z S_1).$  (5)

Rayleigh equation (2) : seek  $u_1(t, x, z) = \hat{u}(x)e^{\gamma t} \cos kz$  in the homogeneous equation :

$$-\hat{u}'' - k_0(x)\hat{u}' + (k^2 + \frac{gk^2}{\gamma^2}k_0(x))\hat{u} = 0$$
(6)

Solution

$$(\hat{u}_1\cos kz, -\hat{u}_1'\frac{\sin kz}{k}, \frac{k_0(x)}{\gamma}\cos kz, -\frac{\gamma}{k^2}\hat{u}_1'\cos kz)e^{\gamma t}.$$

Rayleigh equation (2'):

$$-\frac{d}{dx}(\rho_0\hat{u}) + (k^2\rho_0 + \frac{gk^2}{\gamma^2}\rho_0'(x))\hat{u} = 0$$
(7)

Rayleigh equation (2"):

$$-k^{-2}\frac{d^2}{dx^2}\hat{w} + (1 + \frac{g}{\gamma^2}k_0(x) + \frac{1}{2}k^{-2}(k'_0 + \frac{1}{2}k_0^2))\hat{w} = 0$$
(8)

Assume  $k_0 \ge 0$ . From (2'), integrating, one deduces  $-g > 0 \Rightarrow$  no solution such that  $\gamma^2$  real positive  $- \text{ if } \gamma^2 > (\max k_0)|g| = \Lambda^2$ , no solution  $- \text{ if } \gamma^2 > |g|k$ , no solution the last equality thanks to

$$(1 - \frac{|g|k}{\gamma^2}) \int \rho_0((\hat{u}')^2 + k^2 \hat{u}^2) dx = \frac{|g|k}{\gamma^2} \int \rho_0(\hat{u}' - k\hat{u})^2 dx.$$

#### **3** Semi-classical result

## Condition : $k_0$ has a unique nondegenerate maximum. This condition is satisfied by $k_0(x) = \frac{\xi'(x)}{\xi(x)} = \xi^{\nu}(1-\xi)$ . Maximum at $\xi = \frac{\nu}{\nu+1}$ . This leads to **Proposition**

There exists a  $h_0 > 0$  such that, for all  $k \ge \frac{1}{h_0}$ , there exists a sequence  $\gamma_n(k)$  such that

$$1 - \frac{|g|}{\gamma_n^2(k)} k_0^{max} + \left(\frac{|g|}{2\gamma_n^2(k)} |k_0''|\right)^{\frac{1}{2}} \left(n + \frac{1}{2}\right) = 0(k^{-\frac{3}{2}}).$$

This proves that there exists at least one value of  $\gamma$  and one value of k such that  $\gamma \in (\frac{\Lambda}{2}, \Lambda)$ .

### 4 Nonlinear result

Under the assumptions  $k_0$  bounded (as well as all the derivatives) and admits at least a nondegenerate maximum,  $k_0 \rho_0^{-\frac{1}{2}}$  bounded, there exists an initial value for (2) close (in a  $\delta$  sense) to the stationary solution which departs in a time of order  $\ln \frac{1}{\delta}$  of a given finite quantity from this solution.

No need of  $\rho_0 \ge \rho_{min} > 0$ . Need  $\rho_0 > 0$  everywhere.

### 5 The Duhamel principle

Energy equality :

Note that system (3) leads to the equation on the incompressible quantity  $\vec{u}$ :

$$\partial_{t^2}^2 \vec{u} + k_0 u_1 \vec{g} + \rho_0^{-1} \partial_x (\rho_0 \partial_t q) = \partial_t \vec{S} - S_0 \vec{g}.$$

from which one deduces

$$\begin{split} &\frac{1}{2} \int (\rho_0(\partial_t \vec{u})^2 + k_0 \rho_0 g u_1^2) dx dz = \frac{1}{2} \int (\rho_0(\partial_t \vec{u}(0))^2 + k_0 \rho_0 g u_1(0)^2) dx dz \\ &= \int_0^t (\int \rho_0(\partial_t \vec{S} - S_0 \vec{g}) \partial_t \vec{u} dx dz) ds \end{split}$$

General result : The solution of  $\frac{1}{2} \frac{d}{dt} (\int (\rho_0 (\partial_t \vec{u}_N)^2 - g \frac{\rho'_0}{\rho_0} \rho_0 (u_N)^2) dx dz) = g(t, x, \partial_t \vec{u}_N)$ with initial condition  $\partial_t \vec{u}_N (0), \vec{u}_N (0)$ , with the assumption  $|g(t, x, \partial_t \vec{u}_N)| \leq K(t) ||\rho_0^{\frac{1}{2}} \partial_t \vec{u}_n||_{L^2}$ where K is a positive increasing function for  $t \geq 0$  satisfies the inequalities

$$\begin{aligned} ||\rho_0^{\frac{1}{2}} \vec{u}_N||^{\frac{1}{2}} &\leq [C_1 + \int_0^t \sqrt{K(s)e^{-\Lambda s}} ds] e^{\frac{\Lambda}{2}t} \\ ||\rho_0^{\frac{1}{2}} \partial_t \vec{u}_N|| &\leq [C_1 + \int_0^t \sqrt{K(s)e^{-\Lambda s}} ds]^2 e^{\Lambda t} \end{aligned}$$

where  $C_1$  depends on the initial data.

Result for the linear instability with mixing of modes : Let  $T_1(t)$ ,  $\vec{u}_1(t)$  be the solution of the linearized Euler system. There exists a constant  $C_s$  depending only on the characteristics of the system, that is of  $k_0$  and |g|, such that

$$||\rho_0^{\frac{1}{2}}T_1(t)||_{H^s} + ||\rho_0^{\frac{1}{2}}\vec{u}_1(t)||_{H^s} \le C_s(1+t)^s e^{\Lambda t} (||\rho_0^{\frac{1}{2}}T_1(0)||_{H^s} + ||\rho_0^{\frac{1}{2}}\vec{u}_1(0)||_{H^s})$$

## 6 Weakly non linear approximate solution of order N

We use the Grenier construction, called the weakly non linear expansion (or solution in the physics solutions) :  $(T^N, u^N, v^N, q^N) = (1, 0, 0, 0) + \sum_{j=1}^N \delta^j U_j$ with  $U_j(x, z, 0) = 0$  for  $j \ge 2$  and  $U_1$  solution of the linearized system (with k, a growth rate  $\gamma(k) \in (\frac{\Lambda}{2}, \Lambda)$  and  $\hat{u}$ ). Obviously  $(Emod)(U_j) = S(U_1, ..., U_{j-1})$ . Goal : identify a time T, two constants  $C_0$  and  $M_0$  such that  $||U_j||_{L^2} \le M_0(C_0)^j e^{j\gamma(k)t}, t < T$  $\Rightarrow$  the sequence  $\sum_{j=1}^N \delta^j U_j$  converges normally in  $L^2$  for  $t < \frac{1}{\gamma(k)} \ln \frac{1}{C_0 \delta}$ . Counting terms in the left hand side and estimates :

This result relies on a recurrence, in which one must control the behavior in N of the left hand term of the equation on  $U_N$ . Time derivative  $\partial_t \vec{S} : N$ 

Number of terms (using the **quadratic** nonlinearity) : NBehavior of  $K_N(s) : N^2 e^{N\gamma(k)s}$ 

 $\Rightarrow \int_0^t \sqrt{K_N(s)e^{-\Lambda s}} ds \le \frac{2N}{N\gamma(k)-\Lambda} e^{\frac{N\gamma(k)-\Lambda}{2}t}.$ 

and for  $\gamma(k) > \frac{\Lambda}{2}$  and  $N \ge 2$  bounded independently on N. Rk : not possible to do so when **cubic** nonlinearity.

#### 7 Nonlinear solution

Consider a solution  $(T, \vec{u}, q)$  of the Euler system (2) such that  $(T, \vec{u}, q)(0) = (1, 0, 0) + \delta U_1(0).$ Denote by  $(\tilde{T}, \vec{v}, \tilde{q})$  the difference  $(T, \vec{u}, q) - U^N$ . Moser estimates and usual energy equalities  $\Rightarrow$ control on  $\rho_0^{\frac{1}{2}} \vec{v}, \rho_0^{-\frac{1}{2}} T, \rho_0^{-\frac{1}{2}} q$ Use the nonlinear equation and write as a source term  $k_0 u = (k_0 \rho_0^{-\frac{1}{2}}) \rho_0^{\frac{1}{2}} u$ : one gets estimates on  $\tilde{T}, \vec{v}$  and  $\rho_0^{-1} \tilde{q}$ . End of the proof :  $\vec{u} = \vec{v} + \delta \vec{u}_1 + \delta^2 \sum_{j=2}^N \delta^{j-2} U^j$ First term controlled by  $\epsilon$ Last term controlled by  $\delta^2 C$ Leading term is the second one : can be of order  $\frac{1}{2}$ .